## A micro-reflectance IR spectroscopy method for analyzing volatile species in basaltic, andesitic, phonolitic, and rhyolitic glasses

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### ABSTRACT

Volatile contents of geologic glasses are used to model magma chamber and degassing processes, thus, there is considerable interest in small-scale analytical techniques for analyzing volatiles in glasses. Infrared (IR) spectroscopy has the advantage of determining volatile speciation in glasses (e.g.,  $OH^-$ , molecular  $H_2O$ , molecular  $CO_2$ , and  $CO_2^{2-}$ ). However, sample preparation for the most common IR method used, micro-transmission IR spectroscopy, is complicated because glasses must be prepared as thin, parallel-sided wafers. Raman analysis, while valuable for Fe-poor samples, can be difficult to use for Fe-rich glasses.

We have calibrated a micro-reflectance infrared method for determining volatile species using calculated Kramers-Kronig absorbance (KK-Abs.) spectra that requires that only one side of a glass be polished. The method is easier to use than other reflectance methods where it is difficult to determine the baseline for the IR bands. Total H<sub>2</sub>O wt% =  $m \cdot (3600 \text{ cm}^{-1} \text{ KK-Abs.})$ , where *m*, is the slope of the calibration line that is obtained from a fit to the data. The *m* value is related to the calculated refractive index, *n*, for a range of aluminosilicate glass compositions allowing the technique to be applied to samples with unknown calibration slopes. For calc-alkaline andesite glasses we determined calibration slopes for micro-reflectance IR measurements of molecular H<sub>2</sub>O, molecular CO<sub>2</sub>, and CO<sub>3</sub><sup>-</sup>. The method has been calibrated for glasses with up to 6.76 wt% total H<sub>2</sub>O (but is useful for glasses with more than 20 wt% total H<sub>2</sub>O) and has been calibrated for glasses with up to 0.575 wt% total CO<sub>2</sub>.

This technique provides a means to analyze volatile abundances in samples that are not possible to analyze or prepare for analysis with transmission micro-IR techniques. We have determined volatile contents in fragile samples such as cracked, vesicular, or crystal-bearing glasses formed by volcanic or impact processes or in high-pressure bubble nucleation experiments and H diffusion experiments. We have monitored H uptake during weathering of basaltic glasses that cannot be polished and determined volatiles in melt inclusions and pumice.

Keywords: IR spectroscopy, glass properties, new technique, volatiles, H<sub>2</sub>O, CO<sub>2</sub>, CO<sub>3</sub><sup>-</sup>

#### INTRODUCTION

Volatiles, such as  $H_2O$  and  $CO_2$ , are found in natural glasses and their abundances are used to model the role of volatiles on magmatic processes and degassing (Wallace and Anderson 2000; Larsen and Gardner 2004; Métrich and Wallace 2008); to determine volatile diffusion rates in melts (e.g., Liu et al. 2004); and to measure volatile contents of impact glasses (Harris et al. 2007). Thus, there is considerable interest in small-scale analytical techniques for analyzing glasses. Various techniques are used, such as vacuum-line manometry, secondary ion mass spectrometry, Karl-Fischer titration (for  $H_2O$ ), Raman spectroscopy, and Fourier transform infrared (FTIR) spectroscopy (see reviews in Ihinger et al. 1994; Deloule et al. 1995; Devine et al. 1995; King et al. 2002; Behrens et al. 2006; Thomas et al. 2006; Aubaud et al. 2007). Of these techniques, only Raman

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and FTIR spectroscopy provide information on volatile species [e.g., OH<sup>-</sup>, molecular H<sub>2</sub>O (H<sub>2</sub>O<sub>mol</sub>), molecular CO<sub>2</sub> (CO<sub>2 mol</sub>), and CO<sub>3</sub><sup>2-</sup>] in small sample areas (<50  $\mu$ m across). The abundances of these species are also used in volcanic "geospeedometers" (e.g., Zhang et al. 2000) and it is therefore advantageous to measure them accurately.

Micro-Raman spectroscopy has been developed to determine the total  $H_2O(H_2O_{tot})$ ,  $H_2O_{mol}$ , and  $OH^-$  in silicic, Fe-poor glasses (Thomas 2000; Chabiron et al. 2004; Behrens et al. 2006; Thomas et al. 2006) and a few Fe-bearing glasses (Chabiron et al. 2004; Zajacz et al. 2005; Behrens et al. 2006; di Muro et al. 2006; Thomas et al. 2008a; Mercier et al. 2009). Raman methods use either internal calibrations (e.g., Behrens et al. 2006) or a comparator method where samples are referenced to one calibration glass (Thomas et al. 2008b). However, there are difficulties in applying the micro-Raman method to Fe-bearing glasses due to band overlap in the 100–1100 cm<sup>-1</sup> region and absorption of the

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laser radiation with depth in Fe-bearing glasses that decreases the OH band intensity (Behrens et al. 2006; Thomas et al. 2008a). In addition, glasses with low-melting temperatures, including water-rich samples, may lose volatiles or melt due to strong light absorption and/or fluorescence radiation.

FTIR spectroscopy has been used to successfully analyze Ferich compositions like basaltic glasses (e.g., Dixon et al. 1988; Dixon and Pan 1995). However, the most common method used, micro-transmission IR spectroscopy, is complicated because glasses are generally prepared as thin, parallel-sided, stand-alone wafers (e.g., Paterson 1982; Stolper 1982a, 1982b; Newman et al. 1986; Dixon et al. 1988; Ihinger et al. 1994; Dixon and Pan 1995). Transmission IR techniques using ground glass powders pressed into a KBr disk have the advantage that painstaking thickness measurements on doubly polished glass wafers are not required; however KBr is hygroscopic and therefore it is challenging to ensure that water is not adsorbed (e.g., Izawa et al. 2010). Furthermore, the KBr method is not appropriate for samples in which spatially constrained, fine-scale spectra are needed such as melt inclusions and some experimental glasses. A technique has been developed to investigate melt inclusions in doubly polished olivine crystals (Nichols and Wysoczanski 2007); however that technique is not amenable to studying very small melt inclusions and some volatile species are not possible to measure (e.g., H<sub>2</sub>O<sub>mol</sub>) due to band interferences from the host mineral. An alternative technique is micro-reflectance IR analysis (e.g., Grzechnik et al. 1996; Moore et al. 2000; Hervig et al. 2003), where only one sample face is polished.

The major advantage of the micro-reflectance method is that the sample can be prepared as a single polished surface. Therefore, small samples, such as melt inclusions, do not need to be removed from their matrix for mounting, and fragile pumices and experimental charges may be analyzed in textural context. Also, different depths within a single sample can be analyzed sequentially by successively polishing the sample to deeper levels.

Previous work has shown that reflectance spectra for C-O-H volatiles have band positions similar to that found in transmission IR spectra, but they have different shapes due to the mathematical nature of the superposition of the volatile bands on the Si-O band. Example micro-reflectance IR spectra are shown in Figures 1a and 1b.

For O-H species in glasses, the largest band occurs as a "negative band" in the 3500 cm<sup>-1</sup> region (Hervig et al. 2003). This band, referred to here as a "total H<sub>2</sub>O" band, arises from a combination of symmetric stretching of the H-O-H molecule ( $v_{ss}$ H-O-H), plus symmetric stretching of Si-OH ( $v_{ss}$ Si-OH) in the aluminosilicate network, plus an overtone of the H-O-H bending vibrations (2 $\delta$ H-O-H). The asymmetric stretching vibration ( $v_{as}$ ) of CO<sub>2 mol</sub> gives rise to a "sigmoidal" feature in the reflectivity spectrum (Moore et al. 2000). The dissolved CO<sub>3</sub><sup>--</sup> concentrations (vC-O) in basanite and leucitite glasses have a "doublet" band in micro-reflectance IR spectra that resembles the absorption bands obtained in transmission IR studies (Grzechnik et al. 1996). In this study, we show that a "negative band" for molecular H<sub>2</sub>O



**FIGURE 1. (a)** Reflectance spectrum for calc-alkaline andesite, MHA-23, with insets showing the total  $H_2O$  ( $H_2O_{tot}$ ), molecular  $CO_2$  ( $CO_2 \text{ mol}$ ), and  $CO_3^2$  bands. Note that it is difficult to evaluate where a baseline should be drawn in the case of the C-O species bands. (**b**) Reflectance spectrum for calc-alkaline andesite MHA-41 showing the sigmoidal molecular  $H_2O$  ( $H_2O_{mol}$ ) band at ~1620 cm<sup>-1</sup>. Note that it is difficult to evaluate where a baseline should be drawn in the case of the sigmoidal band. (**c**) KK-Abs. spectrum for MHA-23 with insets showing  $H_2O_{tot}$ ,  $CO_{2 mol}$ , and  $CO_3^2$  bands and the baselines used for quantification. (**d**) KK-Abs. spectrum for MHA-41 showing the  $H_2O_{mol}$  band at ~1620 cm<sup>-1</sup> and the baseline used for quantification. The volatile contents for both of these samples are given in Table 2.

related to H-O-H bending vibrations ( $\delta$ HOH) occurs in the ~1620 cm<sup>-1</sup> region, in agreement with the band observed in transmission IR spectra.

In micro-reflectance IR measurements,  $H_2O_{tot}$  and total  $CO_2$  ( $CO_{2 tot}$ ) concentrations have been related directly to the height of the reflectance band for the species of interest ( $\Delta R_i$ ) relative to the reflectance at a reference wavenumber position ( $R_{ref}$ ) via an empirical equation of the form:

$$(\Delta R_i/R_{ref}) = k_1 \cdot (\text{volatile species}) + k_2$$
(1)

where the positions of the  $\Delta R_i$  and  $R_{ref}$  and the constants  $k_1$  and  $k_2$  depend on the glass composition (see King et al. 2004).

Despite, the ease of micro-reflectance IR measurements, the extraction of band heights from raw R% spectra may be somewhat challenging (e.g., Figs. 1a and 1b). There are significant uncertainties associated with the baseline position, especially for samples with low volatile contents and for sigmoidal bands like  $CO_{2 \text{ mol}}$  and  $H_2O_{mol}$  (Figs. 1a and 1b). Also, assumptions are needed regarding the location of  $R_{ref}$  for different composition glasses. Furthermore, calibrations are not available for  $H_2O_{tot}$  or  $H_2O_{mol}$  in basalt, or C-O species or  $H_2O_{mol}$  in andesite.

In this paper, we present a method for analyzing glasses using micro-reflectance IR spectroscopy that has been manipulated using a Kramers-Kronig (KK) transform. The technique is advantageous because it is straightforward to prepare samples and easier to determine the baseline position for IR bands than using the raw R% spectra (Figs. 1c and 1d). We provide new calibration curves for basalt, calc-alkaline andesite, alkali andesite, phonolite, and rhyolite. We also examine some of the limitations of this technique.

#### **METHODS**

Volatile-bearing glasses were obtained (Perfit et al. 1983, 1999; Westrich 1987; Moore and Carmichael 1998; King et al. 2002; Sinton et al. 2003; Larsen and Gardner 2004; Lui 2005; Mongrain et al. 2008) and their compositions were confirmed using the electron microprobe if necessary (Table 1). The basaltic glasses were analyzed using transmission IR spectroscopy and the  $H_2O_{tot}$  concentrations determined with the Beer-Lambert law (e.g., Stolper 1982b):

nolar concentration of 
$$H_2O_{tot} = A/(d \cdot \rho \cdot \epsilon_{H_2O_{tot}})$$
 (2)

where A is the linear absorbance (i.e., band height); d is the sample thickness that was measured with a Mituyomo micrometer to  $\pm 2 \mu m$  (<2% relative error for most samples);  $\rho$  is the sample density that was calculated following methods in Silver (1988) and data from Church and Johnson (1980) and Mandarino (1976); and  $\epsilon_{H_2O_{tot}}$  is the extinction coefficient for H<sub>2</sub>O<sub>tet</sub> in basalt (from Dixon et al. 1995). Calc-alkaline andesite glasses contained both C-O species as well as H-O species. Table 1 shows that variations in volatile-free compositions are given because slightly different compositions constitute the sample set. Errors in the volatile contents were taken from the literature or for the basalts and phonolite conservatively estimated using curve fitting uncertainties (10 to 14% relative) (Table 1).

Micro-reflectance IR analyses of the glasses were collected using a Nicolet Nexus 670 spectrometer with a Globar source, extended range KBr beamsplitter (XT-KBr), and a Continuµm microscope with a MCT-A detector (cooled with liquid nitrogen). The same instrument was used in King's lab at both the Experimental Analysis Laboratory at the University of Western Ontario and the New Mexico IR Analysis Laboratory at the University of New Mexico. Analyses were made in triplicate if possible, but some glass samples were too small for multiple analyses.

As with any IR analysis of C-O-H species, it is necessary to minimize the background levels of atmospheric gases surrounding the sample. All analyses in this work were performed by running a dry air purge into a volume around the

**TABLE 1.** Compositions of glasses used for calibrations in King's lab

	Basalt	Calc-alk.	Alkaline	Phonolite	Rhyolite			
		andesite	andesite					
wt%, normalized to 100%								
SiO <sub>2</sub>	53.69 (2.69)	60.62	63.68	57.31	75.58 (1.52)			
TiO <sub>2</sub>	1.17 (0.95)	0.89	0.64	0.24	0.17 (0.13)			
$AI_2O_3$	14.29 (2.17)	18.57	17.60	23.12	13.57 (0.83)			
FeO	10.71 (2.77)	5.56	3.89	1.63	1.31 (0.39)			
MgO	6.51 (3.00)	2.71	2.70	0.10	0.18 (0.12)			
CaO	10.82 (2.18)	6.05	5.74	0.66	0.88 (0.43)			
Na <sub>2</sub> O	2.25 (0.58)	4.67	4.12	11.22	4.36 (0.09)			
K₂O	0.56 (0.47)	0.92	1.64	5.72	3.93 (0.30)			
Referen	ce 1	2	3	4	5			

Notes: References are 1 = Galapagos basalt 1652-5 (Perfit et al. 1983, 1999), Manus Basin basalts 15-4, 2249, and 18-3 (Sinton et al. 2003) and experimental basalt DL414 (Lui 2005). 2 = Mount Hood andesite (King et al. 2002). 3 = Mascota andesite (Moore and Carmichael 1998). 4 = Laacher See Tephra phonolite (Larsen and Gardner 2004). 5 = PCD and M3M rhyolites (Westrich 1987).

sample that was ~100 cm<sup>3</sup> (e.g., Fig. 47 in King et al. 2004). We also monitored the total reflectance signal by examining the interferogram peak-to-peak measurement before each analysis. We prefer to operate at conditions where the peak-to-peak measurement is greater than 6.0 volts and we did not operate at conditions less than 4.0 volts.

Each spectrum was collected in a spectral range of  $5300-650 \text{ cm}^{-1}$ , over 300 scans, with 4 cm<sup>-1</sup> resolution, using a  $100 \times 100 \,\mu\text{m}$  sampling area. For each analysis, a background spectrum was collected on a gold-coated glass slide because it has a reflection coefficient of ~100% over the wavelength region measured, and so the absolute value of the sample reflectivity is known without further calibration. The raw spectra were collected in units of percent reflectance (%R). For our purposes, 300 scans appeared to provide sufficiently high signal-to-noise ratio (SNR); however, if a small aperture IR beam is used (e.g., for small melt inclusions), it is necessary to maximize the SNR by increasing the scan times to larger values (>100 scans, Grzechnik et al. 1996).

To convert the R% spectra to absorbance spectra we used the KK transform (McMillan and Hofmeister 1988), recently discussed in detail by Dufresne et al. (2009). The formula for the KK transform is:

$$\theta(v_i) = \frac{2v_i}{\pi} \int_0^\infty \frac{\ln r(v) - \ln r(v_i)}{v_i^2 - v^2} dv$$
(3)

where  $\theta$  is the phase shift, v is the real portion of the frequency, v<sub>i</sub> is imaginary portion of the frequency, and r is the real part of the reflectivity. In this study the reflectance spectra were transformed into KK absorbance (KK-Abs.) units by first smoothing the data points as a function of wavenumber over a ~40.5 cm<sup>-1</sup> window (like Moore et al. 2000) using a "running average type" algorithm (Savitsky-Golay). A smoothed spectrum minimizes error because the subsequent KK transform requires extrapolating the wings of the reflectance spectrum (Efimov 1995). The KK transform was performed using the OMNIC software (version 6.1) and to further ensure uniform extrapolation, we sampled over a constant range of 5300–650 cm<sup>-1</sup>.

#### RESULTS

Figures 1c and 1d show representative KK-Abs. spectra and illustrate the shapes of the major bands that are observed in this type of spectra (discussed above). Each spectrum was examined in the 700–1300 cm<sup>-1</sup> region to verify that only a broad glass band was present, ruling out a significant volume of crystals (with sharp bands at specific wavenumbers) or issues related to internal scattering of light. The volatile bands were then studied in detail. In contrast to the subtle features observed in R% spectra (Figs. 1a and 1b), the KK-Abs. volatile species bands (Figs. 1c and 1d) form positive peaks and the  $CO_3^{-2}$  species forms a doublet, like transmission IR spectra. Each band increases in KK-Abs. intensity with increasing volatile contents (Tables 2 and 3).

The baseline under the  $CO_{2 \text{ mol}}$  band is relatively linear, whereas the baselines under the  $CO_{3}^{-2}$  doublet and  $H_{2}O_{\text{mol}}$  bands

Sample	Method for		Volatile concentrations (wt%)			KK-Abs.values (±10% rel. unless noted)				
	H <sub>2</sub> O <sub>tot</sub>	H <sub>2</sub> O <sub>tot</sub>	H <sub>2</sub> O <sub>mol</sub>	CO <sub>2 mol</sub>	CO <sub>3</sub> <sup>2-</sup>	3600 cm <sup>-1</sup>	1620 cm <sup>-1</sup>	2350 cm <sup>-1</sup>	1520 cm <sup>-1</sup>	3600 cm <sup>-1</sup> (H <sub>2</sub> O <sub>tot</sub> )
		±10% rel.			-	(H <sub>2</sub> O <sub>tot</sub> )	(H <sub>2</sub> O <sub>mol</sub> )	(CO <sub>2 mol</sub> )	(CO <sub>3</sub> <sup>2-</sup> )	with MCT-A*
MHA23	SIMS	0.76	nd	0.077 (0.002)	0.337 (0.005)					
23-1						0.0035		0.0019	0.00074	
23-2						0.0046		0.0024	0.00100	
23-3						0.0041		0.0019	0.00076	
23y						0.0031		na		
MHA26	SIMS	2.43	0.75 (0.07)	0.033 (0.002)	0.373 (0.028)					
26-1						0.013	0.00130	0.0012	0.0005	0.0084 (0.0002)
26-2						0.014	0.00137	0.0013	0.0006	
26-3						0.013	0.00150	0.0011	0.0007	
26y						0.013	0.00160			
MHA27	SIMS	0.85	nd	0.077 (0.001)	0.283 (0.016)					
27-1						0.0038		0.00200	0.0007	0.0012
27-2						0.0032		0.00187	0.0008	0.0013 (0.0003)
27-3						0.0043		0.00204	0.0004	
MHA30	Manometry	2.48	1.09 (0.09)	0.004	0.076 (0.011)					
30-1						0.0140	0.00213	nd	0.0004	
30-2						0.0134	0.00180	nd	0.0002	
30-3						0.0144	0.00218	nd		
MHA31	Manometry	1.09	0.36 (0.03)	0.009 (0.001	0.061 (0.032)					
31-1						0.0076	0.00047	nd	0.0001	
31-2						0.0079	0.00050	0.00024	0.0002	
31-3						0.0073	0.00044	0.00044	0.0001	
MHA41	Manometry	2.69	1.34 (0.06)	0.014 (0.002)	0.323 (0.048)					
41-1	Manometry					0.019	0.00271		0.0004	
41-2	Manometry					0.020	0.00327		0.0005	
41-3						0.020	0.00300			
MHA44	Manometry	3.34	2.02 (0.19)	0.014 (0.002)	0.483 (0.040)					
44-1	Manometry					0.0240	0.0051		0.0012	0.0188 (0.0006)
44-2	Manometry					0.0230	0.0047	0.00037	0.0013	0.0156 (0.0002)
44-3						0.0245	0.0051	0.00034	0.0012	0.0175 (0.0003)
J12	SIMS	3.39	2.20 (0.28)	0.009 (0.001)	0.575 (0.054)					
J12-1	SIMS					0.0251		na	0.0012	0.0165 (0.0017)
J12-2	SIMS									0.0169 (0.0017)
J12-3										0.0175 (0.0018)
Note: Mea	surements we	re made M	CT-A in King's	lab unless note	d otherwise.					

TABLE 2. Results of KK-Abs. measurements on calc-alkaline andesite glasses in King's lab

are curved due to superposition on the large Si-O band. We fit the latter baselines using French curves following the techniques discussed in King et al. (2004). That work evaluated the error associated with different baseline correction approaches (e.g., fitting with Gaussian curves) and found that the French curve method allows for a conservative estimate of the fitting error (<10% relative). We assessed reproducibility by fitting maximum and minimum baselines to the spectra and the variations were all within this conservative error (<10% relative). The baseline under the 3600 cm<sup>-1</sup> peak was assumed to be linear between 3800 and ~2800 cm<sup>-1</sup>, and we assume that the line has a 10% error (e.g., King et al. 2002).

### Total H<sub>2</sub>O contents

The KK-Abs. for the  $H_2O_{tot}$  band measured from a linear baseline has good reproducibility, with:  $1\sigma < 5\%$  for 3–4 replicates on most samples and always <20% (e.g., Tables 2 and 3). For each composition, the  $H_2O_{tot}$  KK-Abs. increases linearly as a function of  $H_2O_{tot}$  wt% (Fig. 2), following the relation:

$$H_2O_{tot} wt\% = m \cdot (3550 \text{ cm}^{-1} \text{ band KK-Abs.})$$
(4)

where *m*, the slope of the calibration line, is a function of composition. The following are compositions/ $m/r^2$  of the fit: calc-alkaline andesite/154/0.93; alkali andesite/156/0.99; phonolite/172/0.90; and rhyolite/204/0.99. Our calibration slope for phonolite was presented in Mongrain et al. (2008) and our approach was used by Pauly et al. (2011), but the calibration

data were not shown in either case. We chose to force Equation 4 through zero because there should be no IR signal when  $H_2O_{tot}$  is absent.

The calibration slope for the basaltic glasses (m = 125,  $r^2 = 0.97$ ) is essentially defined by the sample with the high water content and so we suggest that this is a preliminary value. Also, the basaltic glasses with low water contents are all natural samples and may show more scatter because they have greater variation in composition and may have more intrinsic variation in water contents. Nonetheless, we are encouraged that Pauly et al. (2011) found a similar slope for palagonite (m = 123) when comparing KK-Abs. values to electron microprobe deficits.

#### Molecular H<sub>2</sub>O, molecular CO<sub>2</sub>, and carbonate contents

The  $H_2O_{mol}$  and  $CO_{2 mol}$  bands are relatively easy to identify in both the reflectance and KK-Abs. spectra because both form strong, narrow peaks (low degeneracy in the molecular vibration). In contrast, the  $CO_3^2$  doublet at these concentrations is not possible to identify in reflectance spectra (Fig. 1a; cf. Grzechnik et al. 1996; Moore et al. 2000) although the KK-Abs. spectra have a discernable  $CO_3^2$  doublet (Fig. 1c).

For the calc-alkaline andesite glasses we found the following relations:  $H_2O_{mol} = 435 \cdot (1620 \text{ cm}^{-1} \text{ KK-Abs.})$ ,  $r^2 = 0.92$ ,  $CO_{2 \text{ mol}} = 37.2 \cdot (2350 \text{ cm}^{-1} \text{ KK-Abs.})$ ,  $r^2 = 0.91$ ; and  $CO_3^{2-} = 500 \cdot (1520 \text{ cm}^{-1} \text{ KK-Abs.})$ ,  $r^2 = 0.74$ . The data used to constrain these equations are shown in Figures 3 to 5 and, again, we chose to force the fit through zero. The most difficult volatile to quantify with our technique is  $CO_3^{2-}$  based on the large scatter about the calibration

				KK-Abs.values (±10% r unless noted)	
Sample	Analysis no.	$H_2O_{tot}$	$H_2O_{tot}$ wt%	King's lab	Larsen's
		method	(±10% rel.	3600 cm <sup>-1</sup>	lab 3600
		1	unless noted)	) (H <sub>2</sub> O <sub>tot</sub> )	$cm^{-1}(H_2O_{tot})$
Alkali ande	sites				
MA512	MA512-1	Manometry	0 (0.01)	0	
	MA512-2	Manometry	0 (0.01)	0	
M12-1	M12-1-1	Manometry	2.62 (0.01)	0.0163	
	M12-1-2	Manometry	2.62 (0.01)	0.0161	
	M12-1-3	Manometry	2.62 (0.01)	0.0173	
M12-2	M12-2-2	Manometry	5.03 (0.01)	0.0298	
	M12-2-3	Manometry	5.03 (0.01)	0.0312	
	M12-2-4	Manometry	5.03 (0.01)	0.0323	
M12-4	M12-4-1	Manometry	6.76 (0.01)	0.0450	
	M12-4-2	Manometry	6.76 (0.01)	0.0453	
Basalts					
ManB_2249	Man2249-1	Trans IR	1.19	0.0103	
	Man2249-2	Trans IR	1.19	0.0101	
ManB15-4	Man15-4-1	Trans IR	1.32	0.0134	
	Man15-4-2	Trans IR	1.32	0.0129	
Gal1652	Gal1652-5a	Trans IR	1.38	0.0123	
ManB-18-3	Man18-3-1	Trans IR	1.44	0.0127	
	Man18-3-2	Trans IR	1.44	0.0126	
	Man18-11-1	Trans IR	1.49	0.0154	
	Man18-11-2	Trans IR	1.49	0.0119	
DL0414	DL0414_r1	Trans IR	4.63	0.0368	
	DL0414_r2	Trans IR	4.63	0.0357	
	DL0414_r4	Trans IR	4.63	0.0354	
Phonolites					
LPG1	LPG1	Trans IR	4.68 (0.50)	0.0273	
	LPG1-2	Trans IR	4.68 (0.50)	0.0262	
	LPG1-3	Trans IR	4.68 (0.50)	0.0267	
LPG1a	LPG1a_1	Trans IR	4.13 (0.23)	0.0222	
	LPG1a_2	Trans IR	4.13 (0.23)	0.0223	
	LPG1a_3	Trans IR	4.13 (0.23)	0.0205	
	LPG1a_4	Trans IR	4.13 (0.23)	0.023	
LPG3	LPG3-1	Trans IR	4.55 (0.08)		
	LPG3-1	Trans IR	4.55 (0.08)		
	LPG3-2	Trans IR	4.55 (0.08)		
	LPG3-3	Trans IR	4.55 (0.08)		
LHAE2	LHAE2-1	Trans IR	2.93 (0.15)	0.0181 (0.000	)7)
	LHAE2-2	Trans IR	2.93 (0.15)		
	LHAE2-3	Trans IR	2.93 (0.15)	0.0176 (0.000	)9)
	LHAE2-4	Trans IR	2.93 (0.15)	0.0170 (0.001	3)

TABLE 3. Results of KK-Abs. measurements on four glass compositions

TABLE 3. — CONTINUED

KK-Abs.values (±10% rel. unless noted) Sample Analysis no.  $H_2O_{tot}$ H<sub>2</sub>O<sub>tot</sub> wt% King's lab Larsen's method (±10% rel. 3600 cm<sup>-1</sup> lab 3600  $(H_2O_{tot})$   $cm^{-1}(H_2O_{tot})$ unless noted) Phonolites (cont.) LHAE3-1 4.99 (0.72) 0.0279 (0.0001) 0.0330 LHAE3 Trans IR LHAE3-2 Trans IR 4.99 (0.72) 0.0276 (0.0002) LHAE3-3 Trans IR 4.99 (0.72) 0.0282 (0.0003) LHAE4 LHAE4-1 0.0248 (0.0005) 0.0268 Trans IR 4.25 (0.25) LHAE4-2 Trans IR 0.0250 (0.0001) 4.25 (0.25) LHAE4-3 Trans IR 4.25 (0.25) 0.0249 (0.0003) 0.0290 (0.0003) 0.0312 LHAE5 LHAE5-1 Trans IR LHAE5-2 Trans IR 0.0287 (0.0001) 0.0329 LHAE5-3 0.0282 (0.0001) Trans IR LPG4 0.0260 (0.0001) 0.0302 LPG4-1 Trans IR 3.77 (0.05) LPG4-2 3.77 (0.05) 0.0277 (0.0004) Trans IR I PG4-3 Trans IR 3 77 (0.05) 0 0275 (0 0001) I PG6 LPG6-1 Trans IR 3.39 (0.10) 0.0177 (0.0005) 0.0251 LPG6-2 Trans IR 3.39 (0.10) 0.0168 (0.0007) 0.0270 0.0194 (0.0006) 0.0266 LPG6-3 Trans IR 3.39 (0.10) LPG6-4 Trans IR 3.39 (0.10) 0.0199 (0.0004) LSN1 LSN1-1 Trans IR 0.0294 (0.0005) 0.0430 I SN1-2 Trans IR 0.0302 (0.0001) 0.0396 I SN1-3 Trans IR 0.0299 (0.0006) LST3-40-1 LST3-40-1-1 Trans IR 0 (0.10) 0.0035 (0.0004) 0.0045 LST3-40-1-2 Trans IR 0 (0.10) 0.0030 (0.0001) LST3-40-1-3 Trans IR 0 (0.10) 0.0030 (0.0002) Rhvolites UTR2 0.09 (0.04) 0.0012 UTR2 Manometry PCD PCDA Manometry 0.13 (0.04) 0.0013 PCDB 0.17 (0.04) 0.0015 Manometry 2N 2N Manometry 2.07 (0.29) 0.0097 2N\_2 2.07 (0.29) 0.0103 Manometry Manometry 2.07 (0.29) 0.0089 2N 3 M3N M3N Manometry 2.94 (0.16) 0.0142 M3N 3 Manometry 2.94 (0.16) 0.0145 M3N 4 0.0141 Manometry 2.94 (0.16) M6N M6N Manometry 5.11 (0.38) 0.0239 0.0257 M6N 3 Manometry 5.11 (0.38) 0.0259 M6N 4 Manometry 5.11 (0.38)

line (Fig. 5), likely due to the nature of the  $CO_3^{2-}$  vibration and its resulting IR doublet band.

In sum, these results indicate that  $H_2O_{tot}$ ,  $H_2O_{mol}$ ,  $CO_{2 mol}$ , and possibly  $CO_3^{2-}$  may be quantified in aluminosilicate glasses using reflectance IR techniques. We anticipate that the technique may be further developed to quantify C-O-H species in a range of aluminosilicate compositions.

### DISCUSSION

#### Errors in regression analysis

To examine the goodness of fit (precision) of the calibration line with an ordinary least-squares fit, we determined how the calibration line fit varied when using a robust regression, employing the Isoplot program for Excel (Ludwig 2003). The purpose of using the robust regression is to examine whether outliers in the data might artificially influence the line fits and to also estimate how the slope of the calibration lines may vary accounting for the scatter in the data used for each composition. When forced through zero, the robust regression returns slopes that are within 7% of those determined using the ordinary least-squares fit (Table 4). The errors on the slope of the line determined using the robust regression method vary from 5 to 10% (Table 4). This is comparable with other methods, such as SIMS, where the data used for calibrations has  $\pm 10\%$  relative error. In contrast, errors on transmission FTIR and manometry data can be significantly smaller in the case of high-quality data where exceptional care is taken with the analyses. However, the error on transmission FTIR analyses is commonly significant in the case where thickness estimates are difficult, and thus the error on the analysis commonly approaches 10% or more (King et al. 2004).

To examine the predictive nature (accuracy) of the linear regressions for determining the compositions of unknown samples, we report the root mean square error prediction (RMSEP). This value provides a way to estimate the accuracy error involved in using the calibration line to predict the volatile content of an unknown sample. RMSEP values take into account the measurement errors, random errors, prediction errors (residuals), and bias. Values of RMSEP that are close to zero indicate smaller prediction errors for the linear regression. The formula for RMSEP is:

$$\text{RMSEP} = \sqrt{\frac{1}{n}} \sum_{i=1}^{n} \left( y_i - \hat{y}_i \right)^2 \tag{5}$$

where *n* is the number of measurements in the calibration set (including a sample with 0% of the volatile), *i* refers to a particular observation, and  $\hat{y}_i$  is the predicted value derived from a linear



**FIGURE 2.** KK-Abs. band height for the ~3600 cm<sup>-1</sup> band vs. total  $H_2O$  ( $H_2O_{tot}$ ) wt% determined via manometry or secondary ion mass spectrometry for a range of glass compositions: basalt, calc-alkaline andesite, alkali andesite, phonolite, and rhyolite.



**FIGURE 3.** KK-Abs. band height for the ~1620 cm<sup>-1</sup> band vs. molecular  $H_2O$  ( $H_2O_{mol}$ ) wt% determined via transmission IR spectroscopy for calc-alkaline andesitic glasses (King et al. 2002).

regression on the calibration set leaving out the measured value for that observation (i.e., a jack-knife or segmented approach). For H<sub>2</sub>O<sub>tot</sub>, the RMSEP values given in Table 4 have a range of 0.10 wt% (rhyolite) to 0.20 wt% (alkali andesite) for the linear fits. For the calc-alkaline glasses, very good prediction errors are observed for H<sub>2</sub>O<sub>mol</sub> (RMSEP=0.16 wt%; Fig. 3) and for CO<sub>2mol</sub> (RMSEP=0.003 wt%; Fig. 4) for linear fits. The CO<sub>3</sub><sup>-</sup> contents were so poorly correlated with the 1520 cm<sup>-1</sup> band height that we did consider it useful to calculate the RMSEP. The RMSEP are larger for the robust linear regressions (Table 4), as would be expected since those fits include errors due to outliers. The andesites, rhyolite, and basalt glasses show a small differences



**FIGURE 4.** KK-Abs. band height for the  $\sim 2350 \text{ cm}^{-1}$  band vs. molecular CO<sub>2</sub> (CO<sub>2 mol</sub>) wt% determined via transmission IR spectroscopy for cale-alkaline andesitic glasses (King et al. 2002).



**FIGURE 5.** KK-Abs. band height for ~1520 cm<sup>-1</sup> band vs.  $CO_3^{--}$  wt% determined via transmission IR spectroscopy for calc-alkaline andesitic glasses (King et al. 2002).

(up to 4 wt%) between the RMSEP for the standard linear regression vs. the robust linear regression, which indicates that the accuracy is for these glasses not greatly affected by outliers. The difference is greater for the phonolites (standard RMSEP = 0.19 wt% vs. robust RMSEP = 0.38 wt%), indicating that outliers do affect the phonolite calibration curves. In general, the RMSEP indicate that this method has accuracy errors for samples with 1 wt% H<sub>2</sub>O<sub>total</sub> that are less than 20% relative, with the relative accuracy improving at higher H<sub>2</sub>O<sub>tot</sub> contents.

# A general model for micro-reflectance IR measurements of total $H_2O$

A systematic increase in the calibration slope for  $H_2O_{tot}$ , *m*, with increasingly more mafic compositions is observed in the linear least-squares fit calibration lines (Fig. 2 and Table 4), as

	Basalt	Calc-alk.	Alkaline	Phonolite	Rhyolite
		andesite	andesite		
Molar Si+Al	0.66 (0.04)	0.76	0.78	0.73	0.85 (0.00)
Refr. index	1.590	1.544	1.533	1.506	1.491
	(0.013)				(0.005)
	Lea	st-squares fit	t results for H	2 <b>O</b> tot	
Slope	0.0080	0.0065	0.0064	0.0058	0.0049
m	125	154	156	172	204
RMSEP	0.12	0.14	0.20	0.19	0.10
	Robus	st regression	fit results for	H <sub>2</sub> O <sub>tot</sub>	
Slope	0.00795	0.00697	0.00627	0.00522	0.00480
	(+0.00080	(+0.00034	(+0.00044	(+0.00043	(+0.00018
	-0.00071)	-0.00150)	-0.00034)	-0.00033)	-0.00024)
т	126	143	159	192	208
	(+12/-12)	(+39/-7)	(+9/-10)	(+13/-15)	(+11/-8)
RMSEP	0.18	0.18	0.23	0.38	0.13

TABLE 4. Molar Si+Al, calculated refractive indices, fit parameters, and root mean square error of prediction for H<sub>2</sub>O<sub>tot</sub> calibrations for five different glass compositions measured in King's lab

well as in the robust regressions (Table 4). Such a systematic change suggests that there is a compositional or optical control on the calibration slope, similar to that observed for extinction coefficients determined by transmission IR spectroscopy (e.g., Dixon et al. 1995; Jakobsson 1997; Mandeville et al. 2002). We observe a reasonable negative correlation between (Si + Al) mole fraction and the calibration slope, m (Table 4). However, a better correlation is observed between the calculated refractive index, n, and m (Table 4 and Fig. 6) for both types of fits. We hypothesize that this correlation indicates that the optical properties of the glass are the major control on the slopes of the calibration lines. This finding indicates significant promise in developing a general reflectance IR calibration that covers a range of aluminosilicate glass compositions.

#### Analysis of glass wafers

Figure 7a shows the type of spectrum observed when a calcalkaline glass wafer (in this case 90 µm) is placed on a Au-coated glass slide and analyzed in reflection mode (referred to hereafter as "wafer reflectance"). Instead of a reflectance spectrum (Fig. 1a), the wafer reflectance spectrum (Fig. 7a) has strong similarities to a transmission spectrum (Fig. 7b) combined with a KK-Abs. spectrum (Fig. 1b). This effect is likely due to some of the light transmitting through the sample, being reflected by the Au-coated glass slide substrate, and returning to the detector in addition to some signal reflecting off the surface of the glass. Mixed transmission-reflectance behavior is supported by the observation that both the transmission and reflectance features are affected by thickness of the sample (Fig. 8). For example, in calc-alkaline andesite samples that are <~100 µm both transmission and reflectance features are observed, whereas those that are thicker show reflectance features only (like Fig. 1a). However, the thickness at which the behavior of the IR light through a wafer changes from mixed to reflectance is expected to depend on refractive index (a function of composition). Previous studies that only observed transmission features for rhyolites on a Au backing (Nowak and Behrens 1995) only examined the near-IR part of the spectrum where transmission features dominate and therefore the reflectance features were not observed.

In the wafer reflectance spectrum, the KK-Abs.intensity is amplified: relative to the transmission spectrum, the amount of



**FIGURE 6.** Slope of the calibration line (*m*) for total  $H_2O$  ( $H_2O_{tot}$ ) concentration vs. calculated refractive index, *n* for the range of compositions studied. Data are shown for both a linear least-squares regression approach (no error bars for *m*) and also a robust regression approach (with error bars for *m*). Note that a range of basaltic compositions were used in the calibration and only an average is reported.



FIGURE 7. (a) Reflectance spectrum for a wafer of representative calc-alkaline andesite glass (MHA-23) placed on a gold-coated glass slide. The numbers refer to the KK-Abs. of the band. (b) Transmission spectrum in absorbance units for a representative calc-alkaline andesite glass (MHA-23). Note that the KK-Abs. values are greater for the wafer in reflectance mode than the wafer in transmission mode.

absorbance amplification is a function of the wavenumber. For example, absorbance for  $H_2O_{tot}$  and  $CO_{2 mol}$  are amplified ~2×, whereas absorbance for  $CO_3^{-}$  is amplified ~1.5× (Figs. 7a and 7b). We predict that the absorbance amplification is non-linear due to variations in the optical constants of the glass at different frequencies (e.g. refractive index, dielectric constant etc.). Therefore, glasses of different compositions likely have specific "wafer reflectance calibration coefficients,"  $m_{WR, species}$ , for each



FIGURE 8. (a) Reflectance spectra of an initially 60  $\mu$ m thick calcalkaline andesite MHA-31 wafer that has been polished or roughened on its back to 1, 40, and 90  $\mu$ m . (b) Reflectance spectra of an initially 63  $\mu$ m thick calc-alkaline andesite MHA-44 wafer that has been polished to 1  $\mu$ m or roughened to 130  $\mu$ m on its back and set in epoxy. Note that with increased roughening the intensity decreases in the downward pointing transmission bands (T) related to C-O-H species and the intensity increases in the reflectance band (R) at ~1000 cm<sup>-1</sup> related to O-Si-O asymmetric stretching.

of the volatile species similar to transmission IR spectroscopy. Until we have better constrained  $m_{\text{WR, species}}$ , it would be advantageous to minimize the effect of transmission in a thin sample. For example, melt inclusions are commonly <100 µm and thin sections are ~30 µm; thus both types of samples provide the impetus for developing this micro-reflectance IR technique for thinner glass wafers.

To this end, we experimented with roughening the back of andesite wafers that were less than 100 µm thick. Figure 8a shows reflectance spectra for a calc-alkaline andesite glass wafer (MHA31, initially a 60 µm thick wafer) roughened on the back to 1, 40, and 90  $\mu$ m. The spectra are seen to become progressively less transmission-like and more reflectance-like with increasing roughness. Unfortunately, the wafer was damaged when it was roughened with 200 µm grit. However, Figure 8b shows a different andesitic glass wafer (MHA-44, initially a 63 µm thick wafer) roughened with a ~130 µm polishing grit and set in epoxy resulted in an almost "pure" reflectance spectrum. The reflectance spectrum from the roughened sample was converted to KK-Abs. units, resulting in a spectrum that is very similar to a KK-Abs. spectrum derived from a reflectance spectrum collected from a thick glass (Fig. 9). The total water content calculated using these two KK-Abs. spectra was identical, with both spectra giving 4 wt% H<sub>2</sub>O<sub>tot</sub>. This value is within error of the measured H<sub>2</sub>O<sub>tot</sub> content of 3.34 wt%.

In sum, glass wafers may be analyzed with reflectance techniques using samples where the back has been roughened to ~130  $\mu$ m and set in epoxy. However, glass wafer methods deserve further development because spectra collected on such samples have the advantage of higher absorbance signal (Fig. 7a) and would be useful for analysis of melt inclusion and thin sections.

# Interlaboratory comparison, comparison of IR detectors, and caveats

To determine if our reflectance IR technique depends on the instrument setup we undertook two separate studies. First, measurements were made using different detector types (MCT-A\* vs. MCT-A in King's lab); and second, we measured the phonolite glasses at the University of Alaska (Larsen's lab). The MCT-A detector has a lower detectability rating (D\* =  $4.5 \times 10^{10}$ ) than the MCT-A\* (D\* =  $6.5 \times 10^{10}$ ) according to the manufacturer (ThermoFisher). Larsen's lab uses a Nicolet 6700 FTIR (compared to King's Nicolet Nexus 670), but both used a Globar source, XT-KBr beamsplitter, and a Continuµm microscope with a liquid nitrogen cooled MCT-A detector.

Figure 10a shows that the KK-Abs. value for  $H_2O_{tot}$  using the MCT-A\* detector is consistently lower than that determined using the MCT-A detector (Fig. 10a). This result is surprising because the MCT-A\* detector is designed with higher detectability, thus we cannot rule out detector alignment issues. Nonetheless, the data show that such variables (detector, alignment etc.) are important for accurate calibration curves.

Figure 10b shows the results of phonolite analyses in Larsen's lab using the spectrometer configuration noted above. The reflectance peak heights from Larsen's lab are systematically higher than those obtained in King's lab using an older instrument, although some results are within error (Fig. 10b).

Together, the results of analyses with different machine configurations indicate that volatile calibration curves depend on a particular instrument's configuration (e.g., light throughput and detector efficiency). Therefore, to use these techniques it is necessary to create instrument-specific calibration curves and to check these periodically. This is similar to other micro-analytical techniques currently used for volatiles that also depend on instrument



**FIGURE 9.** KK-Abs. spectra for two samples of calc-alkaline andesite MHA-44 wafer: (1) solid spectrum is from a wafer (initially 63  $\mu$ m thick) roughened to 130  $\mu$ m on its back and set in epoxy; and (2) dashed spectrum is from a thick glass that was singly polished.

setup and efficiency, such as secondary ion mass spectrometry (SIMS) and Raman spectroscopy. One of the advantages of the IR method relative to SIMS is that "blanks" as well as calibration standards do not need to be measured at each analytical session. Furthermore, we have shown that the IR calibrations for  $H_2O_{tot}$  are related to calculated refractive indices, which means that once calibrations have been done in a laboratory using certain glass compositions that it is possible to estimate the slope of the calibration line, *m*, to determine  $H_2O_{tot}$  in other compositions.

As implied above, the KK transform results are sensitive to the input parameters, especially the wavelength range. Smoothing to window spacings that are slightly different results in negligible changes to the calculated KK-Abs. spectrum. However, we find that spectra that are measured over different wavelength ranges do not necessarily give the same KK-Abs. as our technique measured over 5300–650 cm<sup>-1</sup> due to variations in extrapolating the wings of the reflectance spectrum prior to KK transformation. Despite the potential *accuracy* errors associated



FIGURE 10. (a) KK-Abs. band height for the  $\sim$ 3600 cm<sup>-1</sup> band vs. total H<sub>2</sub>O (H<sub>2</sub>O<sub>tot</sub>) wt% in calc-alkaline andesites measured using either a MCT-A or MCT-A\* detector. (b) KK-Abs. band height for the  $\sim$ 3600 cm<sup>-1</sup> band vs. total H<sub>2</sub>O (H<sub>2</sub>O<sub>tot</sub>) wt% in phonolites measured in King's lab and Larsen's lab.

with the wavelength extrapolation, this study shows that two laboratories or different detectors produce results with similar *precision* errors. Furthermore, the technique produces slopes for the calibration lines for a range of glass compositions that may be logically explained through a physical property of the glass: the refractive index. We recommend using the same spectral range as our measurements when using our approach.

# Application of the micro-reflectance IR techniques to challenging samples

We have been able to determine volatile contents in fragile experiments on bubble nucleation (Mongrain et al. 2008) and H<sub>2</sub>O<sub>tot</sub> diffusion from wet basaltic to dry rhyolitic melt (glass) (Lui 2005). We have monitored H<sub>2</sub>O<sub>tot</sub> uptake during weathering of basaltic glasses that cannot be polished (our unpublished data) and others have measured H2Otot contents of fragile palagonites (Pauly et al. 2011). Also, we have measured the H<sub>2</sub>O<sub>tot</sub> content of fragile impact melt glasses (Harris et al. 2007). The impact glasses had in excess of 20 wt% H2Otot determined by alternate methods (Harris et al. 2007) and would have been exceedingly difficult to prepare as very thin wafers for micro-transmission IR analyses; thus, the reflectance method is advantageous for unusually H-rich samples. We recommend that the µ-R-IR technique is best for samples with high volatile contents. In such samples, the method allows for rapid determination of the H2Otot and H<sub>2</sub>O<sub>mol</sub> contents.

Our findings are consistent with technique development using attenuated total reflection (ATR; Lowenstern and Pitcher 2013) and synchrotron reflectance IR techniques. Combined with these techniques, IR analysis with a mapping stage with overlapping spectral collection (several micrometer stepsize) should improve the areal resolution of micro-reflectance IR analyses and potentially allow smaller volumes to be sampled than with transmission IR methods, which are commonly limited by the thickness of the sample. Spectra can be obtained on glassy areas and checked for the presence of minerals by verifying that the 700–1300 cm<sup>-1</sup> region only contains a broad glass band.

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