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2 **Raman Spectroscopy of the Ilmenite – Geikielite Solid Solution**

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10 *Revision 1*

11 **Abstract**

12 Ilmenite ($\text{Fe}^{2+}\text{TiO}_3$) and geikielite (MgTiO_3) are important terrestrial minerals relevant to
13 the geology of the Earth, the Moon, Mars, and meteorite samples. Raman spectroscopy is a
14 powerful technique that allows for mineral cation determination for the ilmenite – geikielite solid
15 solution. We report on a sample suite of nine samples within the ilmenite – geikielite solid
16 solution and provide context for their quantitative interpretation. We compare a univariate
17 Raman peak position model for predicting ilmenite composition with a multivariate machine
18 learning model. The univariate model is currently recommended, though the multivariate model
19 may become superior if the data set size is increased. This study lays the groundwork for
20 quantifying Fe (ilmenite) and Mg (geikielite) within oxides minerals using a cheap, portable, and
21 efficient technology like Raman spectroscopy.

22 **Key Words:** ilmenite, geikielite, Raman spectroscopy

23 **Introduction**

24 Ilmenite is an important mineral group on planetary surfaces, especially for Earth, the
25 Moon (e.g., Papike et al., 1976; Papike et al., 1991; Lemelin et al., 2013; Surkov et al., 2020),
26 and Mars (e.g., Morris et al., 2006). This group of minerals is also found in a variety of meteorite
27 samples (e.g., Snetsinger and Keil, 1969; Bunch and Keil, 1971). Ilmenite group minerals occur
28 in all rock types, making their study broadly relevant. They are mined on Earth as an important
29 resource material. Most terrestrial occurrences are Fe-rich with the exception of kimberlites,
30 where Mg substitutions exist, acting as a kimberlite indicator (e.g., Wyatt et al., 2004). Lunar
31 ilmenite can have Mg substitutions along the ilmenite – geikielite solid solution (e.g., Papike et
32 al., 1991; Tokle and Robertson, 2019; Robertson et al., 2022). Raman spectroscopy may act as an
33 important tool in the upcoming decades for lunar exploration (e.g., Cloutis et al., 2022) for
34 mineral identification and quantification. Notably, ilmenite group minerals provide information
35 about lunar magmatic evolution (e.g., Sato et al., 2017) that affect interpretations of the Moon's
36 interior. The presence of ilmenite in the SNC (shergottites, nakhlites, chassignites) meteorites is
37 of note in regard to Mars (McSween, 1994; Rull et al., 2004; Wang et al., 2004). Ilmenite group
38 minerals contain valuable compositional information relevant to oxygen fugacity and mineral
39 stability (e.g., see Szymanski et al. (2010) for applicability to Mars). Overall, characterizing the
40 composition of ilmenite on Earth and remote planetary bodies informs geologic interpretations
41 and resource identification (e.g., Heiken and Vaniman, 1990).

42 There is a solid solution between ilmenite ($\text{Fe}^{2+}\text{TiO}_3$) and geikielite (MgTiO_3), as well as
43 with other minerals like pyrophanite (MnTiO_3) at high temperatures. Here, Raman spectroscopy
44 is used to measure the composition of synthetic samples covering the solid solution between
45 ilmenite and geikielite. The cation ratio of Fe to Mg ($\% \text{ilmenite} = (100 \times \text{Fe}) / (\text{Mg} + \text{Fe})$) affects
46 positions of the Raman peaks, allowing for the prediction of the mineral compositions. This

47 investigation lays the groundwork for quantifying the cation ratio for the ilmenite – geikielite
48 solid solution using Raman spectroscopy.

49 **Background**

50 Ilmenite and geikielite have three acoustic modes and 27 optical modes. Of these, 10
51 major Raman spectroscopy bands for ilmenite group minerals occur in the common measurement
52 range from 200 to 800 cm^{-1} . These Raman features are caused by the combination of $5A_g$ and
53 $5E_g$ modes (Tibshirani, 1995; Okada et al., 2008). One of the A_g bands, located between ~ 681 –
54 715 cm^{-1} , is primarily utilized here.

55 This work builds on previous studies focused on X-ray diffraction and visible and near-
56 infrared calibrations for the ilmenite – geikielite solid solution using many of the same samples
57 (Tokle et al., 2018; Tokle and Robertson, 2019). Electron microprobe and X-ray diffraction are
58 important tools for distinguishing the mineralogy of the ilmenite – geikielite solid solution. Our
59 well-characterized synthetic mineral samples provide a unique opportunity to evaluate
60 compositional differences within the ilmenite mineral group using Raman spectroscopy, a more
61 accessible and less expensive tool.

62 Over the last 50 years, a wide variety of studies investigated ilmenite minerals using
63 Raman spectroscopy. Early work included Raman spectral measurements and identification of
64 peak positions (e.g., Beattie and Gilson, 1970; White, 1975; Pinet et al., 1986). Subsequent
65 studies investigate ilmenite properties using Raman spectroscopy like heat capacity and mineral
66 stability (McMillan and Ross, 1987; Chopelas, 1999; Linton and Navrotsky, 1999). Additional
67 research of ilmenite group minerals focused on high-pressure experiments (e.g., Reynard and
68 Guyot, 1994; Okada et al., 2008). Part of the motivation behind these studies is the structural
69 similarity of ilmenites and materials like MgSiO_3 –ilmenite that are relevant to Earth’s mantle.

70 Raman spectroscopy is useful for applied geologic investigations like geikielite exsolution in
71 spinel (Reusser et al., 2001) and ilmenite detection in Martian meteorites (Wang et al., 2004).
72 Wang et al. (2004) established a correlation between the Raman A_g peak positions ($\sim 681 - 715$
73 cm^{-1}) to %ilmenite content in a set of ilmenite – geikielite samples, providing a framework for
74 our investigation.

75 **Methods**

76 All nine synthetic ilmenite and geikielite powders utilized in this study were synthesized
77 from oxides in a 1 atmosphere CO:CO₂ furnace. Details on the mineral synthesis of these
78 samples are provided in Tokle and Robertson (2019). All powders were sieved to a grain size
79 fraction of 10 – 20 μm . Microprobe analysis shows impurities in all powders are < 1 weight%
80 and chemically homogeneous (Tokle et al., 2018). The electron microprobe analysis values are
81 provided in Table 1 of Tokle and Robertson (2019). The nine samples within the suite include
82 pure ilmenite and geikielite as well as samples compositionally between these end-members
83 (1%, 5%, 10%, 20%, 40%, 60%, and 80% ilmenite). Compared to the natural samples used by
84 Wang et al. (2004), the Mg-rich samples with narrow compositional gaps augment existing
85 Raman data.

86 Raman spectra were acquired on a Bruker BRAVO Raman dual laser (785 and 852 nm)
87 spectrometer with a spot size of 2 mm in diameter and fixed laser power that did not exceed 100
88 mW to reduce the risk of material alteration. Five sample scans and an integration time of 10
89 seconds were utilized at a spectral resolution of $2.0 \text{ cm}^{-1}/\text{channel}$.

90 Univariate data analysis included gaussian peak fitting for the diagnostic ilmenite feature
91 located between $\sim 681 - 715 \text{ cm}^{-1}$ (A_g). A linear model was used to fit the univariate Raman peak
92 position data when regressed against composition. Partial least squares (PLS) regression, a

93 multivariate machine learning model, was also used to predict %ilmenite content. The PLS
94 method regresses one response variable (%ilmenite) against multiple explanatory variables
95 (intensity at each channel of the spectra), assigning coefficients to every channel (Geladi and
96 Kowalski, 1986; Wold et al., 1983; Wold et al., 2001). Multivariate techniques can exploit broad
97 spectral ranges including multiple diagnostic Raman peaks and do not depend on any single
98 feature's position.

99 For the univariate and multivariate methods, R^2 values, internal root-mean-square error
100 (RMSE) values, and cross-validated RMSE (RMSE-CV) statistics are reported. RMSE-CV
101 values were calculated by creating three folds of data (square root of the total number of
102 samples) and building a prediction model with the remaining data, then averaging the resultant
103 errors. Here, the RMSE-CV value represent the most accurate prediction error associated with
104 data outside the models. Internal RMSE values allow for comparisons to the literature however,
105 these values underestimate the error of predicting data outside the model. All the RMSE values
106 are in units of %ilmenite.

107 **Results**

108 Seven of the previously reported (e.g., Linton and Navrotsky, 1999; Wang et al., 2004)
109 Raman peaks were directly observed (Figure 1) (Breitenfeld et al., 2023). The geikielite
110 spectrum (purple) has bands at roughly 306, 327, 352, 397, 485, 640, and 715 cm^{-1} (**Table 1**).
111 Raman features of pure ilmenite (yellow) shift to lower wavenumber positions as Fe increases,
112 although several bands are absent or poorly resolved.

113 Figure 2 depicts the univariate model for predicting %ilmenite using the diagnostic
114 Raman peak position between $\sim 681 - 715 \text{ cm}^{-1}$ (A_g) with error bars from the peak fitting. The

115 linear equation from the univariate model results in an R^2 value of 0.99 with internal RMSE and
116 RMSE-CV values equivalent to ± 3.9 and ± 11.0 %ilmenite, respectively.

117 Figure 3 shows the internal RMSE and RMSE-CV values for multivariate models with
118 varying numbers of PLS components. The lowest RMSE-CV value is associated with a four
119 component PLS model. However, the small spectral dataset size and large difference between the
120 internal RMSE and RMSE-CV values may indicate that these PLS multivariate models are
121 overfitting the small data set. Overall, all RMSE-CV values for the PLS models are larger
122 (worse) than those for the univariate Raman peak position method.

123 Discussion

124 Fewer Raman features are observed for the pure ilmenite sample compared to the pure
125 geikielite sample (**Table 1**). This is consistent with other Raman measurements of these minerals
126 (e.g., Wang et al., 2004). Equivalent to geikielite, 10 Raman modes (5 A_g and 5 E_g) are predicted
127 for ilmenite (Ross and McMillan, 1984). We are interested in understanding this difference in
128 Raman spectral expression further.

129 Raman spectroscopy allows for the quantification of Fe to Mg within the ilmenite –
130 geikielite solid solution. The reported univariate peak position method outperforms the
131 multivariate PLS method for the current data set size. The Raman peak position of the ilmenite
132 A_g feature between $\sim 681 - 715 \text{ cm}^{-1}$ should thus be used to predict the ilmenite and geikielite
133 content of unknown samples.

134 It must be cautioned that the reported errors (internal RMSE and RMSE-CV) apply only
135 to analyses of ilmenites under these analytical conditions and will likely not be directly
136 comparable to data from different Raman instruments. The effect of predicting the %ilmenite

137 content for natural samples with more diverse cations rather than the pure synthetic samples must
138 also be examined.

139 A larger sample suite should improve the multivariate model, as observed in other types
140 of spectroscopic investigations (e.g., Dyar and Ytsma, 2021). This work lays the groundwork for
141 an improved multivariate model that exploits valuable spectroscopic information beyond a single
142 Raman feature.

143 **Implications**

144 Raman spectroscopy is a useful tool for mineral identification and quantification. Sample
145 characterization can be performed through cation ratio determinations. This work will aid future
146 workers in answering specific geologic questions related to the ilmenite – geikielite solid
147 solution and relevant applications. For example, inquiries may be pursued related to ilmenite and
148 geikielite mineral associations, alteration conditions, and resource identifications.

149 Ilmenite is an important mineral group for many planetary bodies including Earth, the
150 Moon, and Mars. This work is particularly relevant to the characterization of terrestrial
151 kimberlites and lunar basalts. In addition to laboratory analyses of terrestrial and extraterrestrial
152 samples, this work is also applicable to handheld Raman spectroscopy measurements. This can
153 be particularly useful for real-time terrestrial field work or planetary surface exploration by
154 astronauts.

155 **Acknowledgments**

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157 you to Professor Steven Jaret and the other reviewer for their insightful feedback that improved
158 the quality of this work.

159 **Data Availability**

160 The Raman spectral data are archived in an external data repository at
161 <https://zenodo.org/records/10210991> (Breitenfeld et al., 2023).

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251

List of Figure Captions

252 **Figure 1.** Ilmenite and geikielite Raman spectra acquired on Bruker's BRAVO spectrometer.

253 Spectra are color-coded based on %ilmenite content, where pure $\text{Fe}^{2+}\text{TiO}_3$ is represented with

254 yellow and pure MgTiO_3 with purple.

255

256 **Figure 2.** Univariate model of %ilmenite versus Raman peak position (cm^{-1}) of the diagnostic

257 Raman feature between $\sim 681 - 715 \text{ cm}^{-1}$ (A_g) with error bars of the gaussian peak fit. The linear

258 model, R^2 value, internal RMSE, and RMSE-CV are reported.

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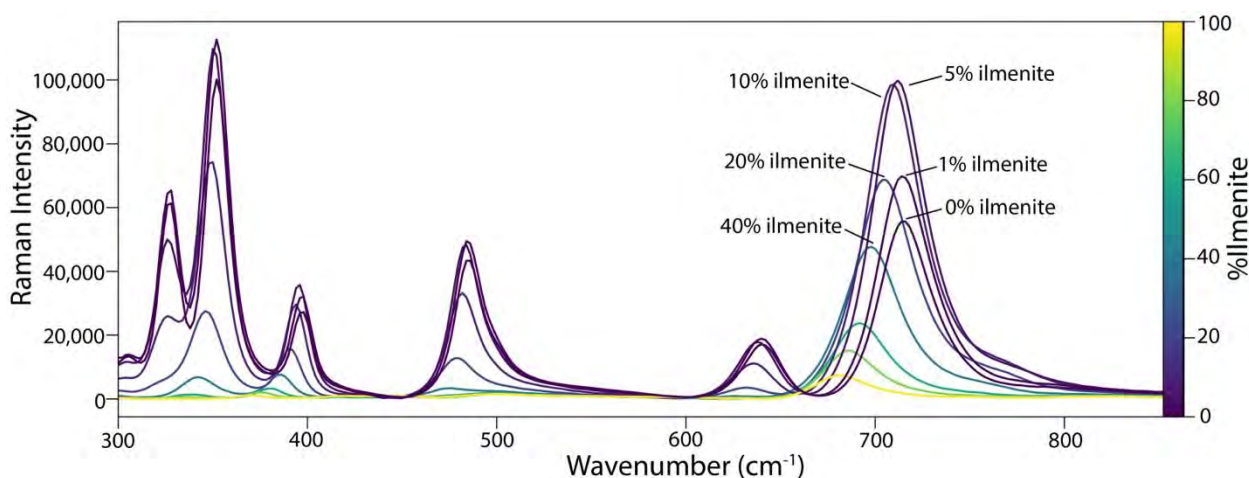
260 **Figure 3.** Multivariate model errors (internal RMSE and RMSE-CV) for 2 – 6 PLS model

261 components.

262

Figures and Tables

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264

265 **Figure 1.** Ilmenite and geikielite Raman spectra acquired on Bruker's BRAVO spectrometer.

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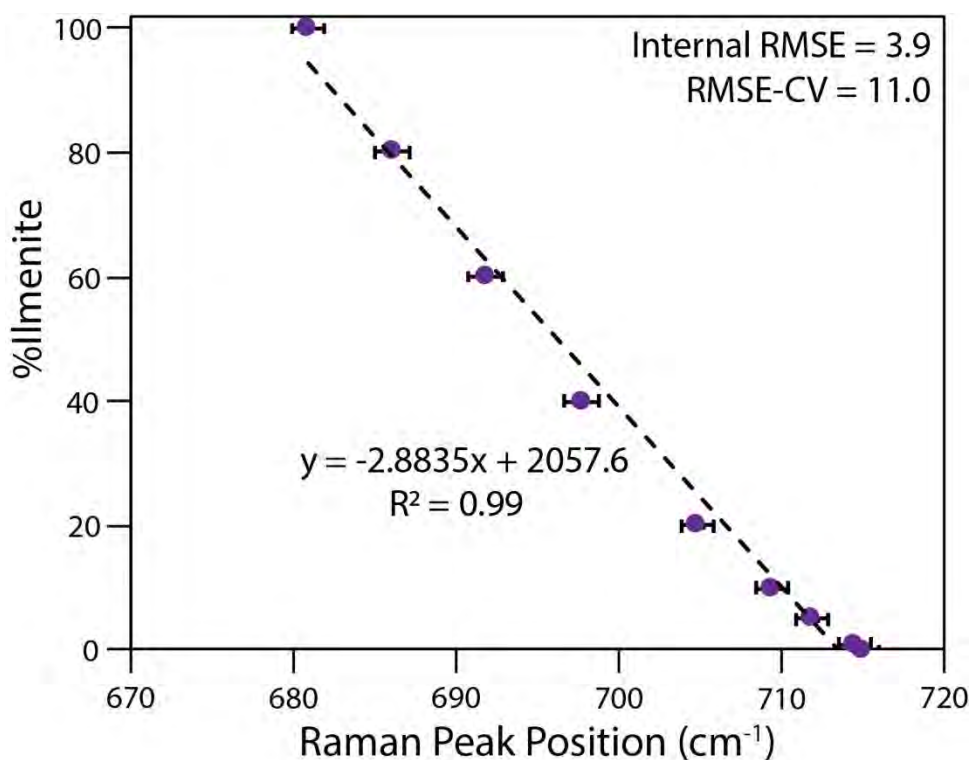
Table 1. Raman peak positions and modes of geikielite and ilmenite for our dataset compared to the literature.

Raman Peak Positions (cm ⁻¹)				Raman Modes	
geikielite		ilmenite		Wang et al., 2004; Okada et al., 2008	
this study	Linton and Navrotsky, 1999	this study	Wang et al., 2004		
306	306	-	333	E _g	translation
327	327	370	368	A _g	translation
352	352	-	-	A _g	bending
397	397	-	-	E _g	bending
485	485	-	-	E _g	bending
-	502	-	-	A _g	bending
640	641	-	-	E _g	stretching
715	714	681	683	A _g	stretching

Note: Two additional Raman modes are predicted and observed outside the wavenumber range of the Raman instrument utilized in this investigation.

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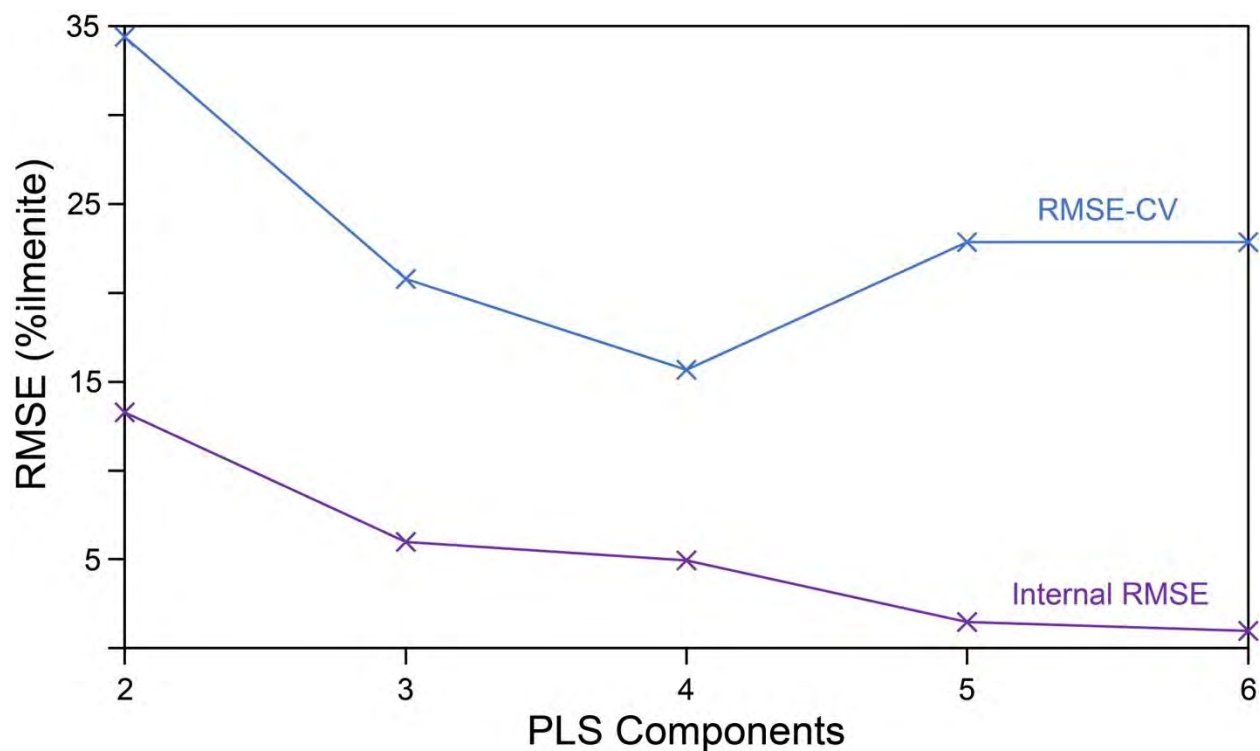
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Figure 2. Univariate model of %ilmenite versus Raman peak position (cm⁻¹) of the diagnostic Raman feature between ~681 – 715 cm⁻¹ (A_g) with error bars of the gaussian peak fit. The linear model, R² value, internal RMSE, and RMSE-CV are reported.



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Figure 3. Multivariate model errors (internal RMSE and RMSE-CV) for 2 – 6 PLS model components.