

1 **Tungsten mineralization during evolution of a magmatic–hydrothermal system:**
2 **mineralogical evidence from the Xihuashan rare-metal granite in South China**

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4 ***Revision 2***

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ABSTRACT

19 Tungsten deposits are usually associated with granitic intrusions that record a long
20 and complex evolution of the magmatic–hydrothermal system. However, the genetic
21 link between magmatic–hydrothermal evolution and tungsten mineralization remains
22 unclear. The Xihuashan tungsten deposit in South China, an important vein-type
23 wolframite deposit, is closely associated with greisen and multiphase intrusive activity
24 that produced biotite granite, two-mica granite, and muscovite granite. From the biotite
25 granite to the two-mica granite to the muscovite granite, micas vary from siderophyllite
26 to lithian siderophyllite, with decreasing K/Rb and Nb/Ta ratios and increasing Rb and
27 Cs contents. The zoned micas in the muscovite granite and greisen display
28 fluorine-depleted rims, reflecting subsolidus replacement by external aqueous fluids.
29 The presence of siderite indicates a Fe, Mn and CO₂-rich fluid under reducing
30 conditions. The micas in the greisen have higher F contents and lower Fe³⁺/Fe²⁺ ratios
31 than those in the muscovite granite, suggesting that the fluids contributing to greisen
32 formation had a relatively high fluorine content and were reduced. The increase of CO₂
33 in the fluid enhanced its ability to unlock W from melts/rocks into fluids. The reducing
34 environment also facilitated the tungsten mineralization. During greisenization, the pH
35 value of the fluid increased, which destabilized the polymeric tungstates to form WO₄²⁻.
36 The mixture of W-rich solution and Fe, Mn-rich external fluid eventually precipitated as
37 vein-type wolframite in favorable locations. An empirical equation ($\text{Li}_2\text{O} = 0.0748 \times \text{F}^2$
38 $+ 0.0893 \times \text{F}$) was introduced for estimating the Li₂O contents of hydrothermal micas

39 using the F contents determined by EPMA.

40 **Key words:** rare-metal granite; tungsten mineralization; magmatic–hydrothermal

41 evolution; mica; siderite; South China

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INTRODUCTION

43 Most tungsten deposits are closely associated with granitic intrusions, and the styles
44 of mineralization may be diverse, including quartz veins, stockworks, breccia, skarn,
45 greisen, pegmatite, and porphyry. Wolframite-bearing quartz veins, regarded as an
46 important type of hydrothermal deposit, commonly occur in differentiated granitic
47 plutons that show varying degrees of hydrothermal alteration (e.g., Gleeson et al. 2001;
48 [Schaltegger et al. 2005](#); [Vigneresse 2006](#)). The relationship between magmatic–
49 hydrothermal activities and tungsten mineralization is therefore a critical issue in
50 understanding the formation of tungsten deposits. For individual deposits, the role of the
51 intrusion as the main source of fluids and metals is usually unclear due to the multiple
52 overprint/reopening of the veins and to the lack of geochronological data constraining
53 the timing of magmatic and hydrothermal events. It has been proposed that the
54 ore-forming elements may be concentrated during multiple stages of magmatic activity
55 and then extracted by magmatic fluids (e.g., [Che et al. 2013](#); [Harlaux et al. 2018a](#)), or be
56 leached from granitoids and/or metamorphic rocks and transported by external
57 (metamorphic or meteoric) fluids (e.g., [Linnen and Williams–Jones et al. 1995](#); [Vindel
58 et al. 1995](#); [Vallance et al. 2001](#)). Most previous studies on the relationship between
59 magmatic–hydrothermal episodes and tungsten mineralization have been based on the
60 geochronology of deposits and their host granites (e.g., [Wang et al. 2011](#); [Li et al. 2013](#)),
61 experiments on partition coefficients between minerals and fluids/melts (e.g., [Linnen
62 and Cuney 2005](#); [Che et al. 2013](#)), analyses of fluid inclusions (e.g., [Wei et al. 2012](#)),
63 and analyses of mineral geochemistry and stable isotopes ([Legros et al. 2019](#)). Although

64 recent studies have successfully dated wolframite from tungsten deposits (e.g., Harlaux
65 et al. 2018b; Deng et al. 2019), limitations still exist, owing to the lack of suitable
66 techniques for precisely dating both hydrothermal deposits and associated granitoids,
67 the intrinsic differences between modeled and natural granitic systems, and
68 uncertainties on whether fluid inclusions are representative of ore-forming fluids.

69 The textures and chemical compositions of ore-bearing oxide minerals,
70 rock-forming minerals, and even accessory minerals may be used to track magmatic–
71 hydrothermal processes in granitic systems and related metal mineralization (e.g.,
72 Roda–Robles et al. 2007; Van Lichtervelde et al. 2008; Yang et al. 2013; Li et al. 2015;
73 Yin et al. 2019). Micas are ubiquitous minerals in peraluminous granitic systems, and
74 these primary magmatic or secondary minerals (due to fluid interactions) have the
75 potential to record the processes of fractional crystallization and hydrothermal evolution.
76 As the major rock-forming mineral that hosts tungsten (e.g., 1.9–198 ppm in micas from
77 peraluminous granites; Simons et al. 2017; El Korh et al. 2020), micas provide an
78 effective proxy for tracing the progressive enrichment and ore-forming processes during
79 magmatic–hydrothermal evolution (Legros et al. 2016, 2018; Yin et al. 2019). In
80 addition to CO₂-rich fluid inclusions, carbonates (including synchysite, siderite, and
81 calcite) may occur frequently in rocks associated with tungsten deposits, suggesting that
82 the fluids involved in W-mineralization are enriched in CO₂ (Castorina and Masi 2008;
83 Wei et al. 2012).

84 The widespread rare-metal deposits in South China are closely associated with
85 Mesozoic granitic rocks, and an example is the Nanling W–Sn polymetallic

86 mineralization region (Fig. 1a) (Yin et al. 1995; Wang et al. 2003; Li et al. 2015). The
87 Xihuashan tungsten deposit located in the Jiangxi Province is a typical vein-type
88 wolframite deposit with minor Sn–Be mineralization (Nie and Wang 2007), and it is one
89 of the major tungsten deposits in the world (Sinclair et al. 2014). The Xihuashan deposit
90 is spatially and temporally associated with Jurassic granitic intrusions (Wang et al. 2011;
91 Li et al. 2013; Yang et al. 2018), and it shows genetic links with the magmatic and
92 hydrothermal evolution of the granitic intrusions. This deposit therefore provides an
93 opportunity to investigate the mechanism of W enrichment and mineralization in
94 rare-metal granites during the evolution of the magmatic–hydrothermal system.
95 Previous studies on the Xihuashan deposit were focused mainly on the geochronology
96 of the deposits and related granites (Wang et al. 2011; Guo et al. 2012; Hu et al. 2012;
97 Li et al. 2013), fluid inclusions and wolframite formation (Wei et al. 2012; Huang et al.
98 2013), and magmatic fractionation or hydrothermal alteration of the granites (Le Bel et
99 al. 1983; Maruéjol et al. 1990; Wang et al. 2003; Yang et al. 2013). This paper presents
100 the results of whole-rock major and trace element analyses of the Xihuashan rare-metal
101 granite as well as the chemical compositions of the micas and siderite, to elucidate the
102 relationships between the tungsten mineralization and the magmatic–hydrothermal
103 evolution of highly fractionated granitic melts, and to reveal the relative contributions of
104 magmatic and hydrothermal processes to the formation of tungsten ore deposits. The
105 results show the critical role of greisenizing fluids and magma fractionation in the
106 tungsten mineralization, and provide new insights into the metallogensis of tungsten in
107 rare-metal granites.

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GEOLOGICAL SETTING AND PETROGRAPHY

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The South China Block (SCB) comprises the Yangtze Block in the northwest and the Cathaysia Block in the southeast (Fig. 1a). The two blocks were amalgamated during the early Neoproterozoic (e.g., Li et al. 2009; Zhao et al. 2011). During the Phanerozoic, the Cathaysia Block was subjected to intensive and widespread multiphase orogenesis and magmatism (Li et al. 2012). The igneous rocks of Jurassic and Cretaceous age are predominantly granites and rhyolites with subordinate mafic and rare intermediate lithologies (Zhou et al. 2006; Li et al. 2012). The Jurassic rocks occur over large parts of the Cathaysia Block, while the Cretaceous rocks cover the coastal area of SE China and the middle to lower reaches of the Yangtze River in the Yangtze Block (Fig. 1a). The majority of highly evolved and peraluminous granitic rocks were formed during the Jurassic with a few during the Cretaceous (Legros et al. 2020). These rocks in the Nanling area are associated with W–Sn–Mo–Bi–Be–Nb–Ta mineralization, and the area is well known as the Nanling tungsten–tin polymetallic mineralization region (NPMR, Fig. 1a) (e.g., Wang et al. 1982; Pei and Hong 1995).

The Xihuashan mining district is located in the southwestern part of the Jiangxi Province, South China (Fig. 1). It contains more than 700 economically viable ore veins, including W–Sn–Be mineralization (Nie and Wang 2007). There are six main tungsten mines in the district, namely Xihuashan, Dangping, Shenlongkou, Luokeng, Xialuogushan, and Niuzishi (Fig. 1b). The Xihuashan Mine hosts the largest vein-type wolframite deposit in the district, with the total estimate of WO₃ being 81,300 tons and

130 an average ore grade of 1.08 % WO₃ (Wei et al. 2012). The deposit is spatially and
131 temporally related to a multiphase granitic intrusion that crops out over an area of about
132 20 km² (Guo et al. 2012). Previous geochronological studies indicated a close temporal
133 relationship between the granitic magmatism and the W mineralization. For example,
134 the crystallization ages of the major intrusions are 156–161 Ma (Guo et al. 2012; Li et al.
135 2013; Tang et al. 2020), and the Re–Os ages of molybdenites in the ore are identical
136 (157 ± 3 Ma; Wang et al. 2011). The Xihuashan granitic complex was emplaced in
137 upper Cambrian low-grade metamorphosed flysch-like sedimentary rocks, which
138 underwent contact metamorphism (Le Bel et al. 1984). Wei et al. (2012) suggested the
139 ore-forming fluids to be a mixture of magmatic and hydrothermal fluids derived from
140 the granitic melts and meteoric water. With the knowledge obtained from geological
141 mapping and exploratory drilling, the Xihuashan granitic complex may be divided into
142 four successive intrusions (G1–G4; Fig. 1b) (Maruéjol et al. 1990; Wang et al. 2003).
143 The G1 (coarse-grained biotite granite) occurs as the outermost part of the complex. The
144 G2 (medium-grained biotite granite) hosts tungsten mineralization at the Xihuashan
145 Mine. The G3 (medium- to fine-grained biotite granite) hosts tungsten mineralization at
146 the Dangping Mine, whereas the G4 (fine-grained biotite granite) occurs as a barren
147 border of G3.

148 The collected samples of G2 (medium-grained biotite granite) in the Xihuashan
149 Mine include biotite granite, two-mica (lithian siderophyllite–muscovite) granite,
150 muscovite granite, and greisen (Fig. 2a–d). The biotite granite crops out over an area of
151 about 5.2 km². The greisen, two-mica granite and muscovite granite are only exposed in

152 the sub-surface and were investigated using drilled samples and exposure in the mine
153 tunnels. Liu (2005) and Huang et al. (2006) showed that the greisen was usually
154 associated with tungsten ore veins cutting the fine-grained two-mica granite, the
155 muscovite granite, or the country rocks. There is a transition between the two-mica
156 granite and the muscovite granite, both of which intruded the biotite granite (Liu 2005;
157 Huang et al. 2006). The main accessory minerals of the Xihuashan intrusive rocks are
158 fluorite, apatite, zircon, thorite, xenotime, and rutile (Supplementary Table 1).

159 The *biotite granite* consists of quartz (32–38%), K-feldspar (19–20%), plagioclase
160 (32–34%), and siderophyllite (8–13%) (Fig. 2a). Plagioclase is subhedral or euhedral
161 and shows compositional zoning. The siderophyllite grains are coarse and up to 2 mm in
162 diameter. The plagioclase and siderophyllite always host inclusions of ilmenite and
163 apatite. Rutile is common as a secondary mineral associated with chloritization of the
164 micas. The fluorite grains are anhedral and contain high yttrium contents ($Y_2O_3 = 8\text{--}15$
165 wt%). The accessory minerals enriched in REEs include apatite, monazite, allanite, and
166 zircon.

167 The *two-mica (lithian siderophyllite–muscovite) granite* is composed mainly of
168 quartz (30–35%), K-feldspar (14–15%), plagioclase (43–49%), and micas (6–8%) (Fig.
169 2b). The micas are dominantly lithian siderophyllite (~6%) with lesser amounts of
170 muscovite (~2%). The primary mica crystals are coarse and up to 2 mm in size, while
171 most secondary micas are fine-grained sericite within large plagioclase and K-feldspar
172 phenocrysts. The lithian siderophyllite always contains inclusions of ilmenite and
173 apatite, and the rims always show porosity due to dissolution. Xenotime is the principal

174 phosphate mineral in the two-mica granite, and other common accessory minerals are
175 zircon, garnet, and fluorite.

176 The *muscovite granite* consists of quartz (39–45%), K-feldspar (19–22%),
177 plagioclase (32–35%), and muscovite (~4%) (Fig. 2c). The accessory minerals include
178 spessartine, zircon, fluorite, xenotime, and siderite. The siderite crystals are anhedral
179 and elongate, and they occur mostly along the cleavages of the micas.

180 The *greisen* is composed of quartz (49–55%), lithian siderophyllite (30–35%), and
181 muscovite (9–11%) (Fig. 2d). The principal accessory minerals include zircon, xenotime,
182 fluorite, siderite, niobotantalates, and wolframite. Xenotime and zircon are the
183 predominant REE-rich minerals, while niobotantalates and wolframite are the
184 tungsten-bearing minerals.

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ANALYTICAL METHODS

187 The whole rock geochemical analyses were carried out at the State Key Laboratory
188 of Isotope Geochemistry (SKLaBIG), Guangzhou Institute of Geochemistry (GIG),
189 Chinese Academy of Sciences (CAS). To eliminate potential contamination, samples
190 were cut into slabs and the central parts (> 200 g) were selected for whole-rock analyses.
191 The rocks were crushed into pieces with size of < 0.5 cm in diameter, cleaned with
192 deionized water in ultrasonic bath, and then pulverized in a corundum mill. Major
193 element oxides were determined using a Rigaku RIX 2000 X-ray fluorescence (XRF)
194 spectrometer. Samples were prepared as glass discs by a Rigaku desktop fusion machine,
195 mixing 0.50 g of rock powder (dried at 110 °C) with 4.0 g of Li₂B₄O₇ for 15 min at

196 1100 °C in 95% Pt - 5% Au crucibles. Calibration lines used for quantification were
197 produced by bivariate regression of data from 36 reference materials. Matrix corrections
198 were also incorporated into calibration using the empirical Traill-Lachance procedure.
199 For whole rock analyses, total Fe is reported as Fe₂O₃ ([Supplementary Table 2](#)). The
200 analytical uncertainty was between 1 % and 5 %. The additional quantitative analysis of
201 F was conducted by ALS Chemex (Guangzhou) Co Ltd., China, using the methods of
202 KOH fusion and ion selective electrode, or Na₂O₂ fusion, citric acid leach and ion
203 selective electrode. Analytical precision of F was lower than 10%.

204 Trace elements were determined using inductively coupled plasma-mass
205 spectrometry (ICP-MS). The rock powder was digested in a mixture of HF and HNO₃ in
206 high-pressure Teflon bombs. An internal standard solution containing single element Rh
207 was used to monitor signal drift. The detailed procedure was described in [Li et al.](#)
208 ([2006](#)). The US Geological Survey and Chinese National Standards, including SARM-4,
209 W-2, BHVO-2, AGV-2, GSR-1, GSR-2 and GSR-3, were used for calibration.
210 Analytical precision of REE and other incompatible elements was in the range of 1 to
211 5 %.

212 The siderites were examined using HORIBA Xplora laser Raman
213 microspectroscopy at the Key Laboratory of Mineralogy and Metallogeny (KLMM),
214 GIG-CAS. An Ar ion laser operated at 44 mW was used to produce an excitation line
215 with wavelength of 532 nm. The scanning range of spectra was set between 100 and
216 4000 cm⁻¹ with an accumulation time of 10 s for each scan. The spectral resolution was
217 0.65 cm⁻¹. The Raman shift of a monocrystalline silicon piece was measured to be 520.7

218 cm⁻¹ before analyses.

219 Back-scattered-electron (BSE) images were obtained by a Zeiss Supra 55
220 field-emission scanning electron microscope (FE-SEM) at the SKLaBIG-GIG-CAS.
221 The *in-situ* mineral composition analyses were performed using the JEOL JXA8100
222 electron probe micro-analyzer (EPMA) at the SKLaBIG-GIG-CAS, with the operating
223 conditions of 15 kV accelerating voltage, 20 nA beam current and 1 μm beam size
224 (siderite) or 5 μm beam size (micas). A variable peak counting time of 7–40 s was
225 designed for the intensity of the characteristic X-ray lines and desired precision
226 ([Supplementary Table 3](#)). The raw data were corrected using ZAF correction. For micas,
227 the detection limits for most elements are lower than 251 ppm, and those for F and Rb
228 are 568 ppm and 631 ppm, respectively ([Supplementary Table 3](#)). Relative uncertainties
229 were estimated to be ca. 7% for F and 10–15 % (1σ) for Cl. Chemical formulae of micas
230 were calculated from mineral analyses based on 24 anions (O, F, OH), and Fe³⁺ was
231 estimated following the method by [Lin and Peng \(1994\)](#) described in [Supplementary](#)
232 [Table 4](#). The CO₂ content in siderite was calculated by stoichiometry.

233 *In-situ* trace element concentrations were measured in mica and siderite grains
234 using an Agilent 7500a quadrupole ICP-MS coupled with a RESolution M-50 laser
235 ablation system at the KLMM-GIG-CAS. The operating conditions included a spot size
236 of 42 μm, repetition rate of 5 Hz, maximum energy of 90 mJ and laser fluence of 7
237 J/cm². Argon was used as a make-up gas and mixed with helium as the carrier gas via a
238 cyclone mixer prior to entering the ICP torch. The National Institute of Standards and
239 Technology (NIST) 610 glass and 612 glass were used as external standards for quality

240 control ([Jochum et al. 2011](#)). Every analysis of six samples was followed by NIST 610
241 and 612 standards for time-dependent calibration of sensitivity drift. The internal
242 standards are ^{27}Al and ^{57}Fe for mica and siderite, respectively. The off-line data
243 processing was performed using the commercial software ICPMSDataCal 6.7 ([Liu et al.](#)
244 [2008](#)). The dwell time of each spot is 70 s of analysis time (30 s background, 40 s
245 ablation). Limits of detection (LOD) were calculated using the 3σ criterion detailed by
246 [Longerich et al. \(1996\)](#). The relative standard deviation (RSD) varies from 1% to 10%.
247 More details on the standards are reported in [Supplementary Table 5](#).

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WHOLE-ROCK COMPOSITIONS

250 Eight samples (two of each rock type) were analyzed for their whole rock major
251 and trace element compositions ([Supplementary Table 2](#)). The granite samples are all
252 peraluminous ($A/\text{CNK} = 1.02\text{--}1.10$; [Supplementary Table 2](#)), with high SiO_2 (72.50–
253 82.54 wt%), F (0.08–0.67 wt%), and alkali contents ($\text{Na}_2\text{O} + \text{K}_2\text{O} = 6.36\text{--}8.78$ wt%) but
254 low contents of MgO (0.02–0.38 wt%), Fe_2O_3 (0.44–2.02 wt%), and TiO_2 (<0.18 wt%).
255 From the biotite granite to the two-mica granite to the muscovite granite, the TiO_2 , MgO,
256 P_2O_5 , CaO, and F contents decrease gradually with decreasing Fe_2O_3 contents ([Fig. 3](#)),
257 consistent with previous data reported by [Guo et al. \(2012\)](#). The two greisen samples
258 display higher contents of SiO_2 (79.81–81.42 wt%) but lower contents of Al_2O_3 (8.98–
259 9.74 wt%), MgO (0.06–0.09 wt%), MnO (0.15–0.19 wt%), TiO_2 (0.02–0.03 wt%), CaO
260 (0.21–1.64 wt%), and alkalis ($\text{Na}_2\text{O} + \text{K}_2\text{O} = 3.37\text{--}3.38$ wt%) than the granite samples
261 ([Supplementary Table 2; Fig. 3](#)). The greisen samples have high Fe_2O_3 (3.28–3.36 wt%)

262 and F (1.05–1.22 wt%) contents (Fig. 3), due to abundant lithian siderophyllite,
263 wolframite, and fluorite.

264 The granite and greisen samples have a narrow range of REE concentrations
265 (80.4–129 ppm; Supplementary Table 2) and similar chondrite-normalized REE patterns
266 with strong negative Eu anomalies ($\text{Eu}/\text{Eu}^* = 0.01\text{--}0.31$) and a weak lanthanide tetrad
267 effect, consistent with previous data reported by Guo et al. (2012) (Fig. 4a). The
268 $[\text{La}/\text{Yb}]_N$ and Eu/Eu^* values decrease following the sequence of biotite granite,
269 two-mica granite, muscovite granite, and greisen. The muscovite granite and greisen are
270 depleted in LREEs and Eu, but enriched in HREEs, and this differs from the typical
271 “wing-shaped” REE patterns for highly evolved granites (Irber 1999; Li et al. 2015). In
272 the multi-element diagrams normalized to the mean composition of upper continental
273 crust, all the samples are enriched in Cs (11.1–73.5 ppm), Rb (312–1020 ppm), W
274 (4.82–65.3 ppm), Ta (3.35–27.2 ppm), Sn (5.65–198 ppm), and Li (37.8–987 ppm), but
275 depleted in Ba (3.31–266 ppm), Sr (1.31–90.6 ppm), P (9.29–204 ppm), and Ti (12.1–
276 1069 ppm) (Fig. 4b). Overall, the muscovite granite and greisen have much higher Li,
277 Cs, Rb, and Sn abundances but lower Ba, Sr, and P than the biotite granite and two-mica
278 granite.

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MINERAL COMPOSITIONS

281 **Mica**

282 Most micas in the biotite granite and two-mica granite are homogeneous in their
283 chemical compositions, but those in the muscovite granite and greisen show

284 compositional zoning with distinct core and rim on BSE images or have a porous rim
285 due to dissolution (Fig. 5). A total of 267 EPMA analyses were performed on eight
286 samples, along with 76 LA-ICP-MS analyses on the same spots previously analyzed by
287 EPMA. Altogether, these analyses cover all the rock types in the pluton and different
288 areas of compositionally zoned mica grains (Supplementary Table 6).

289 The micas in the Xihuashan pluton show a wide range of elemental compositions.
290 Siderophyllite in the biotite granite has the highest contents of FeO^{Tot} (21.70–24.57 wt%,
291 2.757–3.253 apfu), TiO_2 (0.53–3.20 wt%, 0.061–0.373 apfu), and MgO (3.96–5.12 wt%,
292 0.925–1.165 apfu) of all the micas, but the lowest contents of SiO_2 (34.44–37.00 wt%,
293 5.469–5.667 apfu) and Al_2O_3 (16.34–20.77 wt%, 2.333–2.531 apfu) (Fig. 6;
294 Supplementary Table 6). In the two-mica granite, the lithian siderophyllite has higher
295 contents of FeO^{Tot} (19.40–23.17 wt%, 2.439–3.002 apfu) and TiO_2 (0.03–1.57 wt%,
296 0.004–0.183 apfu) than the muscovite ($\text{FeO}^{\text{Tot}} = 5.00\text{--}8.85$ wt%, 0.576–1.051 apfu;
297 $\text{TiO}_2 = 0.01\text{--}0.40$ wt%, 0.001–0.042 apfu). Micas in the muscovite granite and greisen
298 have lower contents of FeO^{Tot} (3.43–19.97 wt%, 0.388–2.504 apfu), TiO_2 (below the
299 detection limit to 0.51 wt%, bdl.–0.058 apfu), and MgO (0.04–0.70 wt%, 0.008–0.147
300 apfu), but higher contents of SiO_2 (37.77–47.99 wt%, 5.683–6.529 apfu) than those in
301 the biotite granite and two-mica granite (Fig. 6). Micas in the biotite granite, two-mica
302 granite, and muscovite granite have similar F contents, all of which are lower than the
303 value of the cores of greisen micas (Fig. 6). Micas in all samples have extremely low Cl
304 contents (mostly less than 0.06 wt%), which decrease roughly from siderophyllite,
305 lithian siderophyllite, and muscovite along with a negative correlation with SiO_2

306 contents (Fig. 6f).

307 Micas in the biotite granite have much higher V, Sc, Cr, Co, and Ni contents than
308 the micas in other rocks (Fig. 7a; Supplementary Table 6). Overall, the micas in the
309 Xihuashan pluton have high contents of Rb, Cs, Nb, and Ta (typically Rb > 1000 ppm,
310 Cs > 200 ppm, Nb > 20 ppm, Ta > 10 ppm) but the contents of most REEs are below
311 detection limits (Supplementary Table 6). The lithian siderophyllites in the two-mica
312 granite and the cores of micas in the muscovite granite and greisen have higher Rb and
313 Cs contents than siderophyllites in the biotite granite (Fig. 7b, c), while the Nb and Ta
314 contents of micas increase from the biotite granite to the two-mica granite, but decrease
315 slightly in the greisen (Fig. 7d, e). Micas in the muscovite granite and greisen have
316 lower Nb/Ta ratios (0.84–5.63) than siderophyllites in the biotite granite and lithian
317 siderophyllites in the two-mica granite (Nb/Ta = 1.77–17.15) (Fig. 7f). The Sn and W
318 contents of micas in the analyzed samples increase gradually from the biotite granite to
319 the two-mica granite and to the cores of micas in the muscovite granite.

320 The micas in the muscovite granite and greisen generally exhibit core–rim
321 zonation on BSE images (Fig. 5). The cores of zoned micas in the muscovite granite and
322 greisen commonly have lower SiO₂ (37.77–42.47 wt%, 5.683–6.333 apfu) and Al₂O₃
323 (23.33–26.91 wt%, 1.667–2.317 apfu) contents but higher FeO^{Tot} (11.32–19.97 wt%,
324 1.356–2.504 apfu), Li₂O (0.26–1.74 wt%, 0.141–1.023 apfu), F (0.73–4.26 wt%, 0.342–
325 1.981 apfu), Rb (4753–5860 ppm), Cs (377–1691 ppm), Ta (15.8–171 ppm), and Nb
326 (33.9–195 ppm) contents than the rims (Figs 5, 6e, 7b–e; Supplementary Table 6). The
327 cores of zoned micas in the muscovite granite and greisen have lower Nb/Ta ratios

328 (0.97–1.55 and 0.84–2.14, respectively) than the rims (1.29–5.63 and 2.01–3.40) (Fig.
329 7f). In addition, the mica cores in the muscovite granite have higher W (70–84 ppm) and
330 Sn (694–848 ppm) contents than the rims (14.9–63 ppm and 265–694 ppm,
331 respectively). The zoned micas in the greisen samples exhibit an increasing trend of W
332 and Sn from core to rim, with values higher than the rims of zoned micas in the
333 muscovite granite (Fig. 7g, h).

334

335 **Siderite**

336 All the siderites in the muscovite granite and greisen samples occur along the
337 cleavages or the crystal borders of zoned micas (Fig. 5), which were most likely formed
338 as secondary minerals due to the presence of CO₂-rich hydrothermal fluids (Buckley
339 and Woolley 1990). The siderites were identified using Raman spectroscopy and EPMA.
340 The Raman spectrum of the investigated mineral shows a strong intensity band at
341 1081.73 cm⁻¹ and a low intensity band at 276.97 cm⁻¹ (Fig. 8), corresponding to the
342 symmetric stretching and external vibration of the CO₃ group, respectively, which are
343 consistent with the characterization of carbonates reported by Buzgar and Apopei
344 (2009). Minor shifts in position may be due to the effect of natural impurities in the
345 studied samples. The siderites have high contents of FeO (30.14–56.01 wt%), MnO
346 (3.47–20.39 wt%), and CO₂ (35.29–37.60 wt%), but low contents of MgO (<0.22 wt%)
347 and alkalis (Na₂O + K₂O < 1 wt%) (Supplementary Table 7). The siderites in the
348 muscovite granite have higher REE contents (ΣREE = 62–109 ppm) than those in the
349 greisen (ΣREE = 6.92–34.4 ppm; Supplementary Table 8). They show chondrite

350 normalized LREE-depleted and HREE-enriched pattern with very low values of
351 $[La/Yb]_N$ (0.01–0.05) (Fig. 8c; Supplementary Table 8) and strongly negative Eu
352 anomalies ($Eu/Eu^* < 0.32$; Fig. 8c; Supplementary Table 8). On the multi-element
353 diagram normalized to average upper continental crust, the siderites are enriched in Li,
354 Rb, Cs, Ta, W, and Sn, but depleted in Ba, Zr, and Ti, similar to the whole rock
355 composition (Fig. 8d).

356

357

DISCUSSION

358 Estimation of lithium contents in micas

359 The Li_2O content of mica may be measured directly by LA–ICP–MS or estimated
360 by EPMA from empirical equations based on the crystallochemical relationship between
361 Li and other major elements in the mica. Considering the different beam sizes of the
362 LA–ICP–MS (e.g., 42 μm) and EPMA (e.g., 1 μm) methods, the Li_2O content of mica
363 estimated from the EPMA results using empirical equations may provide a better
364 constraint on the compositional variation at the micrometer scale.

365 The micas in the Xihuashan pluton display two trends on the diagram of SiO_2
366 versus Li_2O with contents measured by LA–ICP–MS (Fig. 9a). Siderophyllites in the
367 biotite granite and lithian siderophyllites in the two-mica granite have relatively low
368 SiO_2 contents that show a positive correlation with Li_2O (Fig. 9a), which is consistent
369 with the regression function between Li_2O and SiO_2 (Fig. 9a) for trioctahedral micas
370 (Tischendorf et al. 1997):

$$371 \quad Li_2O = 0.289 \times SiO_2 - 9.658 \quad (1)$$

372 In contrast, muscovites in the two-mica granite and micas in the muscovite granite and
373 greisen have relatively higher SiO₂ contents that show a negative correlation with Li₂O
374 contents (Fig. 9a). In particular, muscovites in the two-mica granite and the cores of
375 micas in the muscovite granite show good correlation with the regression function
376 between Li₂O and F (Fig. 9b) for dioctahedral micas (Tischendorf et al. 1997):

$$377 \quad Li_2O = 0.3935 \times F^{1.326} \quad (2)$$

378 while micas in the greisen samples clearly deviate from this function (Fig. 9b). The Li
379 contents of hydrothermal micas (e.g., micas in the greisen) appear to be overestimated
380 using Eq. (2) for dioctahedral micas (Fig. 9b). While the equations in Tischendorf et al.
381 (1997) were designed for magmatic systems, additional adjustment is required for
382 application to hydrothermal systems (Legros et al. 2016). In this paper, a new empirical
383 equation for the Li₂O and F contents of hydrothermal micas is introduced (Fig. 9c):

$$384 \quad Li_2O = 0.0748 \times F^2 + 0.0893 \times F \quad (3)$$

385 This equation may be used for estimating the Li₂O content of hydrothermal mica from
386 the F content determined by EPMA. The Li contents calculated with this equation are in
387 agreement with the values measured by LA-ICP-MS ($R^2 = 0.82$) (Fig. 9d).

388 The Li₂O contents determined using EPMA indicate that the biotites in the biotite
389 granite and two-mica granite are siderophyllite and protolithionite (lithian
390 siderophyllite), respectively, which is consistent with the trend of magmatic evolution
391 (Fig. 10). Lithian siderophyllites in the greisen have higher Li₂O contents (1.57–1.74
392 wt%) than those in the two-mica granite (0.88–1.33 wt%). They are protolithionites
393 following the classification proposed by Tischendorf et al. (1997) (Fig. 10a).

394 Alternatively, they belong to the transitional series between dioctahedral and
395 trioctahedral micas based on Foster (1960) (Fig. 10b), also termed the ferrous aluminum
396 micas in Monier and Robert (1986). Muscovites in the two-mica granite, muscovite
397 granite, and greisen all belong to the transitional series, but in their octahedral sites they
398 have lower contents of Li_2O and divalent metal cations (Fe^{2+} , Mn^{2+} , Mg^{2+}) than the
399 lithian siderophyllites in the greisen (Fig. 10b).

400

401 **Variation of halogens in the melts and fluids**

402 The micas display a wide range of F contents (0.41–4.26 wt%). Overall, the
403 siderophyllites in the biotite granite have the lowest F contents (0.84–1.95 wt%), while
404 micas in the greisen have the highest (up to 4.26 wt%). The cores of zoned micas in the
405 muscovite granite and greisen typically have higher F contents (0.73–4.26 wt%) than
406 the rims (0.41–3.46 wt%) (Supplementary Table 6). In addition, the micas have much
407 lower Cl contents (up to 0.06 wt%) than F (0.41–4.26 wt%; Supplementary Table 6).
408 Indeed, the substitution of Cl into the OH site of mica is much less favorable than the
409 substitution of F (Munoz 1992), due to the significantly larger ionic radius of Cl^- (1.81
410 Å; cf. $\text{F}^- = 1.31$ Å and $\text{OH} = 1.38$ Å).

411 The exchange of F, Cl, and OH between micas and associated fluids may be used
412 to obtain the halogen fugacity of the fluids (Munoz 1984) as well as the halogen
413 concentrations in the melt (Zhang et al. 2012). Since the measured Cl contents of the
414 Xihuashan micas are extremely low (mostly less than 0.06 wt%, with large errors of
415 10%–15%), the exchange of Cl and OH between micas and associated fluids is not

416 further discussed.

417 The equations used for estimating the F fugacity of a fluid [$\log(f_{\text{H}_2\text{O}}/f_{\text{HF}})$] and
418 the F concentration in a melt (W_F^m) are listed as follows (Munoz 1984; Zhang et al.
419 2012):

$$420 \quad \log\left(\frac{f_{\text{H}_2\text{O}}}{f_{\text{HF}}}\right) = \frac{1000}{T} \times (2.37 + 1.1X_{\text{Mg}}) + 0.43 - \log\left(\frac{X_{\text{F}}}{X_{\text{OH}}}\right)^{\text{mica}} \quad (4)$$

421

$$422 \quad W_F^m = \frac{1}{1 + \frac{f_{\text{H}_2\text{O}}}{f_{\text{HF}}}} \times \frac{19}{18.02} \times \frac{10^6}{0.000943 \times T - 0.574} \quad (5)$$

423 where X_{F} and X_{OH} are the mole fractions of F and OH in the hydroxyl site of mica,
424 respectively, X_{Mg} is the mole fraction of Mg in the octahedral site of mica, and T is the
425 crystallization temperature of mica in degrees Kelvin. Due to the difficulties in
426 calculating the crystallization temperatures of micas directly, the homogenization
427 temperatures of fluid inclusions (Mu et al. 1982; Zhang et al. 1982; Wei et al. 2011) and
428 melt inclusions (Chang and Huang 2002) in the Xihuashan tungsten deposit were
429 adopted to estimate the F fugacity of the fluids and the F content of the melt. The melt
430 inclusions in beryl from the wolframite–quartz veins in the Xihuashan tungsten deposit
431 gave a homogenization temperature of 720 °C (Huang et al. 2006), consistent with the
432 mean value (703 ± 28 °C; range of 643 to 890 °C) of the temperature given by the
433 Ti-in-zircon thermometer for the Xihuashan granites (Yang et al. 2018), which can be
434 regarded as the crystallization temperature of the Xihuashan granitic magma. In addition,
435 there were two generations of fluid in the Xihuashan granites, based on stable oxygen
436 isotope data from fluid inclusions (Mu et al. 1982; Zhang et al. 1982; Wei et al. 2011).
437 The primary ore-fluid had a temperature of 370 ± 30 °C (fluid of magmatic origin, $\delta^{18}\text{O}$

438 = 10‰–13‰) and the late fluid had a temperature of 220 ± 10 °C (meteoric water, $\delta^{18}\text{O}$
439 = -8.7‰ to $+3.1\text{‰}$). It is assumed that the primary ore-fluid was related to the
440 formation of the cores of zoned micas in the greisen and that the late fluid was related to
441 the rims of zoned micas in the muscovite granite and greisen, and this will be discussed
442 later. The results of calculations using Eq (4) and Eq (5) are given in [Supplementary](#)
443 [Table 6](#) and shown on [Fig. 11](#). Overall, the $\log(f\text{H}_2\text{O}/f\text{HF})$ values of the fluid decrease
444 from the greisen (4.15–4.65, mean = 4.33) to the biotite granite (3.95–4.42, mean = 4.14)
445 to the two-mica granite (3.38–3.82, mean = 3.65), and to the muscovite granite (3.26–
446 3.91, mean = 3.56) ([Fig. 11a](#)). In general, F is preferentially concentrated in a residual
447 melt when the volatiles of a hydrous magma are exsolved ([Zhu and Sverjensky 1992](#)).
448 The loss of volatiles from a hydrous magma will enhance the F content in the residual
449 melt, producing a lower OH/F value in the melt but a higher OH/F ratio in the
450 associated fluid. This implies that the F content (W_F^m values) of the melt increases
451 gradually from the biotite granite to the two-mica granite, and to the muscovite granite
452 ([Fig. 11c](#)).

453 With regard to the zoned micas in the muscovite granite and greisen, the fluids
454 related to the rims always have higher $\log(f\text{H}_2\text{O}/f\text{HF})$ values than those related to the
455 cores ([Fig. 11a](#)). In addition, the rims of micas in the muscovite granite and greisen all
456 have higher $\text{Fe}^{3+}/\text{Fe}^{2+}$ ratios than the cores ([Fig. 11b](#)). As the $\text{Fe}^{3+}/\text{Fe}^{2+}$ ratio shows a
457 positive correlation with temperature ([Wones and Eugster 1965](#)), the higher $\text{Fe}^{3+}/\text{Fe}^{2+}$
458 ratios of the rims of zoned micas cannot be attributed to lower temperatures relative to
459 the magmatic cores, but an increasing oxygen fugacity and H_2O content.

460

461 **Magmatic evolution of the Xihuashan intrusive rocks**

462 The Xihuashan granites have extremely variable whole-rock and mineral
463 compositions, indicating that the rocks might have undergone varying degrees of
464 fractional crystallization.

465 The micas show a wide range of compositions (Fig. 10) from siderophyllite to
466 lithian siderophyllite, which corresponds to the evolution of the magma from the biotite
467 granite to the two-mica granite. As a peraluminous melt system evolves, Rb and Cs are
468 compatible in biotite while Y, Th, and REEs are incompatible (Bea et al. 1994). The
469 micas in our samples have low Y, Th, and REE contents but extremely high Rb and Cs
470 contents (Supplementary Table 6), suggesting that fractionation of the micas was
471 insignificant during the magmatic evolution. The increase of Rb and Cs contents with
472 decreasing K/Rb ratios from the biotite granite to the two-mica granite (Fig. 12a, b)
473 further excludes the possibility of significant mica fractionation during the magmatic
474 evolution of the Xihuashan intrusive rocks. With regard to the feldspars, Rb and Cs are
475 compatible in K-feldspar but incompatible in plagioclase (Bea et al. 1994), and an
476 increase of Rb and Cs in micas from the biotite granite to the two-mica granite (Fig. 7b,
477 c) could be attributed to a remarkable fractionation of plagioclase. It is also consistent
478 with the fact that our samples are notably enriched in Li, Rb, and Cs compared with
479 upper continental crust (UCC) (Fig. 4b). In addition, xenotime rather than apatite is the
480 principal phosphate mineral in Ca-poor peraluminous granitic melts (Wang et al. 2003).
481 In the Xihuashan pluton, the two-mica granite and muscovite granite samples contain

482 abundant xenotime with apatite being absent or rare, which indicates an essentially
483 Ca-poor peraluminous melt. The Ca-poor melt would have been the product of the
484 fractional crystallization of Ca-rich minerals, including apatite and plagioclase.
485 However, apatite fractionation, if it existed, could not have been primarily responsible
486 for producing the Ca-poor peraluminous melt, because it would have required a large
487 amount of phosphorus to exhaust the Ca in the system. Therefore, the presence of
488 abundant xenotime in the two-mica granite and muscovite granite may provide evidence
489 for plagioclase fractionation during the magmatic evolution, which is consistent with
490 decreasing MgO, TiO₂, Fe₂O₃ and CaO (Fig. 3) and gradually pronounced Eu negative
491 anomalies (Fig. 4a) from biotite granite to two-mica granite to muscovite granite.

492 Collectively, the variations in composition of mica in the Xihuashan pluton are
493 consistent with a trend of magmatic evolution from the biotite granite to the two-mica
494 granite to the muscovite granite. This evolutionary process was controlled mainly by the
495 fractional crystallization of plagioclase.

496

497 **The hydrothermal evolution recorded in zoned micas and siderite**

498 Micas in highly evolved granites and pegmatites always show compositional
499 zoning as a result of the evolution of the magma or sub-solidus fluid interactions
500 (Roda-Robles et al. 2007; Van Lichtenvelde et al. 2008; Li et al. 2015). Volatile
501 components (e.g., H₂O, F, B, and P) and lithophile rare metal elements (e.g., Li, Rb, Cs,
502 Ta, Nb, Sn, and W) are incompatible in major rock-forming minerals during the
503 evolution of peraluminous melt systems. Accordingly, the Li₂O, F, Rb, Cs, Nb, and Ta

504 contents of mica may increase with the progressive evolution of the magma (Roda-
505 Robles et al. 2007), leading to differences in brightness in the zoning textures revealed
506 by the BSE images (Fig. 5). Since fluorine is concentrated in the melt during the
507 differentiation of a granitic melt when there is lack of fractionation of fluorine-enriched
508 minerals such as apatite and mica (Li et al. 2015), the supercritical fluids exsolved from
509 a highly fractionated Sn–W granitic magma are usually enriched in F (Thomas et al.
510 2005). The addition of F into the melt may increase the abundance of non-bridging
511 oxygens (NBOs), which increases the proportion of WO_4^{2-} tetrahedra in silicate melt.
512 Indeed, with our samples, the $\log(f\text{H}_2\text{O}/f\text{HF})$ values decrease and the W_F^m values
513 increase from the biotite granite to the two-mica granite and the muscovite granite (Fig.
514 11), consistent with an evolution of magma without involvement of any notable
515 fractionation of apatite or mica. However, the rims of the zoned micas in the muscovite
516 granite and greisen have lower Li_2O , Rb, Cs, Nb, Ta, and F contents but higher Si and
517 Al contents than the cores (Fig. 9; Supplementary Table 6). Moreover, the rims of the
518 zoned micas belong to the transition series (Fig. 10), similar to the ferrous aluminum
519 micas described by Monier and Robert (1986), which were the products of metasomatic
520 alteration. Altogether, the characteristics of the compositional zoning and dissolution
521 textures (Fig. 5) of the micas in our samples indicate that they were produced by fluid–
522 rock interactions rather than melt differentiation. Furthermore, the micas would have
523 been enriched in F when they interacted with supercritical fluids exsolved from a highly
524 differentiated magma. The rims of the zoned micas show significant depletion in F
525 (Supplementary Table 6), corresponding to higher $\log(f\text{H}_2\text{O}/f\text{HF})$ values than the mica

526 cores, and this suggests a process of subsolidus replacement by a dominantly external
527 aqueous fluid.

528 The siderites have variable REE contents ($\Sigma\text{REE} = 6\text{--}110$ ppm; [Supplementary](#)
529 [Table 8](#)) and negative Eu anomalies, which would have been caused by the fluids
530 having variable REE contents or pH values ([Michard 1989](#); [Möller et al. 2004](#)). In
531 general, the total REE content of hydrothermal solutions increases with decreasing pH
532 values ([Michard 1989](#); [Möller et al. 2004](#)). In acidic solutions, LREEs are likely to be
533 enriched in aqueous phases, while HREEs are preferentially adsorbed on mineral
534 surfaces ([Schwinn and Markl 2005](#)). Thus, siderites precipitated from fluids with
535 variable pH values would have variable LREE/HREE ratios. The siderites in our
536 samples of muscovite granite and greisen have similar $(\text{La}/\text{Yb})_{\text{N}}$ ratios. Instead, different
537 sources of fluid with variable REE contents are more likely to have been responsible for
538 the wide range of REE contents that characterizes the siderites, and this is consistent
539 with the $\log(f\text{H}_2\text{O}/f\text{HF})$ values of the rims of zoned micas in the greisen being generally
540 lower than those in the muscovite granite ([Fig. 11](#)).

541 The stability of siderite depends on the CO_2 fugacity and the oxygen fugacity
542 ([Fusswinkel et al. 2013](#)). Siderite is stable under conditions of low oxygen fugacity (i.e.,
543 $f\text{O}_2 = -38$ to -54) and high CO_2 fugacity (i.e., $f\text{CO}_2 = -1$ to $+2$) ([Fusswinkel et al. 2013](#)).
544 It is replaced by magnetite in the reaction “6 Siderite + $\text{O}_2 = 2$ Magnetite + 6 CO_2 ”
545 when the oxygen fugacity increases or the activity of CO_2 decreases ([Hovis et al. 1991](#)).
546 The lack of magnetite in the muscovite granite and greisen ([Table 1](#)) may be attributed
547 to a relatively reducing environment where the siderite crystals were stable. Therefore,

548 the presence of siderite in the Xihuashan granite indicates a reducing environment
549 during the fluid–rock interactions, and the oxygen fugacity (fO_2) would have been close
550 to the value constrained by the “C + O₂ = CO₂” buffer (Yang et al. 2013).

551 The fluids associated with the tungsten mineralization would have been of
552 external origin or exsolved from the melt (e.g., Wang et al. 2003; Wei et al. 2012;
553 Harlaux et al. 2018). It is noteworthy that the cores of zoned micas in the greisen and
554 muscovite granite show much lower $\log(fH_2O/fHF)$ values and Fe^{3+}/Fe^{2+} ratios than the
555 rims (Fig. 11), and this indicates that the fluids associated with the formation of the
556 greisen and muscovite granite had a higher fluorine fugacity and were even more
557 reducing than the fluids associated with the fluid–rock interactions. Such fluids might
558 have been supercritical fluids exsolved from the highly fractionated magma, and this is
559 consistent with the observation that the $\log(fH_2O/fHF)$ values associated with the cores
560 of micas in the greisen are close to those associated with magmatic micas in the granite
561 samples (Fig. 11a, b). In addition, based on the fluid inclusion studies, Wei et al. (2012)
562 proposed that a mixture of fluids from two different sources could have been
563 responsible for the deposition of wolframite in the Xihuashan deposit. Fluids with a low
564 oxygen fugacity and enriched in fluorine may be derived from fluid oversaturation of a
565 granitic melt (Maruejol et al. 1990; Wang et al. 2003), which may contribute to the
566 formation of the cores of zoned micas in the greisen and muscovite granite. The
567 fluid-rock interactions, however, involved mainly meteoric fluids that were associated
568 with alteration of early Li-rich micas into late muscovite (Legros et al. 2016). These
569 fluids were also enriched in H₂O and CO₂, as recorded in the rims of zoned mica and the

570 siderite in the greisen and muscovite granite. In particular, the rims of zoned micas in
571 the greisen samples have lower $\log(f\text{H}_2\text{O}/f\text{HF})$ values than those in the muscovite
572 granite samples, which further indicates that the fluids associated with the greisen
573 contained larger amounts of magma-derived fluids.

574

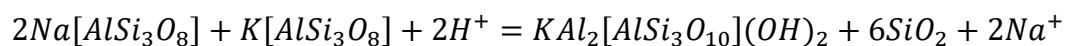
575 **Tungsten enrichment during the evolution of the magmatic–hydrothermal system**

576 Tungsten deposits are usually associated with fractionated granitic magmas and
577 hydrothermal processes (e.g., [Xie et al. 2016](#); [Breiter et al. 2017](#)). In peraluminous
578 granitic systems, tungsten commonly behaves as an incompatible element for most
579 rock-forming minerals and accumulates in melts during fractional crystallization
580 ([Hulsbosch et al. 2016](#)). The WO_3 content of columbite-group minerals increases from 5
581 to 15 wt% during magma evolution ([Xie et al. 2018](#)), which confirms that tungsten can
582 be concentrated in melts. In addition, mica may be used as an indicator of ore-forming
583 potential for tungsten granites ([Yin et al. 2019](#)). In the Xihuashan granite, the tungsten
584 content of micas increases from the biotite granite (0.36–12.8 ppm) to the muscovite
585 granite (14.9–84.1 ppm) while the K/Rb ratio of micas decreases ([Fig. 12c](#)), consistent
586 with the trend of magma evolution. However, magmatic wolframite can crystallize
587 directly from a highly evolved granite melt under restricted conditions, as for example
588 from a peraluminous melt with high concentrations (6 wt%) of fluxing compounds (B,
589 Li, F, and P) and with crystallization temperatures lower than 550 °C at 800 bars ([Che et](#)
590 [al. 2013](#)). In most cases, the required conditions for saturation of magmatic wolframite
591 cannot be reached in granitic melts due to the low levels of fluxing compounds. It is rare

592 to have a wolframite of purely magmatic origin, and hydrothermal fluids typically play
593 a significant role in W mineralization. In the Xihuashan granite, the lack of wolframite
594 in the muscovite granite suggests that the tungsten concentration in the melts did not
595 become saturated relative to wolframite as a result of magmatic fractionation. Indeed,
596 the whole-rock content of W in the Xihuashan granite (4.82–65.3 ppm) is much lower
597 than the solubility of wolframite in peraluminous granitic melts (~1000 ppm WO₃ at
598 750–800 °C; Štemprok 1990). Therefore, although tungsten may be concentrated in a
599 melt via magmatic fractionation, in the magmas of the Xihuashan pluton the W contents
600 were always much lower than the oversaturation level during the evolution of the
601 magma.

602 Hydrothermal activities may play an important role in W mineralization (e.g.,
603 Štemprok 1990; Li et al. 2018). The wolframite crystals in the greisen are intergrown
604 with zoned mica (Fig. 5d), and they would have formed during hydrothermal activity.
605 However, wolframite is absent from the muscovite granite, even though the abundant
606 zoned micas with dissolution textures (Fig 5a, b) provide evidence of fluid alteration.
607 The rims of zoned micas in the muscovite granite have lower W contents than the cores
608 (Fig. 7f), which was probably due to hydrothermal alteration. The hydrothermal fluids
609 liberated W from the micas and resulted in formation of a W-depleted rim, whereas W
610 was concentrated in the hydrothermal fluid. The elevated CO₂ concentration may
611 increase the acidity of fluid and the solubility of tungstates, which facilitates
612 concentrating W in the fluids (Liu et al. 2017; Wang et al. 2020). The rims of zoned
613 micas in the greisen have higher W contents than those in the muscovite granite (Fig.

614 7f), which was probably due to different CO₂ and W concentrations in the hydrothermal
615 fluids. The W-rich rims of zoned mica in the greisen clearly show the involvement of
616 W-rich fluids in the formation of the tungsten deposit (Figs 5d, 7h). Furthermore, a fluid
617 rich in F, Li, and rare metals can induce greisenization in the upper part of the granitic
618 cupola, causing dissolution of feldspars and crystallization of new quartz, zinnwaldite,
619 fluorite, and topaz (Breiter et al. 2017). Greisenization consumes H⁺ and increases the
620 pH value of fluids following the reaction:



621 Albite K-feldspar Muscovite Quartz

622 An increase in fluid pH during greisenization will destabilize the polymeric
623 tungstates to form WO₄²⁻ and other monomeric tungstates (Wang et al. 2020), which
624 may interact with metal cations (i.e., Fe, Mn or Ca) to form wolframite or scheelite.
625 Thus, greisenization promoted the generation of WO₄²⁻ in the fluids, and potentially
626 contributed to the formation of tungsten deposit (Audéta et al. 2008; Harlaux et al.
627 2017; Legros et al. 2019).

628

629 **IMPLICATIONS FOR TUNGSTEN MINERALIZATION**

630 The hydrothermal fluid related to the W deposits would have different sources,
631 including an early magmatic fluid and a later metamorphic or meteoric fluid (Wei et al.
632 2012; Hulsbosch et al. 2016). However, the signature of the early magmatic fluid is
633 usually overprinted by the metamorphic or meteoric fluid (Wei et al. 2012; Hulsbosch et
634 al. 2016). Thus, attempts to determine whether a magmatic or metamorphic fluid was

635 responsible for W mineralization associated with a granite typically provide
636 inconclusive results. In the Xihuashan tungsten granite, the composition and zoning
637 texture of micas may provide some valid constraints on the nature of fluids and the
638 mechanism for precipitation of wolframite. The interaction between external aqueous
639 fluids and pre-existing micas has played an important role in leaching the ore-forming
640 elements (i.e., W and Fe) from the micas into the fluids, which is proved by remarkably
641 dropped concentrations of W and Fe from the cores to the rims of zoned micas in the
642 muscovite granite (Figs 6c, 7h). The W-rich rims of the zoned mica in the greisen,
643 together with the high whole-rock F contents (up to 1.22 wt%) of greisen samples,
644 demonstrate that the fluid related to the greisenization was enriched in F and W. The
645 F-rich fluids involved in the greisenization were most likely derived from the
646 oversaturation of a rare-metal granitic melt (Maruéjol et al. 1990; Launay et al. 2018;
647 Legros et al. 2018). Therefore, the rims of zoned mica in the greisen are enriched in W
648 due to the effect of W-rich fluids, whereas the rims of zoned mica in the muscovite
649 granite are related to the dominant external fluids that had much lower W contents. This
650 difference between the zoned micas in the greisen and muscovite granite indicates that
651 the process of greisenization was critical for extracting and transporting the W from the
652 melts/rocks into the fluids (Launay et al. 2018).

653 The processes identified in the Xihuashan tungsten granite may have been active
654 in many wolframite-bearing quartz-vein deposits elsewhere. Some famous metallogenic
655 provinces or plutons, such as the French Massif Central, the Erzgebirge, the northern
656 Canadian Cordillera, Iberian Massif and the Nanling and Dahutang mineralization

657 regions in South China, are closely related to magmatic and hydrothermal processes
658 (e.g., Legros et al. 2016, 2018; Breiter et al. 2017a, b; Legro et al. 2016, 2018; Xie et al.
659 2018; Yin et al. 2019). Furthermore, the micas in these tungsten-bearing granites are
660 commonly zoned and enriched in fluxing compounds (e.g., Breiter et al. 2017; Harlaux
661 et al. 2018a; Xie et al. 2018; Yin et al. 2019), which might also record the processes of
662 tungsten mineralization related to fluids during their evolution through the magmatic–
663 hydrothermal transition. Collectively, the external and magmatic fluids all facilitated the
664 tungsten mineralization. The overall scenario of the ore-forming processes for the
665 Xihuashan tungsten deposit is summarized as follows (Fig. 13):

666 1) The W and F were gradually concentrated in the magmatic fluids via magmatic
667 fractionation.

668 2) The W and F-rich magmatic fluids pooled in the granitic cupola and induced
669 greisenization. This process destabilized the polymeric tungstates to form WO_4^{2-} in the
670 fluids.

671 3) The pooled fluid flow might have been overpressured, eventually being
672 injected into the vein system.

673 4) Tungsten and the fluids would have migrated along faults and fractures with
674 tungsten being precipitated as vein-type wolframite at favorable sites when the W and
675 F-rich fluid was mixed with an external meteoric fluid.

676 Therefore, tungsten mineralization is closely related to the mixture of magmatic
677 and external fluids. Furthermore, the presence of siderite in the greisen and muscovite

678 granite indicates that the fluids are rich in Fe and Mn, which are ore-forming elements
679 generating metal complexes with WO_4^{2-} during the precipitation of wolframite.

680 The input of Fe and Mn in the fluids would be attributed to the reactions between
681 external CO_2 -bearing fluid and host-rock ([Lecumberri-Sanchez et al., 2017](#); [Wang et al.](#)
682 [2020](#)). In addition, the fluid boiling process might also take the role of triggering the
683 precipitation of wolframite, as the transport of volatile elements (e.g., CO_2 , H^+) into
684 vapor phase changes the chemical composition of the fluid ([Korges et al., 2018](#)).

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1006

FIGURE CAPTIONS

- 1007 **Fig. 1** (a) The distribution of granites and volcanic rocks of different ages in SE China
1008 (modified from [Li et al. 2010](#)), showing the locations of the Nanling W–Sn
1009 polymetallic mineralization region (NPMR) and the Xihuashan tungsten
1010 deposit. (b) Simplified geological map of the Xihuashan pluton (modified from
1011 [Guo et al. 2012](#)), showing four successive intrusions in the Xihuashan mining
1012 district (G1–G4, which correspond to coarse-grained biotite granite,
1013 medium-grained biotite granite, medium- to fine-grained biotite granite, and
1014 fine-grained biotite granite, respectively. After Wang et al. 2003).
- 1015 **Fig. 2** Hand specimen photographs (left) and photomicrographs (right) showing the
1016 petrography of the Xihuashan granites and greisen (in cross-polarized light)
1017 and the different micas encountered in this study. (a) Biotite granite. Biotite
1018 (siderophyllite) and associated apatite and ilmenite inclusions. (b) Two-mica
1019 granite. Biotite (lithian siderophyllite), muscovite and associated apatite and
1020 ilmenite inclusions; the rim of the biotite was eroded, as shown by the irregular
1021 embayed form. (c) Muscovite granite. Muscovite is present between the
1022 feldspar (both plagioclase and K-feldspar) and quartz. (d) Greisen. Biotite
1023 (lithian siderophyllite) and Li-mica occur between quartz grains. Mineral
1024 abbreviations: K-feldspar (Kfs), plagioclase (Pl), quartz (Qtz), biotite (Bt),
1025 muscovite (Ms), apatite (Ap), ilmenite (Ilm), zircon (Zrn), xenotime (Xtm).
- 1026 **Fig. 3** The diagrams of (a) TAS, (b) A/NK vs. A/CNK, and (c–h) TiO₂, MgO, P₂O₅, F,
1027 K₂O, and CaO contents versus Fe₂O₃ for the Xihuashan pluton. The total Fe in

1028 the whole rock analyses is reported as Fe_2O_3 . Plots I–IV from [Guo et al. \(2012\)](#)
1029 represent the data of different intrusions.

1030 **Fig. 4** Chondrite-normalized REE patterns and mean upper crust-normalized
1031 multi-element diagrams for rocks in the Xihuashan pluton. The chondrite and
1032 mean upper crust values are from [Taylor and McLennan \(1985\)](#) and [Rudnick
1033 and Gao \(2003\)](#), respectively. Data in shaded field are from [Guo et al. \(2012\)](#).

1034 **Fig. 5** (a-d) Back-scattered-electron (BSE) images of micas in different intrusions of
1035 the Xihuashan pluton; (e) and (f) representative BSE images of wolframite in
1036 greisen.

1037 **Fig. 6** Diagrams of TiO_2 , Al_2O_3 , FeO^{Tot} , MgO, F, and Cl contents versus SiO_2 for
1038 micas in the rocks of the Xihuashan pluton. BG = biotite granite, TMG =
1039 two-mica granite, and MG = muscovite granite.

1040 **Fig. 7** Diagrams of V, Cs, Rb, Nb, Ta, W, Sn, and Nb/Ta versus $\text{FeO}^{\text{Tot}} + \text{MnO} + \text{MgO}$
1041 contents for micas in the rocks of the Xihuashan pluton. The dashed lines
1042 connect the cores and corresponding rims of zoned mica. BG = biotite granite,
1043 TMG = two-mica granite, and MG = muscovite granite.

1044 **Fig. 8** (a) Representative BSE image, (b) Raman spectra, (c) chondrite-normalized
1045 REE patterns, and (d) mean upper crust-normalized multi-element diagrams for
1046 siderite in the Xihuashan muscovite granite and greisen. Chondrite and mean
1047 upper crust values are from [Taylor and McLennan \(1985\)](#) and [Rudnick and
1048 Gao \(2003\)](#), respectively.

1049 **Fig. 9** (a, b) Variation of measured Li_2O versus SiO_2 and F for micas in the rocks of

1050 the Xihuashan pluton. The regression equations of (1) $\text{Li}_2\text{O} = 0.289 \times \text{SiO}_2 -$
1051 9.658 and (2) $\text{Li}_2\text{O} = 0.3935 \times F^{1.326}$ from [Tischendorf et al. \(1997\)](#) can be
1052 applied to trioctahedral and dioctahedral micas, respectively. (c) Relationships
1053 between measured Li_2O and F in hydrothermal micas, which define a
1054 regression equation of (3) $\text{Li}_2\text{O} = 0.0748 \times F^2 + 0.0893 \times F$ that could be
1055 applied to the hydrothermal micas in our samples. (d) Comparison between
1056 measured and calculated Li_2O for micas in the rocks of the Xihuashan pluton.
1057 Some data for the hydrothermal micas were taken from previous studies ([Van](#)
1058 [Lichtervelde 2008](#); [Akoh and Ogunleye 2014](#); [Xie et al. 2015](#); [Yin et al. 2019](#);
1059 [Ma et al. 2020](#)). BG = biotite granite, TMG = two-mica granite, and MG =
1060 muscovite granite.

1061 **Fig. 10** (a) Variations in the compositions of mica from the Xihuashan pluton on the
1062 diagram of $[\text{Mg} - \text{Li}]$ vs. $[\text{Fe}^{\text{Tot}} + \text{Mn} + \text{Ti} - \text{Al}^{\text{VI}}]$ (according to [Tischendorf et](#)
1063 [al. 1997](#)). (b) Relationships between Li and octahedral cations (based on [Foster](#)
1064 [1960](#)). The arrows indicate the trends in mica compositions due to magmatic
1065 evolution. The shadow area represents the field of the transitional series,
1066 known as the ferrous aluminous micas ([Monier and Robert 1986](#)). D.T. and T.T.
1067 denote the compositional trends of dioctahedral and trioctahedral micas,
1068 respectively. BG = biotite granite, TMG = two-mica granite, and MG =
1069 muscovite granite.

1070 **Fig. 11** Plots of (a) SiO_2 versus $\log(f\text{H}_2\text{O}/f\text{HF})$, (b) $\text{Fe}^{3+}/\text{Fe}^{2+}$ versus $\log(f\text{H}_2\text{O}/f\text{HF})$,
1071 and (c) SiO_2 versus W_F^m for micas in the Xihuashan pluton. BG = biotite

1072 granite, TMG = two-mica granite, and MG = muscovite granite.

1073 **Fig. 12** Plots of Rb, Cs, and W versus K/Rb for micas in the Xihuashan pluton. BG =

1074 biotite granite, TMG = two-mica granite, and MG = muscovite granite.

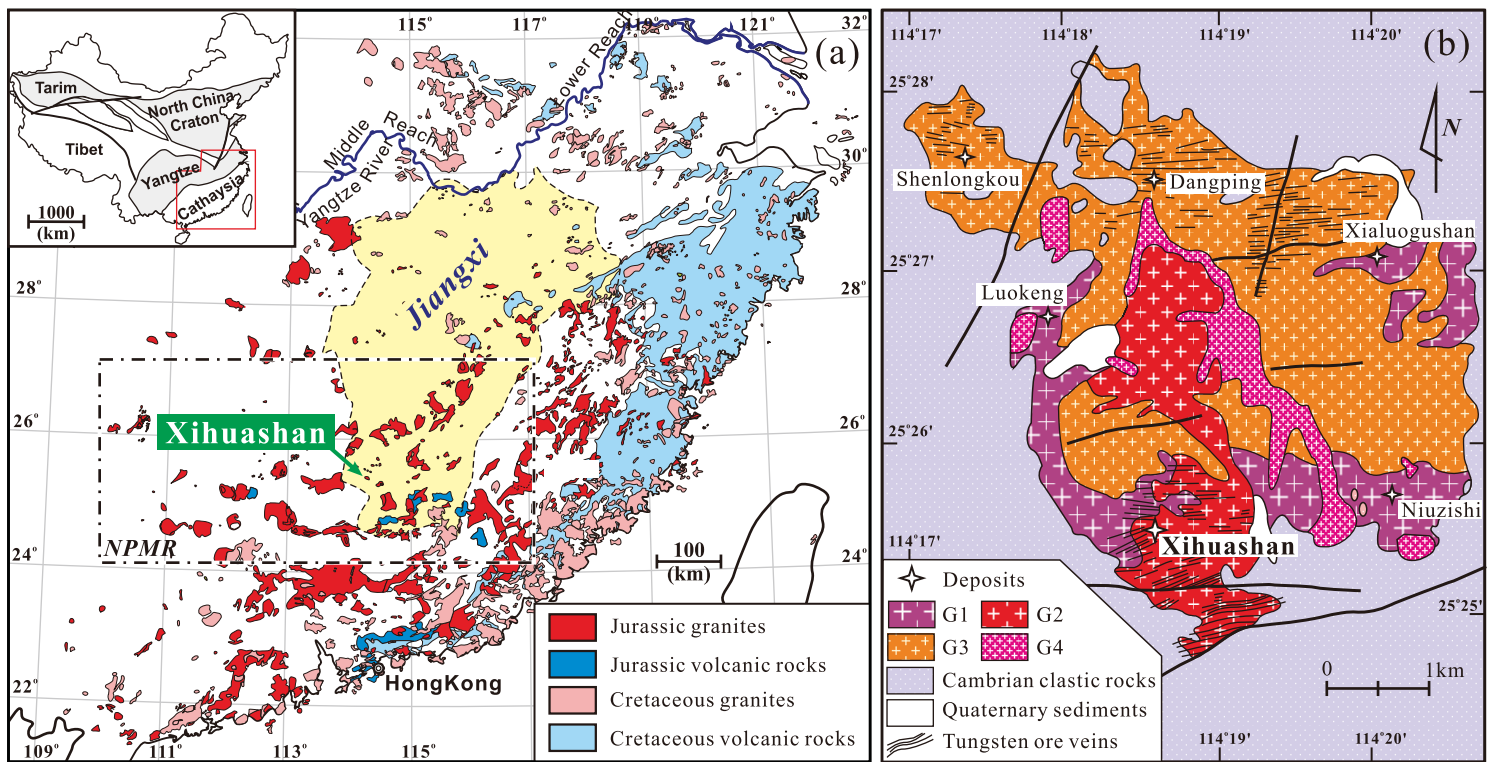
1075 **Fig. 13** Cartoon of tungsten mineralization in the Xihuashan pluton, illustrating the

1076 relationship between the evolution of the magmatic–hydrothermal system and

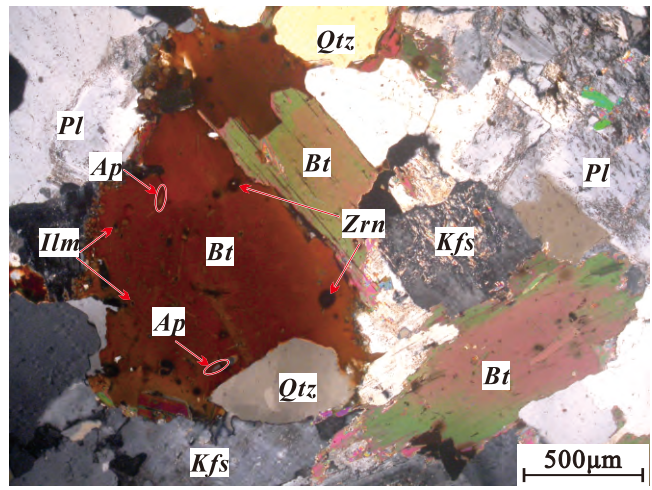
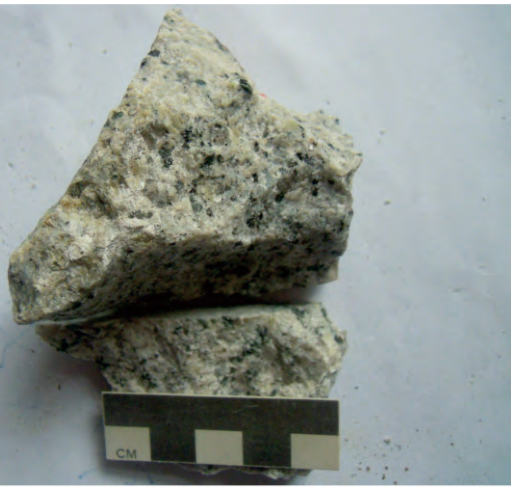
1077 the W mineralization. The spatial relationships between the bodies of granite

1078 and greisen are based on previous investigations ([Liu 2005](#); [Huang et al. 2006](#)).

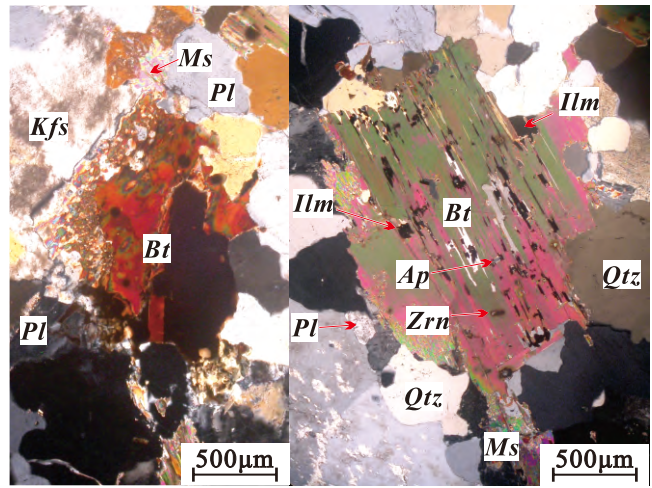
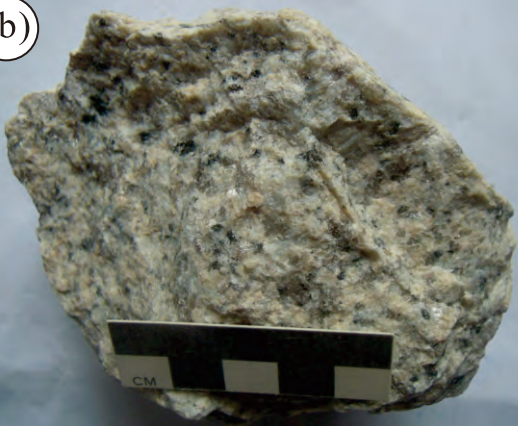
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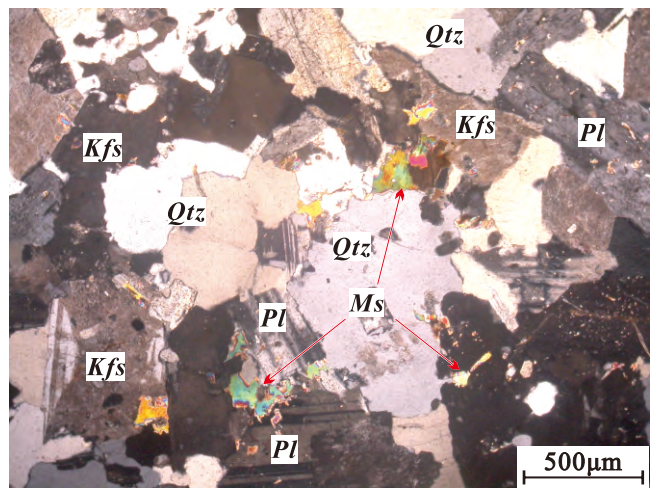
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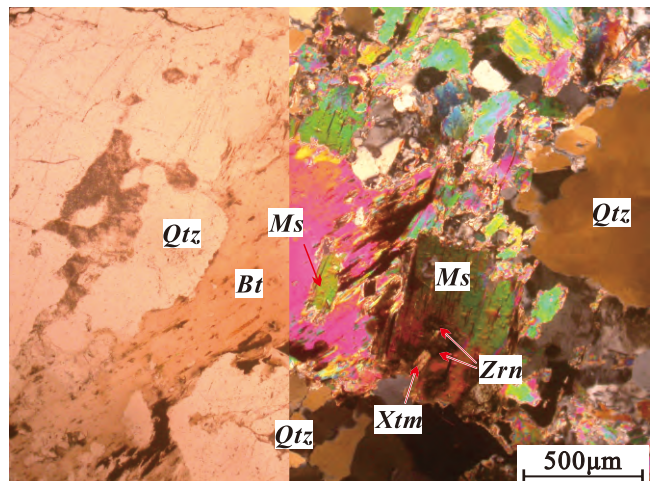
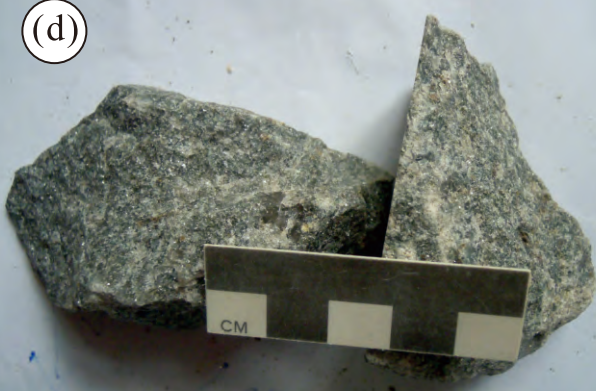
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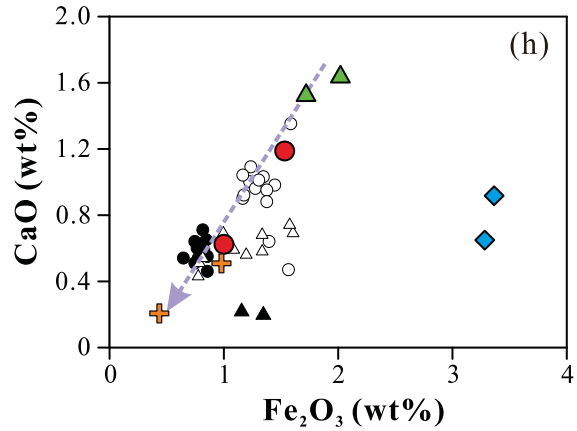
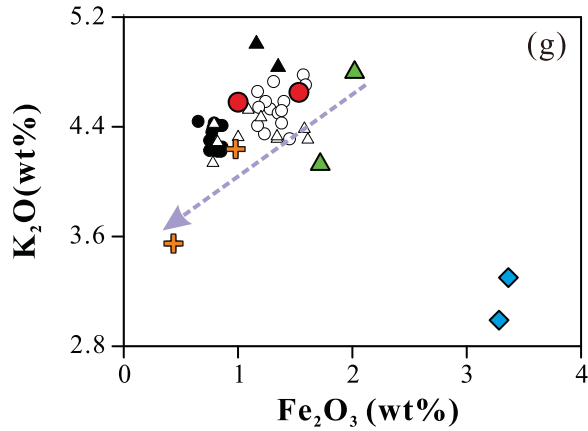
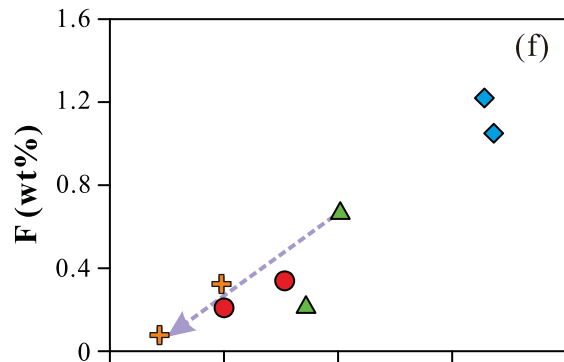
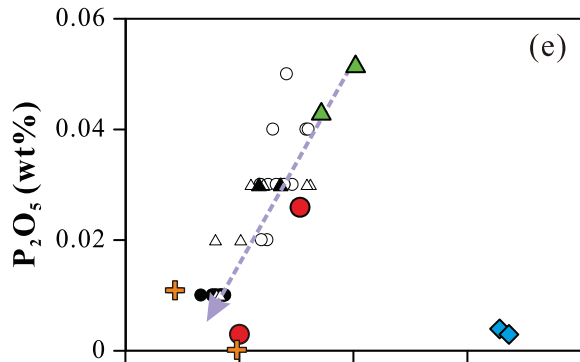
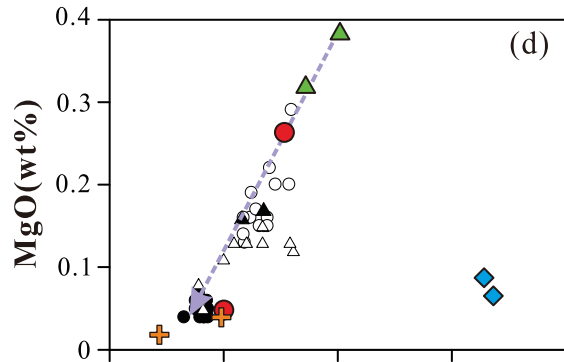
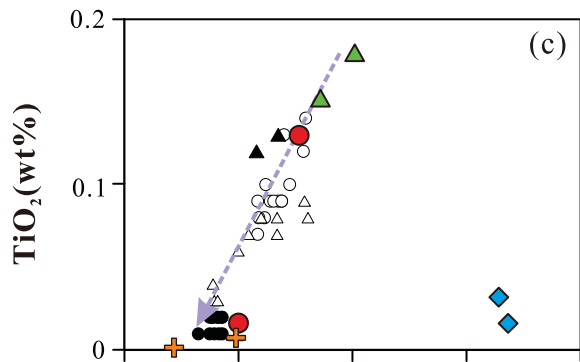
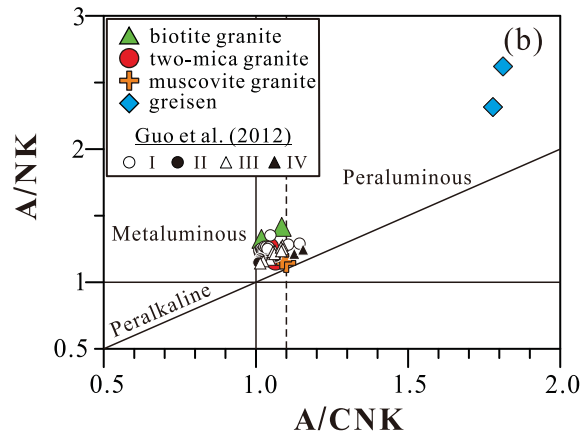
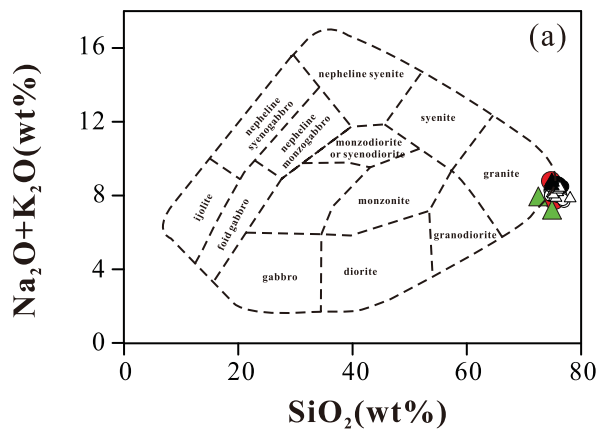


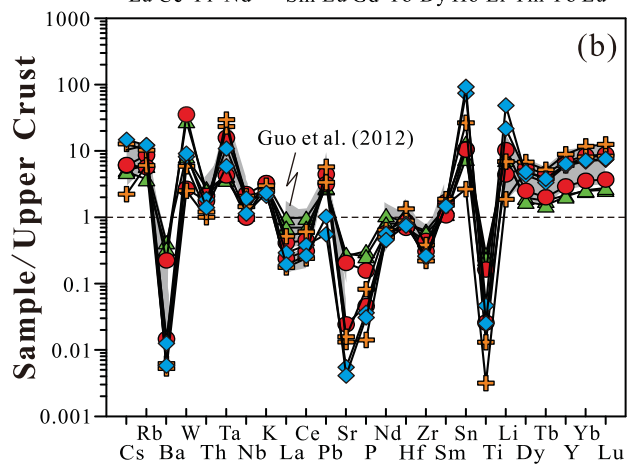
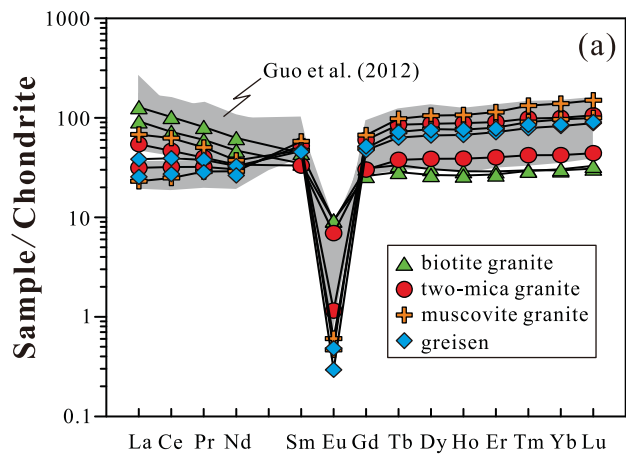
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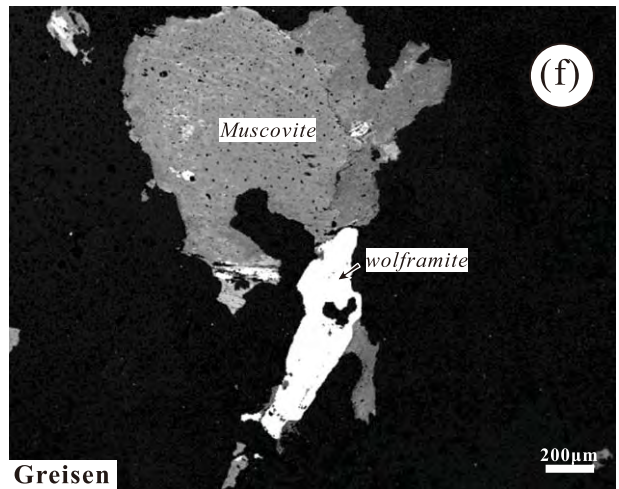
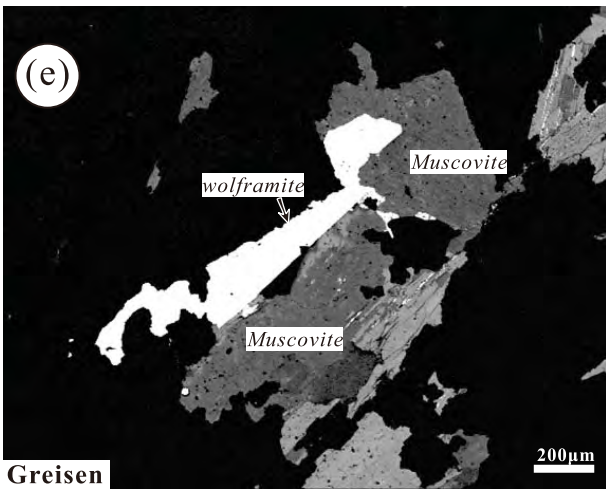
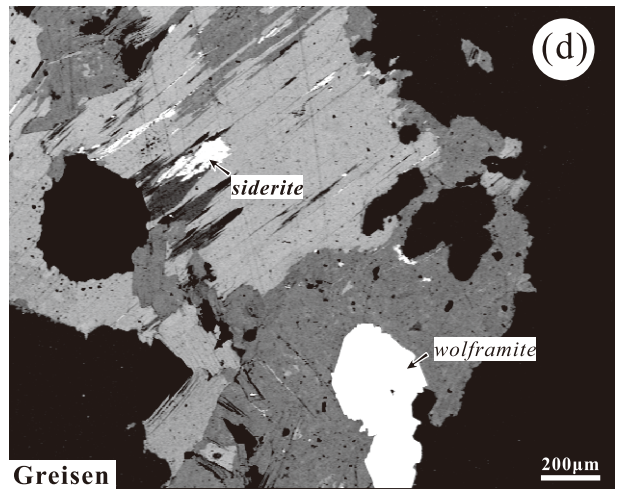
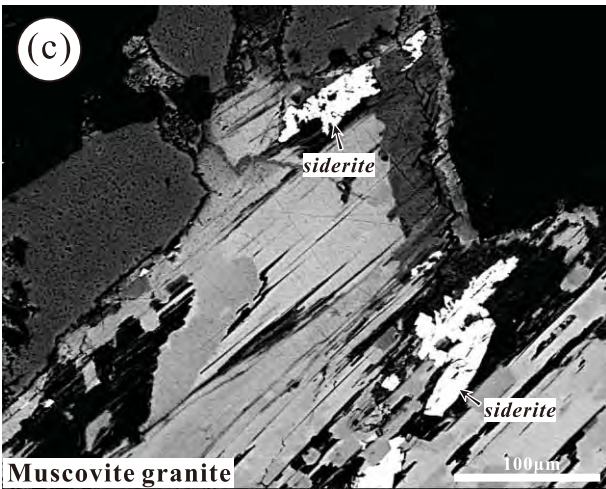
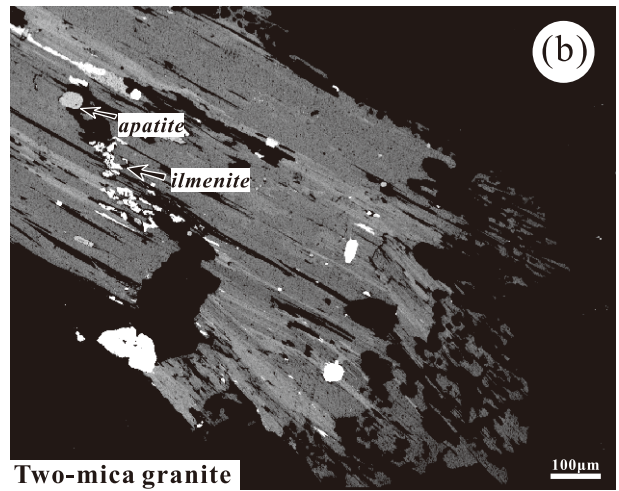
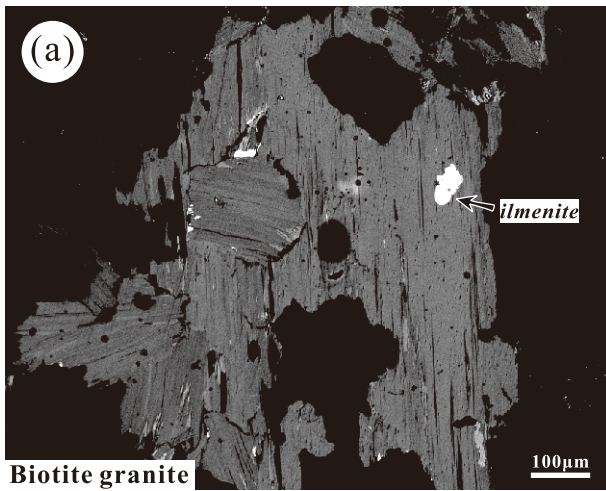


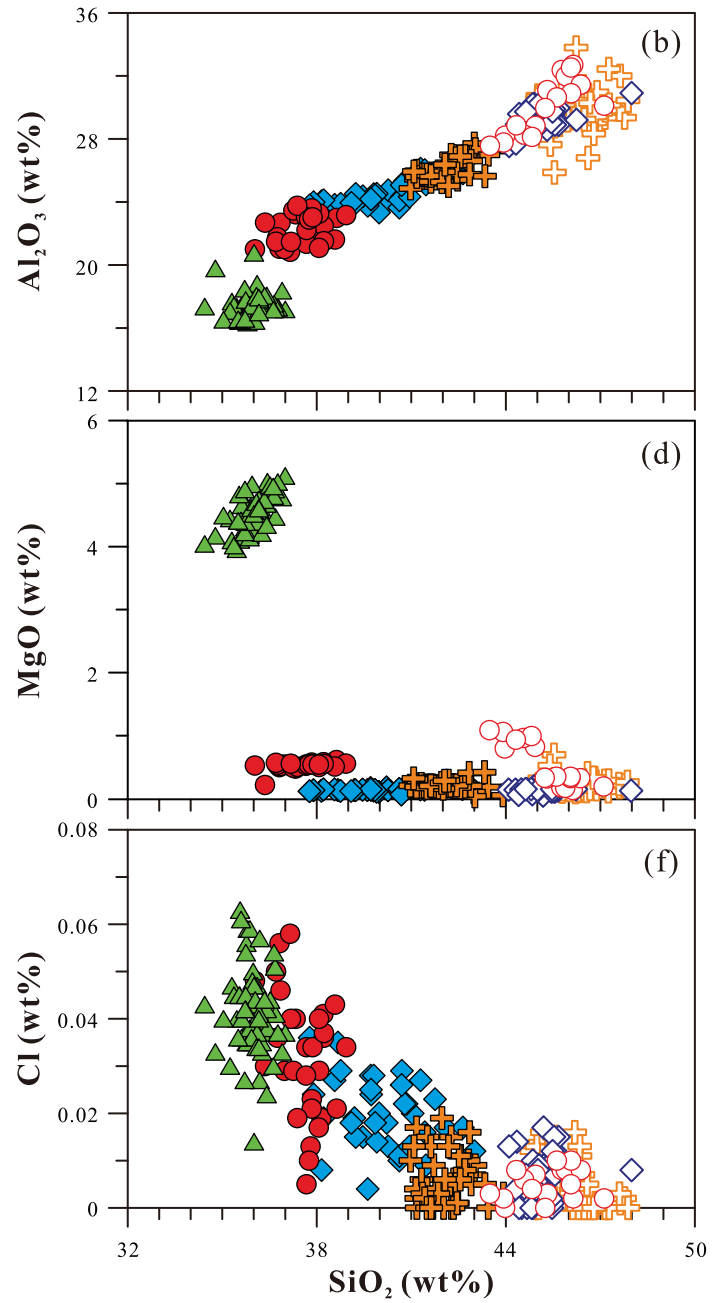
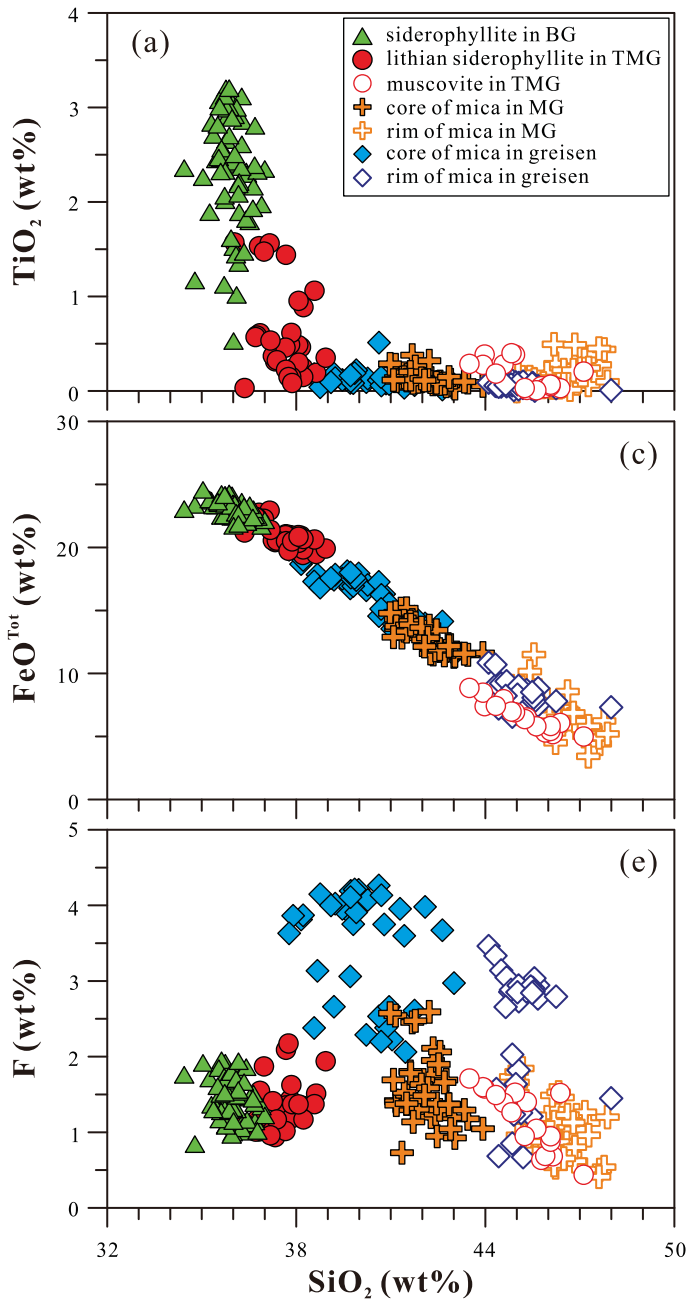
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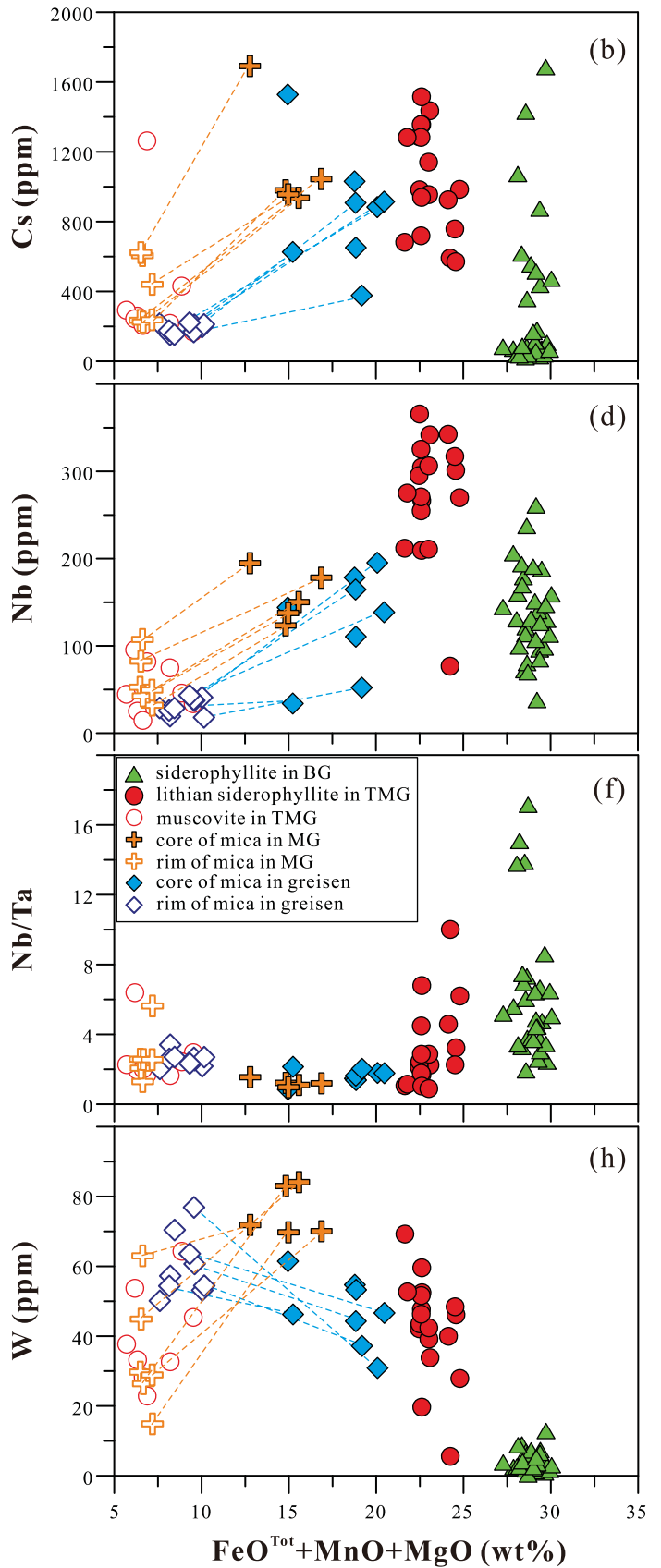
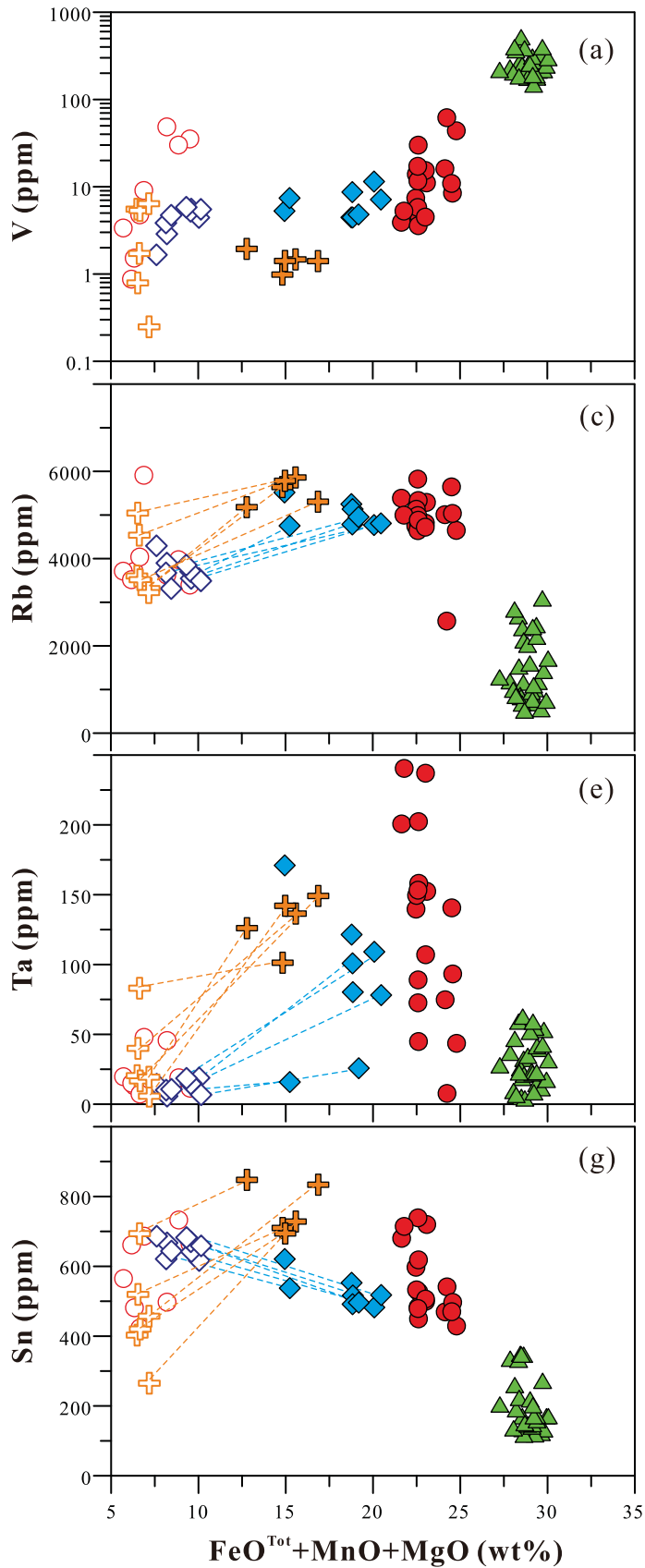


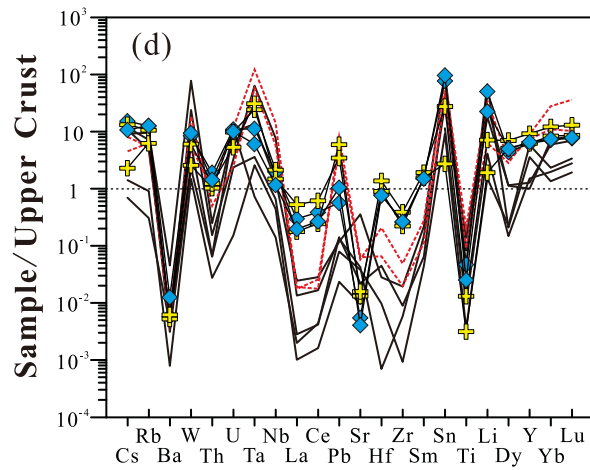
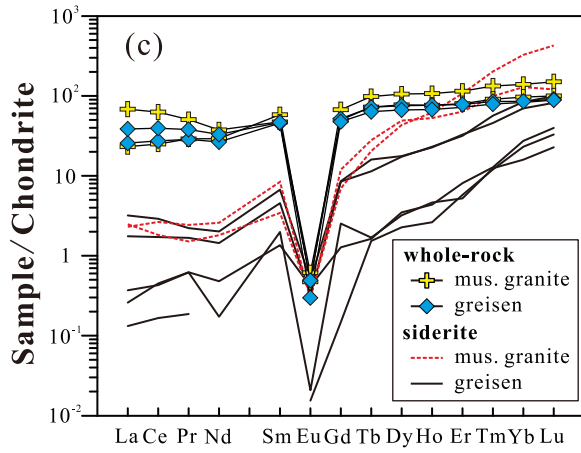
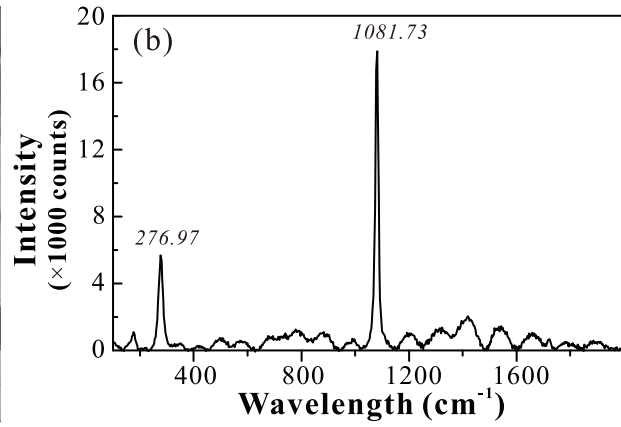
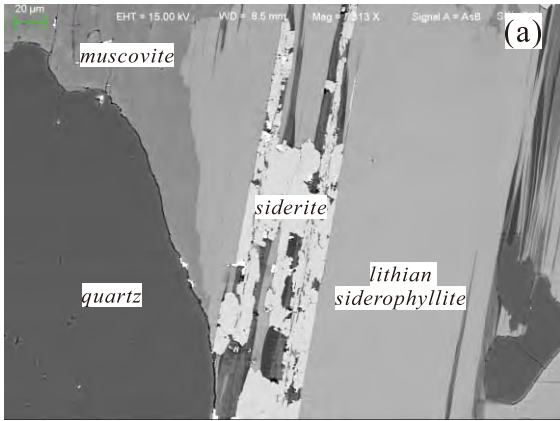


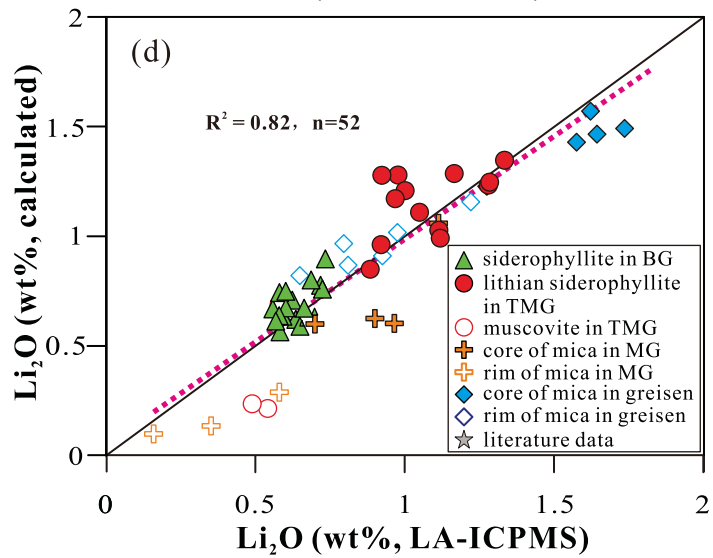
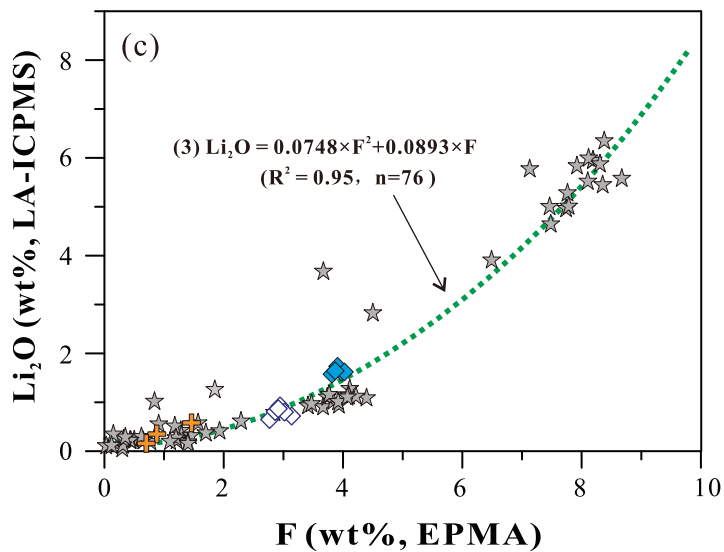
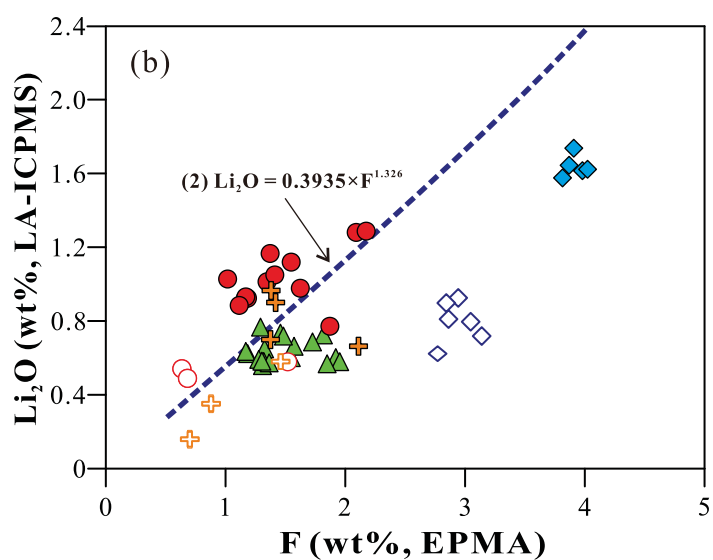
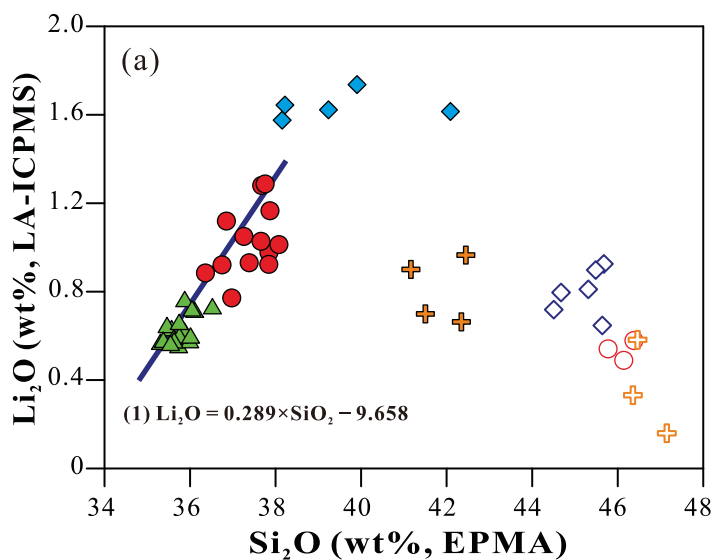


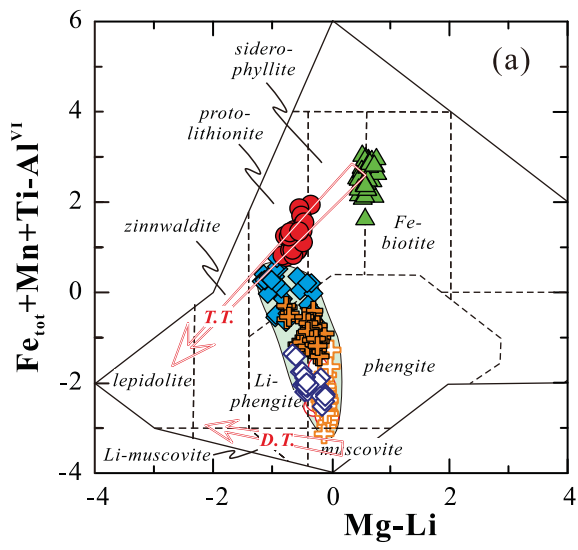




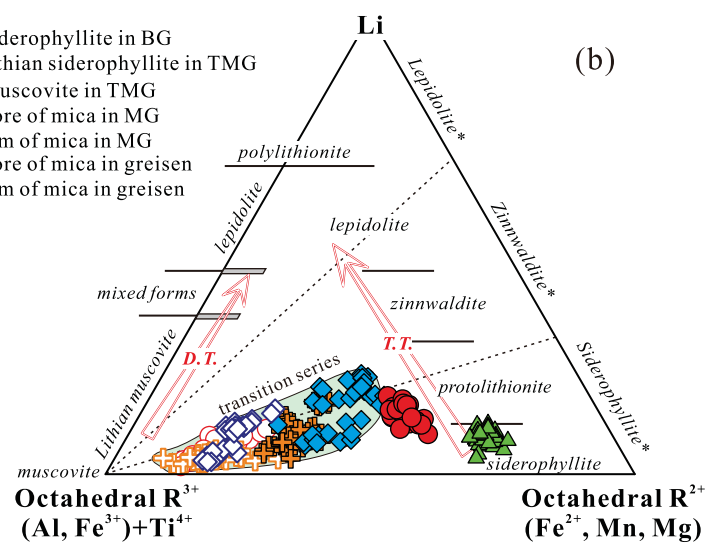








- ▲ siderophyllite in BG
- lithian siderophyllite in TMG
- muscovite in TMG
- ⊕ core of mica in MG
- ⊕ rim of mica in MG
- ◆ core of mica in greisen
- ◇ rim of mica in greisen



(b)

