1   **<u>Revision 2</u>**

2

3   **Jarosite formation in Permian-Triassic strata at Xiakou (South China):**

4   **Implications for jarosite precipitation from H2S upwelling on Mars**

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14   **ABSTRACT**

15   The source of sulfuric acid and associated aqueous alteration of ancient Martian sedimentary

16   rocks remain under debate in the context of divergent models of jarosite formation. Here, we

17   report the formation of sulfates including jarosite in K-bentonites within shallow-water facies of

18   the Permian–Triassic (P-T) transition at Xiakou in South China. In these strata, jarosite is

19   dispersed in the clay matrix or forms aggregates in pore spaces, has a euhedral morphology, and

20   co-exists with variably ^{34}S-depleted paragenetic gypsum and bassanite (δ^{34}S = –37.23‰ to +3.20‰

21   VCDT). Subaqueous alteration of volcanic tuffs concurrently with oxidation of upwelled,

22   biogenically sourced H2S is the process of jarosite formation in the Xiakou K-bentonites. This

23   mechanism of jarosite precipitation and stability over geological time challenges the long-held

24   view of acidic, water-limited conditions leading to iron(III) sulfate precipitation and would be

25   consistent with possible microbial or nanobial life on early Mars.

26   **Keywords:** Tuff; K-bentonite; smectite; jarosite; sulfates; microbial reduction

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31   **INTRODUCTION**

32   Global mineralogical observations of Mars via orbital reflectance spectroscopy have revealed

33   an apparent evolution in the secondary mineral assemblage during the earliest history of the planet

34   (Bibring et al., 2006). The Noachian-aged (4.1-3.7 Ga) stratigraphic units are dominated by

35   Fe/Mg-smectites, with minor amounts of Ca-sulfates and Fe-oxides, whereas the younger

36   Hesperian-aged (3.7-3.1 Ga) units more commonly consist of Mg/Ca/Fe-sulfates, with fewer

37   detections of clay minerals, hematite, and hydrated silica (Ehlmann and Edwards, 2014). The

38   widespread clay minerals on Mars record clear evidence for water-rich surface and near-surface

39   environments in Mars' past, since smectite often forms by supergene weathering under relatively

40   warm and humid climate conditions on Earth (Sulieman et al., 2020; Saricaoglu et al., 2021; Hong

41   et al., 2023). After extensive formation of phyllosilicates in the Noachian era, the abundance of

42   sulfates, especially jarosite, in Hesperian-aged sediments is thought to reflect a substantial change

43   in surface-water chemistry from near-neutral to acidic and saline owing to a dramatic change in

44   the climate of Mars (Bibring et al., 2006; Hurowitz et al., 2017; Poulet et al., 2005).

45   Jarosite on Earth often occurs in arid, acidic, sulfate-rich weathering systems linked to

46   supergene weathering of sulfides (Bauer and Velde, 1997) and in fumaroles and acidic-saline gr—

47   oundwaters associated with lake systems (Benison, 2006). It is thermodynamically stable in a

48   narrow pH range (~3-5) and is considered to be diagnostic of sulfuric acid oxidation and

49   water-limited conditions (Potter-McIntyre and McCollom, 2018; Yoshida et al., 2018). Thus, the

50   formation processes and environmental conditions experienced by jarosite-bearing Martian rocks

51   have been inferred largely from the presence of jarosite.

52   In addition to orbital detection, jarosite has been identified in situ by the *Opportunity* rover

53   on the Meridiani Planum and the *Curiosity* rover in Gale crater. Compositional data and images

54   from the ground allow for more detailed interpretations of environments that led to jarosite

55   formation. The evaporitic sandstones of the Burns Formation on the Meridiani Planum contain up

56   to 10 wt.% jarosite and are thought to have been reworked by eolian and subaqueous processes

57   and further altered by early and late diagenesis, during which jarosite may have formed as an

58   evaporative cement following groundwater recharge (McLennan et al., 2005). Many alternate

59   hypotheses to explain the jarosite have been proposed, including water-limited aqueous alteration

60   followed by extreme aridity to preserve this mineral (Elwood Madden et al., 2004), aqueous

61   oxidation of hydrothermal sulfides (Zolotov and Shock, 2005), fumarolic SO_2-bearing vapor

62   alteration of volcanic ash (McCollom and Hynek, 2005), and sediment reworking in conjunction

63   with acid-sulfate weathering of massive ice deposits (Niles and Michalski, 2009). Jarosite has

64   been detected in early Hesperian-aged, smectite-bearing, and typically Ca-sulfate-bearing

65   fluvio-lacustrine deposits in Gale crater in abundances up to ~3 wt.% (Rampe et al., 2017; 2020).

66   Hypotheses for jarosite precipitation include early or late diagenetic alteration of sulfides

67   (Hurowitz et al., 2017; Rampe et al., 2017). K-Ar dating of one drill powder by the Sample

68   Analysis at Mars (SAM) instrument yielded a relatively young age of 2.12±0.36 Ga for the jarosite,

69   providing further evidence of a late diagenetic origin for jarosite in Gale crater (Martin et al.,

70   2017).

71   Characterizing jarosite-bearing Mars-analog environments on Earth can help us further

72   understand the mechanisms by which jarosite on Mars may have been precipitated and preserved.

73   Here, we report the formation of jarosite and Ca-sulfate in altered volcanic tuffs within

74   shallow-water settings of the Permian–Triassic (P-T) transition at Xiakou, South China. Our

75   observations show that subaqueous weathering of volcanic materials associated with oxidation of

76   upwelling biogenic H_2S from below the chemocline yielded jarosite-bearing smectitic mudstones.

77   The striking mineralogical and depositional similarities to Martian deposits make the formation

78   model and environmental conditions of the Xiakou K-bentonites a reasonable terrestrial analog for

79   fine-grained sediments of the Red Planet, thus providing new insights into the geological

80   processes that precipitated jarosite in Martian rocks.

81

82   **SAMPLING AND METHODS**

83   *Sampling*

84   The Xiakou stratal succession (31°6′51″N, 110°48′13″E) comprises the Permian Chihsian to

85   Lower Triassic Smithian stages. The depositional setting of the Xiakou region was a distal

86   carbonate ramp, characterized by shallow to somewhat deeper water, open-shelf facies during the

87   Permian–Triassic (P-T) transition (Pei et al., 2020). The 1.8-m-thick P-T transition interval

88   consists of gray mudstones and pale-yellow marls, interbedded with eight thin layers (each 1-4 cm

89   thick) of altered volcanic tuff (K-bentonite) (Fig. 1). Most tuff beds in the Xiakou section have

90   sharp contacts with underlying and overlying background deposits, indicating limited vertical

91   sediment mixing. The profile was first cleaned by removing loose surficial materials prior to

92   sampling. Bulk samples, each weighing ~500 g, were collected from each ash layer (*N*=8) and its

93   host rock (*N*=15) from the profile. The soft volcanic ash was excavated using a small trowel from

94   between layers of the hard host rock.

95
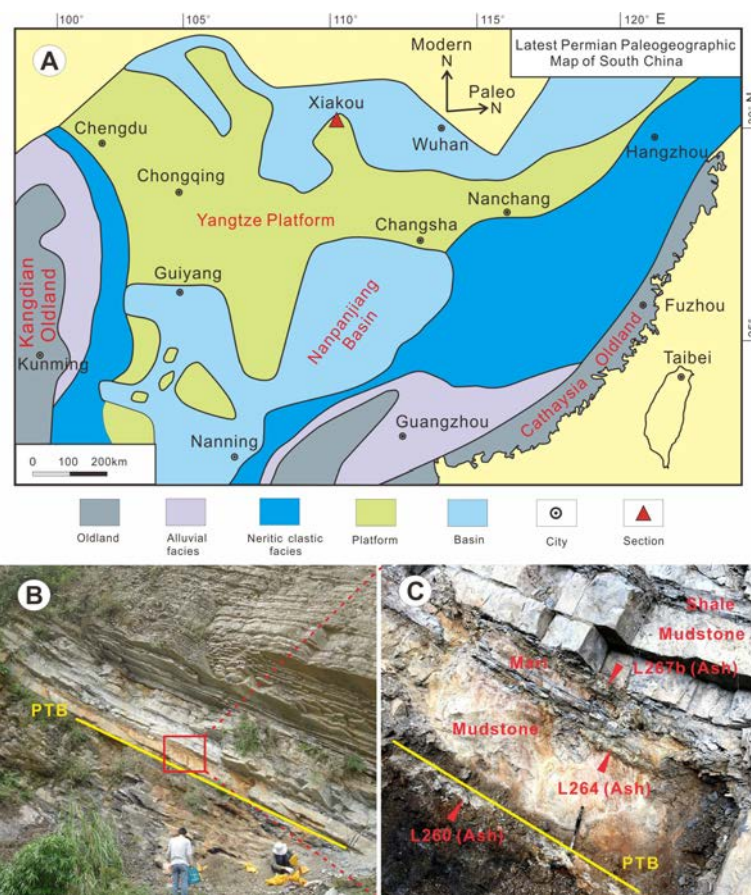
97   Fig. 1 Location and lithology of the Xiakou section. (A) Paleographic map of South China during the
98   Permian-Triassic transition (~252 Ma) (modified from Algeo et al. (2013); (B) The full Xiakou P-T transition
99   sequence; (C) A close-up photograph showing interbedding of tuff beds with marl and mudstone (pen length 15
100 cm).

101

102 *X-ray diffraction（XRD）*

103   Bulk samples were dried in an electric dry oven at 60 ◦C overnight and then ground to powder.

104 The clay fraction (< 2 μm) was extracted using the sedimentation method. The relative proportions

105 in the purified clay samples were determined by fitting the XRD spectra of glycolated oriented

106 clay samples using the NEWMOD II program (Reynolds and Reynolds, 1996). The oriented clay

107 samples were prepared by pipetting the clay suspension onto glass slides, and the glycolated

108 treatment was performed by treating the air-dried oriented samples with ethylene glycol vapor in a

109   desiccator at 65 °C for 4 h. Mineral compositions of the bulk samples were obtained according to

110   the XRD results of non-oriented powdered samples. An internal standard (zincite, ZnO) X-ray

111   diffraction method was used to estimate the abundances of mineral phases in bulk samples. These

112   bulk samples were prepared by mounting the powders into sample holders using a back-press

113   technique to yield non-oriented powdered samples. X-ray diffraction analysis was undertaken on a

114   Panalytical X'Pert PRO DY2198 diffractometer at 35 kV voltage and 30 mA current with

115   Ni-filtered $Cu_{K\alpha}$ radiation. The XRD patterns were collected from 3° to 30° for oriented clay

116   samples and from 3° to 65° 2θ for bulk samples at a scan rate of 4° 2θ/min, which were performed

117   in three times respectively. The mineral intensity factor (MIF) values of these minerals were first

118   measured at routine experimental conditions, and the content (wt.%) of a non-clay mineral in the

119   bulk sample was calculated by the formula: $\%X=(I_X \cdot M_S \cdot 100)/(I_S \cdot MIF \cdot M)$, where M_X and M are

120   the masses of mineral X and the sample to which standard is added, I_X and I_S are the intensities of

121   reflections belonging to X and S in a mixture of a sample with the standard, and %X is the actual

122   amount of the mineral in a sample without the standard (Środoń et al., 2001). The content of total

123   clay minerals in a bulk sample was estimated from the sum of measured minerals departing from

124   100%. The characteristic peaks for mineral identification include 3.11 Å (jarosite), 2.68 Å

125   (gypsum), 3.00 Å (bassanite), 2.70 Å (pyrite), 3.03 Å (calcite), 2.90 Å (ankerite), 2.88 Å

126   (dolomite), 4.26 Å (quartz), 3.19 Å (plagioclase), and 3.22 Å (orthoclase). The mineral intensity

127   factor values for these minerals are 0.512 (jarosite), 0.302 (gypsum), 0.456 (bassanite), 0.94

128   (pyrite), 1.07 (calcite), 0.92 (ankerite), 0.82 (dolomite), 0.29 (quartz), 0.71 (plagioclase), and 0.45

129   (orthoclase). The relative error of measurement of mineral phase content is generally <15%.

130

131   *Scanning electron microscope（SEM）*

132   Small blocks of ash samples were prepared, epoxied with resin, polished to produce a smooth

133   surface, and subsequently carbon-coated. Microscopic observation was performed on a SUPRA 55

134   thermal field emission scanning electron microscope (FESEM) equipped an energy-dispersive

135   spectrometer (EDS) and operated at 20 kV accelerating voltage and a beam current in the range of

136   1–3 nA. The chemical compositions of mineral particles were determined using the EDS system,

137   with backscattered electron imaging used to analyze finer-grained particles.

138

139 *High-resolution transmission electron microscopy （HRTEM ）*

140   HRTEM observations were undertaken on a FEI TECNAI G2 S-TWIN field-emission HRTEM

141 equipped with an EDAX solid-state EDX detector and a Gatan 830 charge-coupled device (CCD)

142 camera system. The instrument was operated at an accelerating voltage of 200 kV, with a

143 point-to-point resolution of 0.24 nm and a line resolution of 0.14 nm. The area selected for

144 electron diffraction was usually at the relatively thin edge of crystals, and the diffraction patterns

145 were analyzed using the Digital Micrograph software version 3.9 (Gatan Ltd.) and compared to

146 standard diffraction patterns of jarosite (PDF 71–1777; Formula of $K(Fe_3(SO_4)_2(OH)_6$,

147 rhombohedral R-3m symmetry, and cell parameters: $a_0 = b_0 = 7.315$ Å, $c_0 = 17.224$ Å).

148

149 *Major- and trace-element analyses*

150   The major element compositions of the samples were measured using X-ray fluorescence (XRF)

151 spectroscopy. Samples were prepared as fused pellets using dilithium tetraborate-polyvinyl

152 alcohol. The measurement of major element compositions was undertaken on a Shimadzu

153 XRF-1800 sequential XRF spectrometer with a rhodium tube at a 2.5 kW generator. The detection

154 limit for major elements was ~0.01 wt%, and the relative standard deviation was <1%. Trace

155 element analysis of whole-rock sample was conducted on an Agilent 7700e ICP-MS. The samples

156 were digested in a mixed solution of hydrofluoric acid and nitric acid. The relative standard

157 deviation was ~4% for REEs and 5–10% for other trace elements.

158

159 $^{34}S_{CaSO4}$ *composition analysis*

160   Gypsum and bassanite grains (usually 30-200 μm) were glued to a sample target with 50 grains

161 per sample. In-situ sulfur isotope analysis of gypsum and bassanite was conducted on a Neptune

162 Plus multireceiver inductively coupled plasma mass spectrometer (MC-CIP-MS) at mass resolving

163 power (Rpower(5, 95%)) of 4996 equipped with an NWR FemtoUC femtosecond system. In the

164 laser system, helium was used as the carrier gas for the ablation cell and was mixed with argon

165 (makeup gas) after the ablation cell. The single spot ablation mode was used with a spot size of 40

166 μm, a slow pulse frequency of 4 Hz, and a laser fluence of ~2.5 J/cm^2. The $^{34}S/^{32}S$ ratio of the

167 standards and samples were directly analyzed using static mode, with the total acquisition time for

168 each measurement being approximately 52 s. $δ^{32}S$ and $δ^{34}S$ values were calculated using the

169   calibration method of standard-sample bracketing (SSB) to avoid matrix effects. The natural barite

170   standard GTS ($\delta^{34}S_{V\text{-}CDT}$ = +4.7±0.4‰) was used as a reference material for calibration of isotopic

171   results. The in-house natural barite standards YF-1 ($\delta^{34}S_{V\text{-}CDT}$ = +18.5±0.7‰) and YF-2 ($\delta^{34}S_{V\text{-}CDT}$

172   = +18.1±0.4‰) were analyzed repeatedly as unknown samples to verify the accuracy of the

173   calibration method (Fu et al., 2017). The final $\delta^{34}S$ value of each sample was the average of 15 to

174   50 point analyses, with a relative standard deviation usually better than 0.7‰.

175

176   **RESULTS**

177   *Mineral compositions of the Xiakou sediments*

178   X-ray diffraction evidence shows that jarosite, gypsum, and bassanite are present as minor

179   components in K-bentonites and only occasionally in their enclosing host mudstones (Fig. 2;

180   Table 1). The mineral compositions of the tuff beds are dominantly clay minerals and minor quartz,

181   plagioclase, jarosite, gypsum, bassanite, and pyrite, with trace amounts of calcite and ankerite in

182   some beds (Fig. 3; Table 2). The mudstones consist dominantly of clay minerals, quartz, and

183   plagioclase, with minor pyrite and occasionally jarosite and gypsum, while the marls contain

184   mainly calcite, ankerite, clay minerals, and minor amounts of quartz, plagioclase, and pyrite.

185   Jarosite exhibits an intimate association with gypsum. Clay phases in the tuffs are dominantly

186   mixed-layer illite/smectite (I/S) clays, whereas those in the mudstones and marls consist mainly of

187   illite and minor chlorite, with I/S present only in the lower part of the profile (Table 1).

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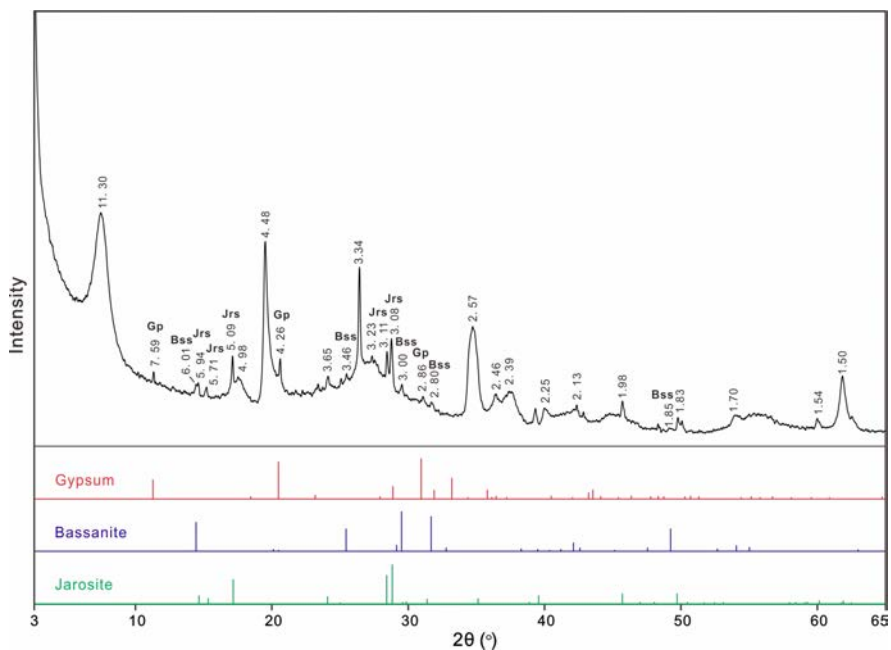
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190   Fig. 2. Determination of jarosite (Jrs), gypsum (Gp), and bassanite (Bss) by XRD analysis of representative

191   whole-rock tuff sample (L255).
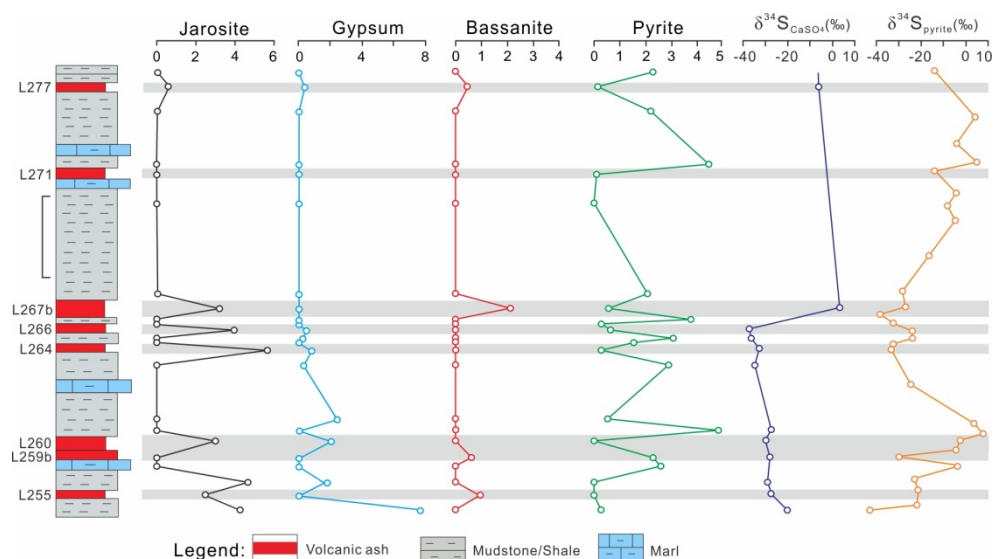
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193   Fig. 3 Sulfate content and sulfur isotopic composition of Xiakou profile. (The scale bar is 30 cm in length; $\delta^{34}S_{pyrite}$

194   curve adapted from Shen et al. (2013))

195

196   *Occurrence of jarosite in Xiakou tuffs*

197   Back-scattered electron (BSE) microscopic observations and elemental mapping by SEM

198   show that jarosite occurs both dispersed through the clay matrix and as aggregates in small pore

199   spaces, rather than coating pyrite grains (Figs. 4A-D and 5). HRTEM observations show that

200   jarosite exhibits both euhedral hexagonal and pseudo-cubic morphologies, with particle sizes

201   mainly of 150-400 nm and 50-150 nm, respectively. Both hexagonal and pseudo-cubic crystals

202   display a uniform diffraction

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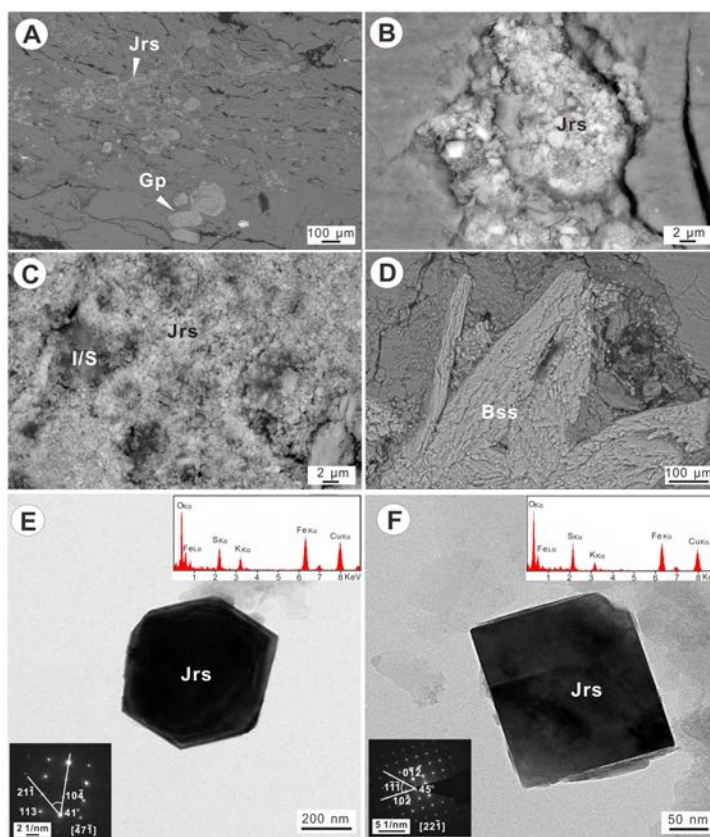
Fig. 4 Microscopic observations of authigenic minerals in the Xiakou tuff. (A) Jarosite nanoparticles associated with euhedral gypsum crystals (BSE); (B) Aggregated jarosite in a pocket within clay matrix (BSE); (C) Enlarged view of jarosite showing fine-grained morphology on clay flake substrate (secondary electron image); (D) Bassanite in subhedral crystals with orientation along the cleavage directions of gypsum showing its formation from dehydration of a gypsum precursor (BSE); Hexagonal (E) and pseudo-cubic (F) crystal habits of jarosite with electron diffraction patterns and energy-dispersive spectroscopy (EDS) spectra by HRTEM observation. (Jrs: jarosite; Gp: gypsum; I/S: mixed-layer illite/smectite; Bss: bassanite; BSE: back-scattered electron).

213   pattern compatible with the trigonal system of jarosite (R-3m), and the chemical composition

214   resembles that of typical jarosite, *i.e.*, $KFe_3^{+3}(SO_4)_2(OH)_6$, as confirmed by EDS data (Fig. 4E, F).

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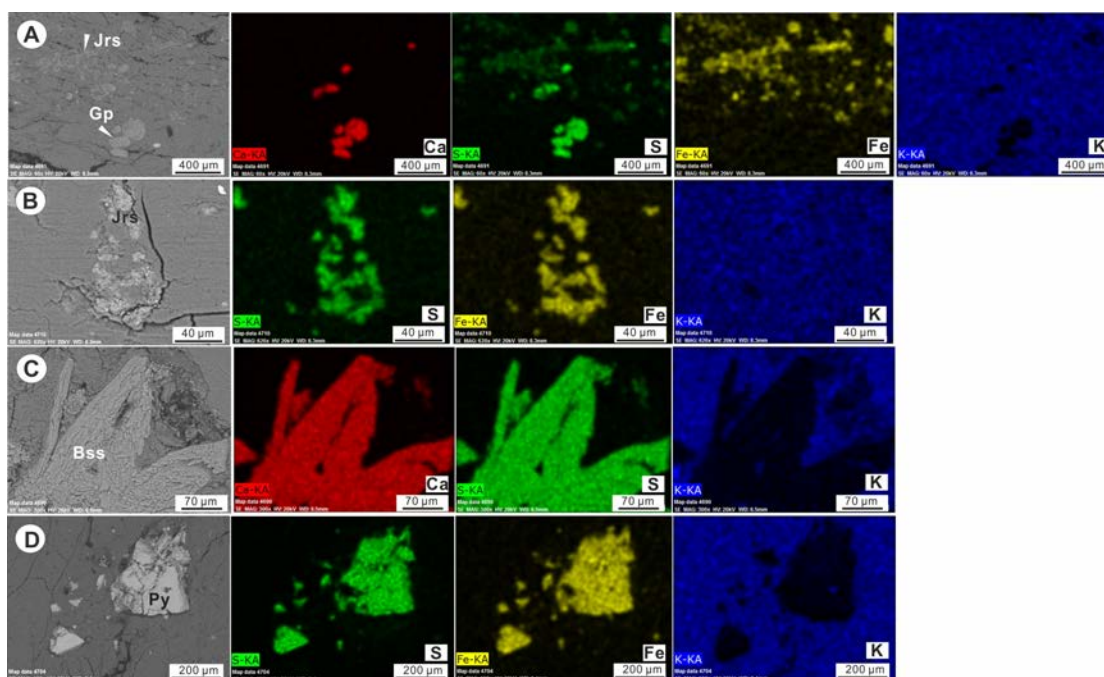
Fig. 5 Elemental mapping for jarosite and co-existing sulfates in the Xiakou K-bentonite. Overlapping distributions of (A) Ca and S indicate probable gypsum, and S, Fe, and K indicate dispersed jarosite; (B) S, Fe, and K reveal jarosite in a pocket-like space; (C) Ca and S indicate bassanite, supported by its crystal morphology; and (D) S and Fe indicate pyrite (Py).

Gypsum, which occurs in pore spaces of the clay matrix as euhedral, micrometer-sized (< 30 to 200 μm), isolated prismatic crystals or intergrown crystals (Figs. 4A and 5). Bassanite occurs in aggregates in pore spaces, where fine-grained (2 to 5 μm) crystals display a subhedral elongated prismatic shape typically oriented along a good cleavage direction {010} or the distinct cleavage planes of gypsum {100} and {011} (Figs. 4D and 5). Pyrite is widely present in the tuff beds and background sediments. It displays euhedral cubic and octahedral and, sometimes, irregular morphologies, being mainly 10-40 μm in size (Fig. 5D).

Geochemical compositions of the Xiakou sediments

The Xiakou tuffs have generally compatible major element compositions with relatively less variation between samples from different layers, consisting of mainly SiO_2, Al_2O_3, and K_2O (Table S1). The contents of SiO_2, Al_2O_3, and K_2O of the tuff samples ranged from 48.22 to 57.82%, 16.32 to 25.49%, and 4.49 to 6.18%, respectively. TiO_2 content is generally low (0.355–0.854%). The $Fe_2O_3(t)$ content changed significantly from 1.33 to 5.92%, consistent with their varied mineral abundances of jarosite and pyrite. CaO ranged greatly from 0.91 to 4.67%, corresponding to

238   changes mainly in sulfate and carbonate contents. The tuff samples have relatively low and quite

239   uniform MgO content (2.43–3.69%) compared to those of adjacent normal background sediments.

240   Na$_2$O content is low (0.07 to 0.58%), with the single exception of sample L267b (1.04%) due to

241   the presence of much more plagioclase. LOI values changed in a narrow range of 6.59 to 11.92%,

242   consistent with the predominance of clay minerals determined by XRD analysis. Results of trace

243   element and rare earth element analyses are reported in Table S2. Plots of TiO$_2$ versus Al$_2$O$_3$ and

244   Zr/TiO$_2$ versus Nb/Y of the Xiakou tuffs show that the tuffs have felsic-intermediate source rocks

245   (Fig. S1).

246

247   *δ^{34}S values of sulfates in the Xiakou sediments*

248   The δ^{34}S of gypsum and bassanite (n.b., jarosite was not measured due to small particle size) in

249   the lower part of the profile ranges from –37.23 to –21.80‰ (Table 2), which is markedly lower

250   than that of contemporaneous seawater sulfate of (~+15 to +25‰) (Kampschulte and Strauss,

251   2004) or magmatic sulfur (~0‰) (Marini et al., 2011). For some relatively larger grains, δ^{34}S

252   value showed zoned distribution from the center to margin of the crystals. For example, the zoned

253   δ^{34}S value of Sample L255 (tuff) increases from -30.74 to -24.88 and to -19.38‰, while that of

254   L277 (mudstone) decreases from -15.12 to -20.11‰ (Table 3), suggesting the evolution of

255   porewater with diagenesis.

256

257   **DISCUSSION**

258   *Precipitation of sulfates in the Xiakou tuffs*

259   Weathering of tuffs can occur quickly in syndepositional to early diagenetic environments,

260   and the dissolved Si and Al released by dissolution of volcanic materials can lead to oversaturation

261   of sediment porewater with respect to smectite (Christidis, 1998). In addition to the Si and Al

262   incorporated by neoformed smectite, excess silicon can form Si-rich phases such as quartz,

263   opal-CT, mordenite, and plagioclase or leach away. Authigenic smectite is usually converted to

264   mixed-layer I/S via K-incorporation in the burial diagenetic environment, depending on burial

265   temperatures and potassium availability (Somelar et al., 2010). The tuff beds in the Xiakou profile

266   contain dominantly I/S clays, with minor plagioclase, quartz, carbonates, and sulfides, and

267   variation in the smectite layer content of the I/S clays between tuff beds is likely attributable to

268   differences in parent-rock composition (Table 1; Fig. S1). This relationship is also reflected by

269   relatively larger amounts of plagioclase, quartz, jarosite, and gypsum in the tuff beds having

270   intermediate source rocks compared to those with felsic source magmas.

271   Microscopic observations showed that there is no evidence of potential precursor ferrous

272   phases such as schwertmannite for formation of jarosite (Fig. 4A-C), suggesting that jarosite is

273   unlikely to have formed via alteration of reduced Fe-minerals. In particular, the euhedral crystals

274   of jarosite indicate that their precipitation took place directly from sediment porewaters under

275   relatively slow precipitation kinetics, and that the two distinct morphologies of jarosite probably

276   reflect two periods of formation under slightly different conditions (Fig. 4A-F). The disseminated

277   and pocket-like occurrences of jarosite, especially on flakey clay substrates (Fig. 4C), suggest that

278   it formed during the post-depositional to early diagenetic stages, prior to obvious compaction of

279   the sediment.

280   The euhedral, isolated or intergrown crystals of gypsum in pore spaces of the Xiakou tuffs

281   exhibit closely association with jarosite, indicative of initial precipitation from porewater (Figs.

282   4A and 5). The precipitation of gypsum likely took place before significant sediment compaction,

283   concurrent with or slightly later than jarosite formation. Also, the large size of single crystals and

284   crystal clusters suggests precipitation of gypsum by relatively slow growth in moderately

285   supersaturated porewaters. In shape and size, the bassanite aggregates resemble those of gypsum

286   crystals and clusters (Figs. 4D and 5), indicating its formation via dehydration of gypsum

287   precursors. Transformation of gypsum to bassanite occurs under water-poor, high-temperature

288   conditions and/or in high-salinity brines, as commonly encountered in the burial diagenetic

289   environment (Hoareau et al., 2011). This transformation was facilitated by the nature of the

290   Xiakou sedimentary environment, which was a moderately deep-water setting characterized by

291   elevated salinity, oxygen depletion, and euxinia, as determined from sedimentary microfacies

292   investigation (Pei et al., 2020). The paragenetic association of jarosite and gypsum in the clay

293   matrix indicates concurrent formation of these minerals, both being related to the presence of

294   sulfuric acid. The basanite originated via dehydration of gypsum and is expected to have inherited

295   its sulfur source. The presence of pyrite is largely confined to void spaces, indicating that it

296   formed during late diagenesis after the sediments had undergone a certain degree of compaction.

297

298   *Formation of jarosite during alteration of tuff by upwelled H_2S*

299   The source(s) of sulfur in the Xiakou profile can be inferred from the S-isotopic compositions

300   of the constituent minerals. Although tuff particles often contain certain amounts of acid volatiles

301   (e.g., SO_2, HF, and HCl) due to adsorption of volcanic aerosols (Cronin et al., 2014), the markedly

302   lower $\delta^{34}S$ values of gypsum and bassanite in the Xiakou sediments (Table 2) exclude its source of

303   seawater sulfate (Kampschulte and Strauss, 2004) or magmatic sulfur (Marini et al., 2011).

304   Gypsum from organic-rich marine sediments has been considered as a marine atmospheric H_2S

305   proxy, and the Xiakou sulfates with notably light isotopic compositions are characteristic of sulfur

306   sourced from sulfate-reducing bacteria in the underlying sediments. Sulfate-reducing bacteria (e.g.,

307   *Desulfovibrio desulfuricans*) enrich metabolically produced H_2S in ^{32}S, resulting in $\delta^{34}S$ values as

308   low as –46‰ (Kaplan and Ritterberg, 1964). The highly ^{34}S-depleted sulfates ($\delta^{34}S$ = –37.23 to –

309   21.80‰) in the lower profile were likely derived from sulfate-reducing bacteria, taking into

310   consideration the organic matter-rich underlying deposits (Pei et al., 2020; Shen et al., 2013) and

311   the notably higher $\delta^{34}S$ values (–6.62 to +3.20‰) in the upper profile (L267b, L277). In particular,

312   some larger crystals of gypsum exhibit zoned distribution of $\delta^{34}S$, increasing gradually from the

313   center to margin of gypsum crystals in tuffs, while showing the opposite trend for gypsum in the

314   adjacent mudstone, indicative of evolution of porewater chemistry in diagenesis. The $\delta^{34}S$ of

315   pyrite fluctuates between +5% and −30% within the Xiakou profile (Fig. 3; Table 2). Heavier $\delta^{34}S$

316   values indicate a dominantly diagenetic origin of pyrite, whereas lighter values reflect a

317   syngenetic origin from a euxinic water column (Shen et al., 2013; Lei et al., 2017).

318   The concurrent formation of jarosite, gypsum, and bassanite required acidic conditions, high

319   salinity, and sources of K^+, Fe^{3+}, and SO_4^{2-} ions. Most the tuff beds in the Xiakou section show a

320   sharp boundary with their enclosing deposits, indicating limited mixing with the normal

321   background sediments, and the highly acidic porewater needed to precipitate sulfates is primarily

322   derived from upwelling of H_2S accompanying sedimentation of volcanic tuffs. The deposition of

323   the volcanic tuffs simultaneously introduced the flux of atmospheric O_2 and resulted in the

324   formation of sulfuric acid ($H_2S + 2O_2 \rightarrow SO_4^{2+} + 2H^+$) (Fig. 6). The existence of acidic porewaters

325   at Xiakou is also confirmed by the distinct crystallographic morphology of authigenic anatase,

326   which has an acute tetragonal-bipyramidal habit in the tuffs (Hong et al., 2022), similar to that of

327   anatase formed in the Verkhnekamskoe salt deposits at a pH of 4.7-5.9 (Chaikovskiy et al., 2019).

328
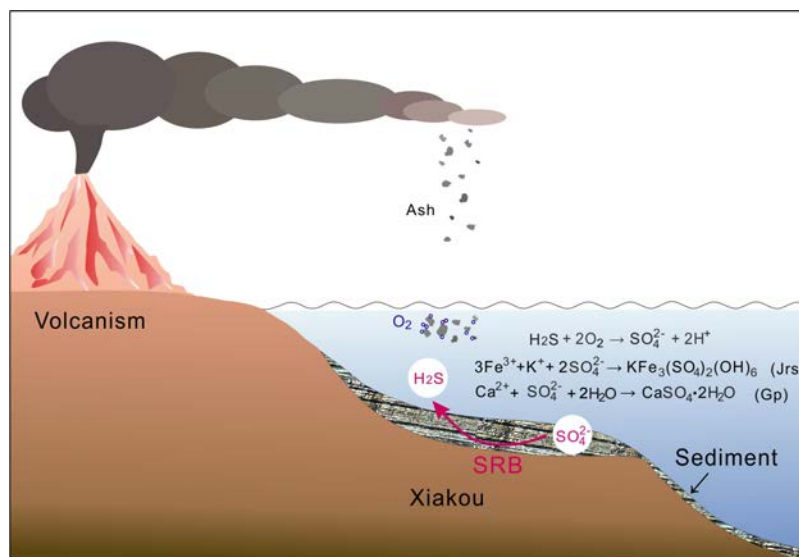
329

330   Fig. 5 Jarosite formation due to volcanic tuff deposition in combination with upwelling of H_2S produced by

331   sulfate-reducing bacteria (SRB). (Jrs- jarosite; Gp- gypsum).

332

333   Potassium is a highly soluble component of volcanic glass that is almost completely leached

334   from tuffs during early diagenesis (De La Fuente et al., 2002). However, diffusive loss of K^+ and

335   other leached cations (e.g., Ca^{2+}, Mg^{2+}, Fe^{2+}) was low due to (1) the low porosity and permeability

336   of the mudstones and marls enclosing the tuff layers at Xiakou, and (2) the mineralogy of the host

337   marine sediments, which consists dominantly of illite and carbonates, yielding a porewater cation

338   assemblage similar to that of the tuffs. These factors led to retention of leached cations including

339   K^+ in the porewater of the tuffs and, thus, ultimately promoted the precipitation of jarosite.

340   The host sediment at Xiakou contains substantial Fe in the form of ankerite precipitated from

341   ferruginous seawater under slightly reducing conditions (Table 2), with incorporation of Fe^{2+} into

342   ankerite as a substituent for Mg^{2+}. However, iron in silicates tends to be released as Fe^{3+} during

343   alteration of volcanic materials, and the formation of jarosite is most likely attributed to very local

344   changes in chemistry, redox conditions, and element concentrations in the traps of sediments. As

345   confirmed by jarosite formation in a silicic chalk elsewhere, weathering of dispersed pyrite and

346   glauconite in the calcareous material led to the local precipitation of jarosite (Bauer and Velde,

347   1997). The absence of Fe(III)-oxide phases in the Xiakou tuffs is not only attributed to the

348   relatively Fe-poor intermediate-felsic source rocks but also to good preservation of jarosite in the

349   low-pH environment, since decomposition of jarosite often produces iron oxides (Elwood Madden

350   et al., 2012). Contrarily, weathering of Fe-rich basaltic rocks under cold, dry climatic conditions in

351   Greenland formed both nontronite and well-crystalline iron oxides (goethite and

352   magnetite/maghemite) in closed crack microenvironments due to water availability (Koch et al.,

353   1995). The formation of jarosite in the Xiakou section is thus likely attributable to highly localized

354   variation in porewater chemistry, pH and redox conditions.

355   The Xiakou section contains notably more sulfur than in P-T sequences elsewhere in South

356   China despite lack of evidence for the presence of sulfide deposits at or near the Xiakou area.

357   Total sulfur values are mostly 0.08-2.50% (versus 0.01-0.03% at Dongpan and 0.00-0.05% at

358   Xinmin), suggesting locally higher sulfur inputs to Xiakou (Kump et al., 2005; Pei et al., 2020).

359   Furthermore, all tuff beds yield higher sulfur contents than their host rocks (Shen, 2014),

360   indicating that sulfur uptake at Xiakou was associated with volcanic ash deposition. The

361   observation that jarosite is found only in tuffs at Xiakou and not in cogenetic volcanic ash beds in

362   other South China P-T boundary sections—despite all of the tuffs having a shared origin in

363   intermediate to felsic subduction-zone volcanism of the eastern Tethys (Shen et al., 2013; Yin et

364   al., 1992)—is an indication that excess sulfur was delivered through upwelling of a H_2S-rich,

365   euxinic deep watermass. The presence of this H_2S-rich watermass allowed sulfates including

366   jarosite to form in freshly deposited ash beds on the Xiakou seafloor. The absence of jarosite and

367   heavy $\delta^{34}S$ values of tuffs in the upper Xiakou profile (e.g., L259b and L271) are attributable to a

368   reduction in upwelling intensity (Fig. 3; Table 2). The presence of jarosite in background

369   sediments (mudstones) adjacent to tuff bed L255 can be interpreted as the product of mixing with

370   volcanogenic ash, as reflected by the presence of I/S clays in the mudstones (Table 1).

371   In certain ancient and recent marine sediments, the formation of gypsum and bassanite can

372   involve mainly alteration of volcanogenic materials with no sulfur input, especially in those

373   having context of convergent margins or basaltic plateaus (Hoareau et al., 2011), since volcanism

374   in this tectonic environment often links to voluminous outgassing and input of volcanic sulfate

375   aerosols (McKenzie et al., 2016). However, formation of jarosite in various magmatic tuffs has not

376   been observed except for limited cases in altered volcanogenic rocks around the volcanic vents

377   due to sulfur-bearing outgassing (Zimbelman et al., 2005). This observation suggests that jarosite

378   is unlikely to form from weathering of tuffs without the flux of sulfur during the alteration process.

379   At Xiakou, it is clear that flux of H_2S upwelling in association with subaqueous weathering of

380   volcanogenic materials is the major process leading to saturation and precipitation of jarosite in

381   sediment porewaters. The preservation of jarosite in the Xiakou sediments is most likely attributed

382   to the acidic brine porewater, owing to the elevated salinity, anoxic, and euxinic depositional

383   environment (Pei et al., 2020), the acidic condition due to lasting H_2S flux, and subsequently the

384   decreasing water activity by elevated temperature and enhanced compaction in burial diagenesis.

385   Devitrification of volcanic ash yields mainly smectitic clay minerals during early diagenesis (40°

386   to 100 °C; Fortey et al., 1996), which will transform into I/S clays and sometimes almost pure

387   illite in burial diagenesis due to increasing temperature (Huff, 2016). The weak activity of fluid

388   flow during burial diagenesis also contributes to the stability of jarosite, as reflected by the

389   presence of mainly I/S clays and absence of diagenetic mineral phases in the tuff beds. To date,

390   jarosite formation observed in the Xiakou tuffs within the P-T strata confirms that extremely

391   acidic conditions linked to upwelling of H_2S formed by sulfate-reducing bacteria (SRB),

392   indicative of the microbe-induced formation of jarosite during weathering of volcanogenic

393   materials in water.

394

395   **IMPLICATIONS**

396   Jarosite in the Xiakou P-T section exhibits striking mineralogical and morphological

397   similarities to occurrences of jarosite in Meridiani Planum sediments, suggesting that the former is

398   a possible analogue for Martian sulfate-bearing settings. The essential conditions for jarosite

399   formation at Xiakou are concurrent deposition of volcanic airfall materials, a flux of oxygen

400   (co-transported with the ash), and upwelling of H_2S from a deep euxinic watermass. Although the

401   felsic-intermediate source rocks at Xiakou differ from the primarily basaltic source rocks on Mars,

402   the implications for sulfate formation during subaqueous weathering of volcanogenic materials

403   concurrent with H_2S influx are likely similar. Smectite-bearing Noachian and Hesperian-aged

404   sedimentary rocks on Mars are analogous to altered volcanogenic materials in an open water body

405   on Earth, since the weathering products of volcanic materials under such conditions tend to be

406   monomineralic (Jeans et al., 2000). The Xiakou depositional environment is analogous to

407   smectite-bearing ancient lake basins on Mars such as the Gale and Jezero craters. The limited

408   illitisation of smectite on Mars was likely the result of low burial temperatures and a lack of K^+ in

409   diagenetic fluids (Borlina et al., 2015; Bristow et al., 2018), analogous to shallow lacustrine

410   settings in Turkey in which thick bentonites have accumulated (Özdamar et al., 2014; Kadír et al.,

411   2017).

412   Calcium sulfate is present as late diagenetic veins at many Mars rover landing sites, but it is

413   also likely present in the sedimentary matrix (Vaniman et al., 2018). As in the Xiakou tuffs, some

414   gypsum in smectite-bearing Martian mudstones can be explained as a non-evaporitic mineral,

415   formed in a closed-basin environment through weathering of volcanogenic materials with limited

416   influx of sulfur. Bassanite and anhydrite can further form from dehydration of gypsum during

417   subsequent burial diagenesis due to increasing porewater temperature and salinity (Hoareau et al.,

418   2011). In particular, sulfates including jarosite in the mudstones of the Murray formation in Gale

419   crater occur as dispersed crystals in the fine-grained matrix or as cement in nodules, dendrites, and

420   veins (Hurowitz et al., 2017), consistent with precipitation from porewaters during burial

421   diagenesis (Kah et al., 2018). Moreover, the relatively young age of jarosite in a single target from

422   the Murray formation indicates a late-diagenetic or a weathering product origin for Martian

423   jarosite (Martin et al., 2017). This mineral assemblage is analogous to that in the Xiakou tuffs.

424   The presence of sulfates (including jarosite) in the Murray formation mudstones has been

425   attributed to an influx of sulfuric acid, likely related to volcanic outgassing (Bibring et al., 2006),

426   or to acidic groundwaters during diagenesis (Rampe et al., 2017). However, an alternative

427   hypothesis for their origin is subaqueous weathering of volcanogenic materials in conjunction with

428   upwelling of H_2S from below a euxinic chemocline in an open water body on the surface of Mars.

429   Although these volcanogenic materials may have contained highly soluble volcanic sulfate

430   aerosols, volcanic-induced acidification likely took place in a relatively short interval of time such

431   that it did not cause widespread acid-sulfate alteration (Cronin et al., 2014; Kump et al., 2005).

432   The extremely acidic environment could be produced due to intensive H_2S upwelling in

433   association with the disturbance of deposition of volcanogenic materials in the basin, analogous to

434   the jarosite formed in the Xiakou tuffs. The accumulation of H_2S in the background sediments of

435   the basin could be possibly derived from sulfate-reducing bacteria, especially because SRB can do

436   this faster - on biological time scales rather than geologic ones, and the upwelling of H_2S-bearing

437   fluids may be induced due to heating from magmatic activity underlying the crater or due to

438   dewatering of smectite during early burial diagenesis. Therefore, jarosite-bearing sediments on

439   Mars are potentially evidence of the former existence of microbial life on the Red Planet, which is

440   consistent with previously reported cryptic mineralogic and organic geochemical evidence of

441   microbial or nanobial life in Martian meteorites (Stern et al., 2022). Furthermore, bacterial sulfate

442   reduction could explain the highly ^{34}S-depleted sulfur isotope signatures (locally to $-47 \pm 14‰$) of

443   sedimentary rocks from Gale crater (Franz et al., 2017; Stern et al., 2022).

444

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636 **Supplementary Materials**

637   Fig S1 to S2

638   Table S1 to S4

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Table 1 Relative proportions of clay species of clay fraction from the Xiakou sediments (wt%)

Sample	Lithology	Illite	Chlorite	Kaolinite	I/S clays
L278	Mudstone	94±3	6±1	/	/
L277	Tuff	/	/	/	100
L276	Mudstone	84±2	16±2	/	/
L272	Marl	73±3	28±2	/	/
L271	Tuff	/	/	6±1	94±7
L270	Marl	97±4	3±1	/	/
L268	Mudstone	97±5	3±1	/	/
L267b	Tuff	40±3	7±1	/	53±6
L267a	Mudstone	90±3	10±2	/	/
L266a	Mudstone	98±3	2±1	/	/
L266	Tuff	/	/	/	100
L265b	Mudstone	96±3	4±1	/	/
L265a	Mudstone	95±6	5±1	/	/
L264	Tuff	3±1	/	/	97±5
L263	Mudstone	98±3	2±1	/	/
L261a	Marl	97±5	3±1	/	/
L261	Mudstone	76±3	11±2	/	13±2
L260	Tuff	/	/	/	100
L259b	Tuff	/	/	/	100
L259a	Marl	75±2	5±1	/	20±3
L256	Mudstone	69±3	3±1	/	28±2
L255	Tuff	/	/	/	100
L254	Mudstone	76±2	9±1	/	15±2

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Table 2 Mineral and sulfur isotopic compositions of the Xiakou sediments

Sample	Lithology	Dolomite	Ankerite	Calcite	Pyrite	Jarosite	Bassanite	Gypsum	Quartz	Plagioclase	K-feldspar	Clay minerals	K_2O	$^{34}S_{CaSO4}$	$^{34}S_{pyrite}$
L278	Mudstone	2	16	1	2	/	/	/	17	2	/	60	2.69	/	−15.96
L277	Tuff	/	/	/	/	1	1	1	1	/	/	96	6.07	−6.62±1.36	/
L276	Mudstone	3	16	5	2	/	/	/	20	3	/	51	2.76	/	+2.07
L272	Mudstone	1	39	7	4	/	/	/	18	/	/	31	1.30	/	+2.87
L271	Tuff	/	/	1	/	/	/	/	1	/	/	98	6.18	/	−15.28
L270	Mudstone	1	6	46	/	/	/	/	10	2	/	35	0.89	/	−6.59
L269	Marl	1	13	3	2	1	/	/	24	8	/	48	3.67	/	−29.89
L267b	Tuff	/	/	/	1	3	4	/	26	7	/	59	4.49	+3.20±1.23	−29.01
L267a	Mudstone	2	20	1	4	/	/	/	17	3	/	53	2.73	/	−39.85
L266a	Mudstone	2	/	/	/	/	/	/	24	4	/	70	0.95	/	+33.74
L266	Tuff	/	/	/	1	4	/	1	1	1	/	92	5.38	−37.23 ±0.71	−25.96
L265b	Mudstone	/	21	/	3	/	/	1	20	2	/	53	4.33	−36.22±1.54	−25.73
L265a	Mudstone	/	6	/	2	/	/	/	19	8	/	65	3.75	/	−34.20
L264	Tuff	/	/	/	/	6	/	1	5	/	/	88	5.32	−33.44±2.28	−34.95
L263	Mudstone	1	34	/	3	/	/	1	19	3	/	39	1.95	−35.18±0.83	−26.62
L261a	Mudstone	3	34	2	1	/	/	2	12	1	/	45	2.63	−28.36±2.40	+1.53
L261	Mudstone	1	24	1	5	/	/	/	19	3	/	47	1.00	/	+5.57
L260	Tuff	/	1	/	/	3	/	2	1	/	/	93	5.76	−29.60 ±1.67	−4.21
L259b	Tuff	/	2	3	2	/	2	/	14	2	/	75	4.91	−28.90±1.42	/
L259a	Marl	2	23	10	3	/	/	/	28	/	1	33	1.72	/	−6.29
L256	Mudstone	/	/	/	/	5	/	2	9	2	/	82	4.85	−34.46±1.29	−23.19
L255	Tuff	/	/	1	/	2	1	/	3	3	/	90	5.57	−29.86±3.94	/
L254	Mudstone	/	/	19	/	4	/	8	26	2	8	33	1.95	−21.80±2.78	−23.73

653 Note: Mineral and sulfur contents in wt% and δ^{34}S values in ‰, the relative error of mineral contents is <15%; δ^{34}S$_{pyrite}$ from Shen et al. (2013); and "/" is under detection limit.

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655   Table 3 The zoned distribution of δ^{34}S value of gypsum showing the evolution of porewater with diagenesis

Sample	Lithology	Jarosite	Bassanite	Gypsum	K$_2$O	^{34}S$_{CaSO4}$(‰)	Zoned grains (center to margin); ^{34}S$_{CaSO4}$ (‰) (errors in accuracy < 0.7‰)
L277	Tuff	0.66	0.87	0.31	6.07	-6.62±1.36	-7.88, -6.34, -4.16, -3.28, -3.55, -3.71, -0.18, -0.39
L266	Tuff	3.93	/	0.41	5.38	-37.23±0.71	-37.80, -35.74
L261a	Mudstone	/	/	2.40	2.63	–28.36±2.40	-24.24, -28.19
L255	Tuff	2.46	0.90	/	5.57	-29.86±3.94	-30.74, -24.88, -19.38 (Grain #1) -28.94, 20.56 (Grain #2)
L254	Mudstone	4.32	/	7.79	1.95	-21.80±2.78	-15.12, -20.11

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